Some Observations from the M3D CDX-U Sawtooth Study

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# **Motivation**

- As a nonlinear code benchmark, model sawtooth events in a tokamak plasma with extended MHD models using realistic physical values to make quantitative predictions.
  - Large tokamaks have large disparities in spatial and temporal scales to be resolved.
    - Resistive MHD: Current sheet thickness ~  $S^{-1/2}$
    - Two-fluid MHD: ion skin depth ~  $n^{-1/2}$
  - Small tokamaks operate in regimes accessible to presentday codes.

# <u>Characteristics of the Current Drive</u> <u>Experiment Upgrade (CDX-U)</u>



- Low aspect ratio tokamak  $(R_0/a = 1.4 1.5)$
- Small ( $R_0 = 33.5 \text{ cm}$ )
- Elongation  $\kappa \sim 1.6$
- $B_T \sim 2300$  gauss
- $I_p \sim 70 \text{ kA}$
- $n_e^P \sim 4 \times 10^{13} \text{ cm}^{-3}$
- $T_e \sim 100 \text{ eV} \rightarrow \text{S} \sim 10^4$
- Discharge time ~ 12 ms
- Soft X-ray signals from typical discharges indicate two predominant types of low-*n* MHD activity:
  - sawteeth
  - "snakes"



10

0.2

 $\psi - \psi_0$ )

- Nonlinear evolution
  - disruption?
  - stagnation?

0.

• repeated reconnections?

# **Baseline Parameters for CDX**

Lundquist Number S	$\sim 2 \times 10^4$ on axis.
Resistivity <i>η</i>	Spitzer profile $\propto T_{eq}^{-3/2}$ , cut off at 100× $\eta_0$
Prandtl Number Pr	10 on axis.
Viscosity <del>µ</del>	Constant in space and time.
Perpendicular thermal conduction $\kappa_{\perp}$	200 m <sup>2</sup> /s (measured value at edge)
Parallel thermal conduction (sound wave)	$V_{\rm Te} = 6 V_{\rm A}$
Peak Plasma <mark>β</mark>	~ $3 \times 10^{-2}$ (low-beta).
Density Evolution	Turned on for nonlinear phase.
Nonlinear initialization	Pure <i>n</i> =1 perturbation such that $\frac{\max(B_{pol}^1)}{\max(B_{\phi}^0)} = 10^{-4}$

# <u>n=1 Eigenmode</u>

### Incompressible velocity stream function U



Toroidal current density  $J_{\phi}$ 

 $\gamma \tau_{\rm A} = 8.61 \times 10^{-3} \rightarrow \text{growth time} = 116 \tau_{\rm A}$ 

# Higher n Eigenmodes

Incompressible velocity stream function U

*n* = 3 *n* = 2 *n* = 4

 $\begin{array}{c} m \geq 5 \\ \gamma \, \tau_{\rm A} = 1.28 \times 10^{-2} \end{array}$ 

 $\begin{array}{c} m \geq 7 \\ \gamma \, \tau_{\rm A} = 1.71 \times 10^{-2} \end{array}$ 

 $\begin{array}{c} m \approx 9 \\ \gamma \tau_{\rm A} = 1.87 \times 10^{-2} \end{array}$ 

#### In Absence of Heat Conduction,

#### Higher n Modes are More Unstable than Internal Kink



### Resistivity Scaling is Consistent with low-*n* Resistive Ballooning Modes



For each toroidal mode number *n*, the linear growth rate  $\gamma$  is found to be proportional to  $\eta^{\alpha}$ :

n	α
2	0.597
3	0.590
4	0.568
5	0.553
6	0.543
7	0.542
8	0.546
9	0.560
10	0.586

#### Parallel Heat Conduction Alone Stabilizes Some Modes, But Does Not Appear to Cause Saturation of Unstable Modes



#### Moderate Isotropic Heat Conduction Has Stronger Stabilizing



Kinetic Energy

#### High Perpendicular Heat Conduction Stabilizes All Ballooning



### Heat Conduction in M3D

Isotropic component in resistive MHD energy equation

$$\frac{\partial p}{\partial t} + \mathbf{v} \cdot \nabla p = -\gamma p \nabla \cdot \mathbf{v} + \rho \nabla \cdot \kappa_{\perp} \nabla \left( \frac{p}{\rho} - \frac{p_0}{\rho_0} \right)$$
(1)

where  $p_0/\rho_0$  is the equilibrium temperature.

Artificial sound wave model for  $\kappa_{\parallel}$ :

$$\frac{\partial T}{\partial t} = s \frac{\mathbf{B} \cdot \nabla u}{\rho} \tag{2a}$$

$$\frac{\partial u}{\partial t} = s\mathbf{B} \cdot \nabla T + v \nabla^2 u \qquad (2b)$$

where *s* is the electron sound speed.

## Parallel Transport

Solving (2) only, holding other quantities fixed, with  $s = 6 v_A$ . Start from equilibrium temperature distribution, with field as shown for t = 1795.61 below.



# Parallel + Perpendicular Transport

• Parallel heat conduction alone results in very fast equilibration, but cannot rapidly smooth out all bumps where field is stochastic.

• From (1), the time scale for perpendicular equilibration is approximately  $\lambda^2/4\pi^2\kappa_{\perp}$ , where  $\lambda$  is the wavelength of the temperature fluctuation being smoothed. Adding some  $\kappa_{\perp}$  and solving (1) as well as (2) can quickly smooth the smallest bumps:

$$\kappa_{\perp} = 10^{-5} (2.2 \text{ m}^2/\text{s}) \rightarrow \tau/\tau_A \sim 2500 \ \lambda^2$$
:



continue, setting  $\kappa_{\perp} = 10^{-4} (22 \text{ m}^2/\text{s}) \rightarrow \tau/\tau_A \sim 250 \ \lambda^2$ :



### Nonlinear Sawtooth History



### **Total Energy and Core Temperature**



## Initial state: *t* = 1266.17

Poincaré plot 1.0 0.5 0.0 -0.5-1.0

1.5

2.0

0.5

1.0





# <u>Island growing: *t* = 1660.70</u>

Poincaré plot 1.0 0.5 0.0 -0.5-1.00.5 1.0 1.5 2.0







# Nonlinear phase: *t* = 1795.61







# <u>After 1st Crash: *t* = 1839.86</u>

Poincaré plot







### Flux surfaces recovered: t = 2094.08

#### Poincaré plot







### Nonlinear Sawtooth History

22 Modes Retained





# During 1st Crash: *t* = 1717.08



# <u>After 1st Crash: *t* = 1725.34</u>



## <u>After Incomplete Recovery: *t* = 1848.30</u>





- Both timing and energy of peaks are different.
- Outer flux surfaces do not heal in highest-resolution case.
- Energy in higher-*n* modes significantly affects sawtooth evolution.
- Further study is needed to assess convergence on this case.

### Toroidal Structure of the Sawtooth Current Sheet



Isosurface of n>0 part of toroidal current density during sawtooth crash shows that the current peak occurs where the 1,1 island is reconnecting. Following this peak around the torus indicates that its variation in  $\phi$  may not be fully resolved in this calculation.

### Toroidal Structure of the Sawtooth Current Sheet,

<u>continued</u>



# **Conclusions**

- Nonlinear MHD simulation with actual device parameters is capable of tracking evolution through repeated sawtooth reconnection cycles.
- The coupling of reasonable  $\kappa_{\parallel}$  with finite  $\kappa_{\perp}$  will rapidly smooth out any temperature fluctuations resulting from instabilities strong enough to render the local magnetic field stochastic. This should cause any pressure-gradient-driven modes to saturate.
- The increased energy, faster cycle, and differing toroidal structure in the current sheet indicate that the problem is not fully resolved at 48 planes. More resolution or an extended MHD model may be necessary to achieve convergence.