## Two-Fluid Benchmarking of the 1/1 Internal Kink

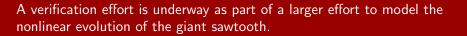
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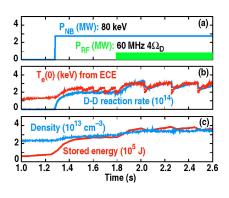




- Motivation
- Review of existing theory
- 3 Benchmarking MHD and two-fluid calculations
- Summary

## Sawteeth are periodic relaxation events of the core plasma.

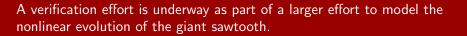
- In tokamaks sawteeth result from the nonlinear evolution of a n = 1 mode.
- The sawtooth cycle is characterized by a slow build-up of the core n and T<sub>e</sub> followed by a rapid crash.
- Two fluid drifts and kinetic effects temporarily stabilize the kink leading to larger but less frequent giant sawteeth.
- Giant sawteeth are a concern for modern tokamaks and ITER.



[Choi et al., PoP 14, 2007]

# Modeling of the giant sawtooth requires an accurate representation of multiple two fluid effects.

- Different two fluid effects modify the stability of the internal kink in opposing ways.
  - Diamagnetic drifts reduce the linear growth rate when the diamagnetic frequency is comparable to the MHD growth rate [Ara et al., 1978].
  - Finite electron compressibility allows the electrons and ions to decouple and increases the linear growth rate in the semi-collisional and collisionless regimes [Zakharov and Rogers, Phys Fluids B. 4, 1992].
  - Electron inertia increases the growth rate in in collisionless regime.
- ullet We present the results of a verification effort to test NIMROD's ability to capture the different two fluid effects for the 1/1 kink in a screw pinch equilibrium.
  - NIMROD accurately captures the transition from ideal to resistive kink in resistive MHD.
  - NIMROD accurately captures the increased growth rate in the semi-collisional regime in the absence of diamagnetic drifts.
  - There is a significant discrepancy between NIMROD and the analytic theory of Zakharov and Rogers when drifts are included.



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# The ideal MHD 1/1 kink is characterized by a "top hat" radial displacement.

• The Euler-Lagrange equation for a screw pinch is:

$$\begin{split} \frac{\delta W}{2\pi} &= \frac{\pi}{\mu_0} \int \left( f \left( \frac{d\xi}{dr} \right)^2 + g\xi^2 \right) dr \\ f &\sim r B_\theta^2 \left( 1 - q \right)^2 \\ g &\sim \frac{B_\theta^2}{r} \left( (1 - q)^2 - 2 \left( 1 + q \right) (1 - q) \right) + 2 \frac{r^2}{R^2} \frac{d\mu_0 p}{dr} \end{split}$$

- The quantity g is negative for |q| < 1.
- $\bullet$   $\delta W$  is negative indicating instability for the top hat trial function:

$$\xi = \begin{cases} \xi_{\infty} & \text{for} \quad r < r_{s} \\ 0 & \text{for} \quad r > r_{s} \end{cases}$$

The dynamics in a thin layer around the discontinuity at the rational surface determines the linear growth rate.

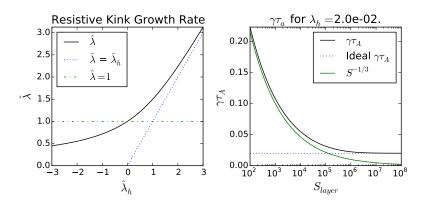
 Coppi et al. worked out the resistive MHD kink dispersion relation in the limit of a thin layer [Sov J. Plasma Phys. 2, 1976]

#### Resistive MHD kink dispersion relation

$$\hat{\lambda} = \hat{\lambda}_h \left( \frac{\hat{\lambda}^{9/4}}{8} \frac{\Gamma\left(\frac{\hat{\lambda}^{3/2} - 1}{4}\right)}{\Gamma\left(\frac{\hat{\lambda}^{3/2} + 5}{4}\right)} \right)$$

- $\hat{\lambda} = \gamma \tau_{\Delta} S^{1/3}$  and  $\hat{\lambda}_{b} = \lambda_{b} S^{1/3}$
- $\lambda_h=rac{-\pi}{B_o^2q'^2r_s^2}\int_0^{r_s}gdr$  is the normalized ideal MHD growth rate.
- The resistive kink growth rate is  $\gamma_r \tau_A = S^{-1/3}$  ( $\lambda_h = 0$ ).
- The Alfven time and Lundquist number are defined relative to the layer:  $\tau_A^2 = \frac{\mu_0 \rho_0}{R^2} \frac{R^2}{q'^2 r_s^2}, \ \tau_R = \mu_0 r_s^2 / \eta, \ \text{and} \ S = \tau_R / \tau_A.$

The resistive MHD dispersion relation captures the transition from tearing mode to ideal kink.



- The ideal MHD growth rate  $\hat{\lambda} = \hat{\lambda}_h$  is recovered in the limit of large  $\hat{\lambda}_h$  (large S).
- The resistive kink growth rate  $\hat{\lambda}=1$  is recover in the limit of  $\hat{\lambda}_h\approx 0$  (small S).
- Tearing behavior is recovered when  $\hat{\lambda}_h \ll 0$ .

# Two-fluid modifications to the internal kink are described by Zakharov and Rogers inner layer equation [Phys. Fluids B 4, 1992].

The two-fluid kink growth rate,  $\Gamma$ , is an eigenvalue of the inner layer equation

$$\left[\lambda_{s}^{2}+\frac{\lambda_{e}^{2}\Gamma\left(\Gamma-i\Omega_{*i}\right)}{q'^{2}x^{2}}Z'\right]'=\left(1+\frac{\Gamma\left(\Gamma-i\Omega_{*i}\right)}{q'^{2}x^{2}}\right)Z-\frac{2L_{h}}{\pi x^{2}}\int_{0}^{\infty}Zdx$$

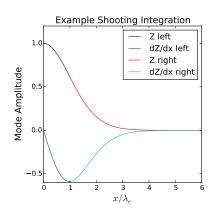
that satisfy the boundary condition

$$\lim_{|x|\to\infty}Z=\frac{2L_h}{\pi x^2}\int_0^\infty Zdx.$$

- $Z=ia\gamma\xi_a'\simeq V_{\theta}$  is approximately the poloidal flow.
- $\Gamma = \gamma \tau_A$  and  $\Omega_{*i} = \omega_{*i} \tau_A$  are the normalized growth rate and ion diamagnetic frequency.
- $\lambda_s^2 = \frac{\rho_s^2 \Gamma(\Gamma i\Omega_{*i})}{\left(\Gamma i\frac{5}{3}\Omega_{*e}^n\right)\left(\Gamma i\frac{5}{3}\Omega_{*i}^n\right)}$  is a modified ion sound gyroradius length scale squared.
- $\lambda_e^2 = \frac{\Gamma}{\Gamma i\Omega_{ma}} \left( \frac{1}{S\Gamma} + d_e^2 \right)$  is the effective resistive skin depth squared.
- $\bullet$   $L_h$  is the inertial ideal MHD length scale.

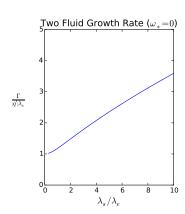
The eigenvalues of the inner layer equation are calculated numerically using a shooting method.

- The layer equation is solved by treating it as an initial value problem.
- The equation is integrated twice:
  - The first integration starts at x = 0 and integrates outwards towards an intermediate value of x<sub>m</sub>.
  - The second integration starts at  $x_r > \max(\lambda_e, \lambda_s)$  and integrates inwards towards  $x_m$ .
  - To seed the integration we initially guess the  $\int Zdx$ .
- We converge on  $\Gamma$  by minimizing the error in Z and Z' at  $x_m$  and the the guess of  $\int Z dx$ .

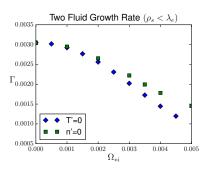


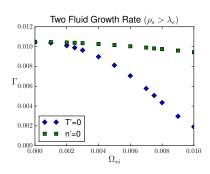
## Decoupling of the electrons and ions leads to faster growth rates at large $\rho_s$ .

- The linear growth rate is enhanced when the ion sound length scale is larger than the resistive skin depth  $(\lambda_s > \lambda_e)$ .
  - Here the two-fluid treatment is accurate provided that the ion-sound Larmor radius is larger than the ion Larmor radius  $(\rho_s > \rho_i)$ .
  - The current sheet width is characterized by the resistive skin depth λ<sub>e</sub>.
  - The flow "sheet" is characterized by the ion sound length scale  $\lambda_s$ .

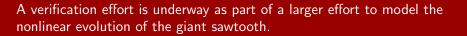


### Diamagnetic drifts decrease the linear growth rate at fixed $\rho_s$ .





- In the collisional limit  $(\lambda_e > \rho_s)$  drifts that arise due to temperature gradients have a similar impact of the growth rate as drifts that arise due to density gradients.
- In the semi-collisional limit ( $\lambda_e < \rho_s$ ) drifts that arise due density gradients have a larger impact of the growth rate than drifts due to temperature gradients.



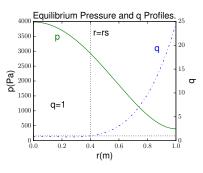
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Linear calculations use the NIMROD code to evolve the primitive fields.

$$\begin{split} &\rho\left(\partial_{t}\vec{V}+\vec{V}\cdot\nabla\vec{V}\right)=\vec{J}\times\vec{B}-\nabla P-\nabla\cdot\pi_{i}\\ &\pi_{i}=-\rho\nu_{iso}W+\frac{P_{i}}{4\Omega_{ci}}\left[\hat{b}\times W\cdot\left(I+3\hat{b}\hat{b}\right)-\left(I+3\hat{b}\hat{b}\right)\cdot W\times\hat{b}\right]\\ &W=\nabla\vec{V}+\nabla\vec{V}^{T}-2/3I\nabla\cdot\vec{V}\\ &\partial_{t}n+\nabla\cdot\left(n\vec{V}\right)=\nabla\cdot\left(D\nabla n-D_{h}\nabla\nabla^{2}n\right)\\ &n\left(\partial_{t}T_{s}+\vec{V_{s}}\cdot\nabla T_{s}\right)=-\left(\gamma-1\right)P_{s}\nabla\cdot\vec{V_{s}}-\left(\gamma-1\right)\nabla\cdot\vec{q}_{s}\\ &\partial_{t}\vec{B}=-\nabla\times\left[\eta\vec{J}-\vec{V}\times\vec{B}+\frac{1}{ne}\left(\vec{J}\times\vec{B}-T_{e}\nabla n\right)+\mu_{0}d_{e}^{2}\partial_{t}\vec{J}\right]+k_{divb}\nabla\nabla\cdot\vec{B} \end{split}$$

- Artificial particle diffusivity and magnetic divergence diffusions are used to provide numerical stability.
- Gyro-viscosity and two-fluid corrections to Ohm's law are included in two-fluid calculations.

## Calculations are performed in a screw pinch that is n = 1 ideal kink unstable.

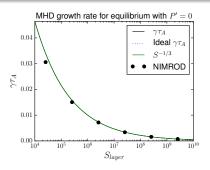


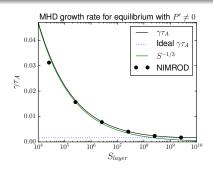
Equilibrium Parameters:

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$q_0$	0.9
q(a)	24.5
$B_0$	1T
R/a	30
$\beta_0$	1%

- Equilibria are generated by specifying the pressure and safety factor.
- Two equilibria are studied: one with a uniform pressure and the other with a spatially varying pressure profile.
  - This allows for the study of two fluid modifications to the kink with and without diamagnetic drift stabilization.
- A strongly sheared q profile is needed to produce thin layers.

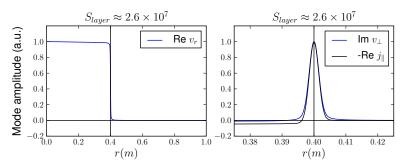
Resistive MHD calculations capture the transition from resistive to ideal kink.





- The resistive interchange scaling  $\gamma \tau_A = S^{-1/3}$  is an excellent approximation for the uniform pressure equilibrium with  $S < 10^{10}$ .
  - Here the ideal drive is weak:  $\gamma_{\it ideal} \, au_{\it A} = 4.2 imes 10^{-5}$
- The pressure gradient is the dominant source of free energy for the nonuniform pressure equilibrium.
  - Ideal behavior is recovered for  $S \gtrsim 10^9$ .
  - Here the ideal growth rate is  $\gamma_{ideal}\tau_A=1.6\times 10^{-3}$ .
- The validity of the analytic theory breaks down due to a finite layer width at small S.

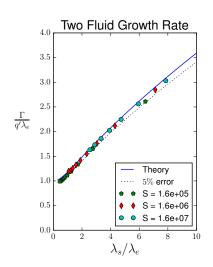
## The radial velocity resembles the "top hot" trial function at large S.



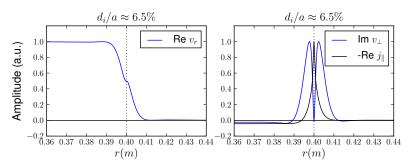
- Figures show the mode structure for the nonuniform pressure equilibrium at  $S=2.6\times 10^7$ .
- The radial flow is proportional to the displacement.
- As predicted from MHD theory, the momentum and current layers have the same width.
  - Note that the horizontal axes use different scales in the two plots.

# NIMROD accurately calculates the growth rate in the transition from the collisional to the semi-collisional regime in the absences of drifts.

- The linear growth rates agree with theory to within 5% error for a wide range or parameters.
  - This agreement has been verified for  $\rho_s/\lambda_e \lesssim 60$  and  $S \lesssim 10^9$ .
- These calculations use the uniform pressure equilibrium.
- The theoretical growth rate is calculated assuming  $\lambda_h = 0$ .

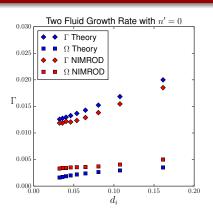


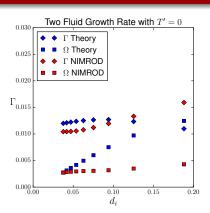
#### The separation of layer widths is observed at large $d_i$ .



- The width of the flow layer scales with  $\rho_s$ .
- The current layer width depends on both the resistive layer width and the electron skin depth.
  - In the collisionless limit the current layer width scales linearly with electron skin depth.
- Figures show the mode structure for  $S = 2.6 \times 10^7$ .

## There is considerable disagreement in the calculated growth rate when drift effects are included.





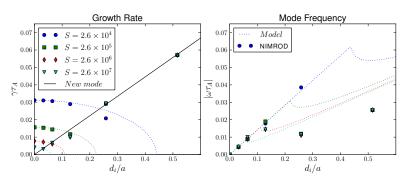
- Reasonable agreement is observed between NIMROD and the theory in the uniform density calculation, where pressure gradients are due to temperature gradients.
  - Here the drifts have a small effect on the growth rate.
- There is considerable disagreement in the uniform temperature calculation.
- Increasing  $d_i$  increases both  $\rho_s$  and  $\omega_*$  in these calculations.

## NIMROD reproduces the two-fluid linear kink behavior in several regimes.

- Resistive MHD calculations correctly calculate the linear growth rate in both the inertial and resistive regimes.
  - Agreement between theory and calculations is limited by the validity of small layer approximation.
- The two-fluid calculations agree with the theory of Zakharov and Rogers to within 5% error in the absence of drifts.
  - Agreement is obtained for a wide range of parameters  $\rho_s$  and S.
- The agreement breaks down when in calculations where the diamagnetic drifts have a significant effect on the growth rate.
  - Further work is needed to understand the source of the disagreement.

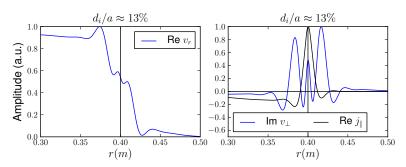
# Extra Slides

Drifts reduce the linear growth rate at small  $d_i$  for all cases with a finite pressure gradient.



- The calculated growth rates are approximated by a simple model:  $\omega = \omega_{*i} + i \sqrt{\gamma_{MHD}^2 \omega_{*i}^2/4}$ .
- Here the nonuniform pressure equilibrium is used.
- The 1/1 kink in not the dominant mode at large  $d_i$ .
  - The new mode is characterized by a large  $v_{\parallel}$ .
  - This modes scales linearly with  $d_i$  and is insensitive to S.

Large oscillations in the mode structure are observed when the drifts have a significant impact on the growth rate.



- These oscillations are characteristic of drift stabilization.
- Figures show the mode structure for  $S = 2.6 \times 10^5$ .
- Here the growth rate is 25% smaller than the MHD growth rate.