#### Code Coupling: Pedestal-ELM Cycle



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#### THE CPES Code Coupling Team

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#### **Research participants**

- Gunyoung Park, NYU
  - XGC0 kinetic code development
- Julian Cummings, Caltech
  - QA Testing, Convergence tests
- Scott Klasky and Roselyne Barreto, ORNL
  - Data visualization and Dashboard development
- Norbert Podhorszki, UC Davis
  - Kepler workflow development and testing
- Linda Sugiyama, MIT; Hank Strauss, NYU
  - M3D code development and expertise



- Phil Snyder, General Atomics
  - ELITE code development and expertise
- Alexey Pankin, Lehigh University
  - XGC NIMROD Coupling
  - C.S. Chang, NYU: Principal Investigator
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## Talk outline

- Motivations for kinetic-MHD code coupling
- Description of code coupling scenario and operation
- Discussion of Kepler workflow automation activities
- Data movement and visualizations for sample run
- Conclusions and future work





# Kinetic <-> MHD coupling in a nutshell

- Overarching goal is to model edge pedestal buildup (kinetic model) followed by ELM crash (MHD model)
- As kinetic code proceeds, we must determine when the pedestal pressure and current density profiles become MHD unstable
- Then switch control to MHD model for nonlinear evolution of ELM and "healing" of MHD equilibrium
- Ideally, we then restart/rerun kinetic code based upon the new equilibrium and continue the cycle
- Process of alternating control should be automated





# Coupling Fusion codes for Full ELM, multi-cycles

- Run XGC until unstable conditions
- M3D coupling data from XGC
  - Transfer to end-to-end system
  - Execute M3D: compute new equilibrium
  - Transfer back the new equilibrium to XGC
  - Execute ELITE: compute growth rate, test linear stability
  - Execute M3D-MPP: to study unstable states (ELM crash)
  - Restart XGC with new equilibrium from M3D-MPP







#### Basic code coupling scenario

- Four different simulation codes in use
  - <u>XGC0</u>: kinetic simulation of edge plasma, including neoclassical transport with ion-electron-neutral dynamics
  - <u>M3D-OMP</u>: MHD analysis code, produces equilibrium & mesh
  - <u>ELITE</u>: ideal MHD linear stability analysis code
  - <u>M3D-MPP</u>: parallel two-fluid nonlinear MHD initial value code
  - DEGAS-2 : atomic physics -not discussed in this talk
- MHD codes accept plasma profile data (n,T,j) from kinetic code, use eqdsk file for magnetic equilibrium
- M3D-OMP generates hi-res equilibrium for ELITE linear analysis and mesh for M3D-MPP nonlinear simulation
  - EFIT eq. 65x65, ELITE eq. 140x200, M3D-MPP mesh 40x200
- If ELM is linearly unstable, M3D-MPP run is launched





## Details of code coupling scenario

- XGC0 code run on Cray XT3, initialized with *geqdsk*
- XGC0 periodically writes profile data *m3d.in* and *peqdsk* – *Plasma profile, equilibrium data transferred to Infiniband cluster*
- M3D-OMP reads *m3d.in* and eqdsk data, performs equilibrium solve, and writes high-res *eqdsk* file
- ELITE does linear stability check on *eqdsk* and *peqdsk* 
  - Linear growth rate of sample mode compared to  $\omega_{*_{pi}}/2$  threshold
  - Kepler enables simultaneous runs to obtain converged growth rates for several toroidal mode numbers across the spectrum
- If stable, return updated eqdsk data file to XGC0
- If unstable, stop XGC0 and do nonlinear M3D-MPP run
  - Requires one-step M3D-OMP run to produce M3D meshfile





#### User's view of scientific workflow

- User launches XGC simulation on production computing platform (e.g., *jaguar*) as usual via PBS script
- User then initiates workflow via shell script (or GUI)
- Workflow monitors simulation output, requests postprocessing & visualization, and directs any code coupling
  - Migrate binpack \*.bp files and NetCDF \*.cdf files to ewok
  - Run bp2h5 utility to convert binpack files to portable HDF5
  - HDF5 files can be input to AVS service for data visualization
  - NetCDF time history data files "split" to recover newest time slices since last monitoring update, then migrated and "fused"
  - NetCDF data is imaged with ncgraph utility or can be monitored
  - All data and images archived into HPSS at simulation end





## Kepler workflow basics

- Open-source, actor-oriented framework using Java
  - Actors encapsulate parameterized actions
  - Communicate via input and output data ports
- Computations performed by flow of control
  - No call-and-return semantics
  - Tokens passed from one actor to the next trigger actions
- Several key advantages for scientific workflows
  - Supports parallel-pipeline processing (sync or async)
  - Not restricted to web service or computing grid applications
  - Open-source development of customized actors





## **CPES** Coupling workflow



# Kepler graphical workflow (cont'd)



Archive ELITE logs and files







#### Running the Kepler scientific workflow

- Kepler is Java-based framework and uses the JRE
  - User must load jre module on ewok for workflow use
- Create "workflow" dir in lustre disk area on ewok
  - Contains all transferred and post-processed simulation files
- Launch XGC0 job on jaguar with PBS script as usual
  - XGC0 uses input parameters to indicate MHD coupling
- Run shell script from /ccs/proj/fus011/Workflows area
  - Script command: start-workflow.sh <shot#> <jobid>
  - Script logs in to specific ewok node to process files and launch M3D codes as needed (user passcode required)
  - Separate scripts exist for just code monitoring or for coupling
- Then, go to ewok (or Dashboard!) and view results





## Sample run of code coupling workflow

- Basic XGC0 simulation of edge pedestal buildup
- Input data drawn largely from DIII-D shot analysis
  - Initial equilibrium from EFIT analysis of shot 096333
  - Model profiles for initial plasma density and temperature
- XGC0 run parameters
  - 320,000 ion particles run on 128 cores of Cray XT4
  - Set to run for 100 ion toroidal transit periods in 50,000 steps
  - Simple neutron physics model with 0.985 recycling rate
  - One dump of plasma profiles and eqdsk update in each ion toroidal transit time
- Equivalent uncoupled run takes ~75 minutes on jaguar
- Full ELM cycle simulation takes 2-3 hours wall clock time on the combined jaguar-ewok system





## Edge pedestal buildup



## Data movement: XGC0 and M3D-OMP

- XGC0 writes *m3d\_<ts>.in* files with (n,T,j) profile data
  - Here <ts> represents the timestep or dump number
- Kepler actor looks for most recent dump file
  - Migrated from Cray XT3 to Infiniband cluster, renamed m3d.in
- Kepler runs serial M3D-OMP code to update eqdsk
  - Kepler is provided with M3D-OMP input file for "eq only" run
  - Initial run starts from same EFIT equilibrium data as XGC0
  - Subsequent runs update from the previous eqdsk file
- Kepler actor transfers updated eqdsk back to XGC0
  - XGC0 rereads equilibrium data in each coupling phase
- Output from each M3D-OMP run saved in subdirectory





#### Data movement and visualization: ELITE

- ELITE run for linear stability check has several inputs
  - High-resolution equilibrium data copied from M3D-OMP run
  - Plasma density data in peqdsk format transferred from XGC0
  - ELITE input file controls grid resolution, modes of interest, etc.
- After each M3D-OMP run, Kepler gathers inputs, runs serial ELITE code, and monitors and graphs results
  - ELITE run has 3 phases: eq solve, vacuum bc, eigenvalue
  - Kepler scans results of each phase, logs possible problems
    - Example: ELITE results may indicate eigenmode not resolved
  - Growth rate is collected, compared with stability threshold
  - Kepler plots ELITE output variables using IDL script
- Output from each ELITE run saved in subdirectory





#### ELITE output: normalized pressure gradient



#### ELITE output: normalized plasma current



#### Data Movement and Visualization: M3D-MPP

- If ELITE returns *unstable*, Kepler launches M3D-MPP
  - Prepare: run one step with M3D-OMP, write meshfile omp.out
  - Run: submit 128-process batch job on cluster with omp.out and input parameters file for nonlinear ELM simulation
  - Typical run has 16 poloidal planes with toroidal period = 3
- M3D-MPP produces periodic output in HDF5 format
- AVS/Express module used to plot output variables
  - Visualization routines made available as a service
  - Kepler launches service to listen for plotting requests
  - Simple Python script extracts data and requests plots
  - Resulting images organized within shot output directory
- Resulting data and images are archived using HPSS





# M3D-MPP: plasma density



Time given in units of the Alfvén period





# **ORNL Dashboard tool**

- Monitor computational resources and simulations within a standard web browser
  - Secure login using one-time-password token
  - View job queues on various ORNL and NERSC machines
  - Access current and archived runs, plus those of collaborators
  - Each run offers list of output variables for plotting
- Graphical display utilizes Flash animation
  - Plots update on-the-fly as new data is discovered
  - View of dashboard can be recorded for later playback
  - Flash capabilities include display of 3d visualizations
- Provenance information





#### Monitoring XGC output on Dashboard







#### Monitoring M3D-MPP output on Dashboard







#### Next steps for code coupling workflow

- Improvements for M3D-MPP
  - Automatic detection of quasi steady-state (n,T,j) profiles
  - Transfer of "healed" equilibrium back to XGC0 code
- Attempt multiple ELM cycles by restarting XGC0
- Extend workflow designs to accommodate alternate MHD equilibrium and stability analysis codes
  - TEQ module in XGC0 along with ELITE stability check and nonlinear ELM simulation with NIMROD
- Enhance use of simulation metadata within workflow
  - Provenance tracking for a simulation database
  - Improved interaction with the ORNL Dashboard tool
- Verification of coupling results
  - Convergence tests have been planned



