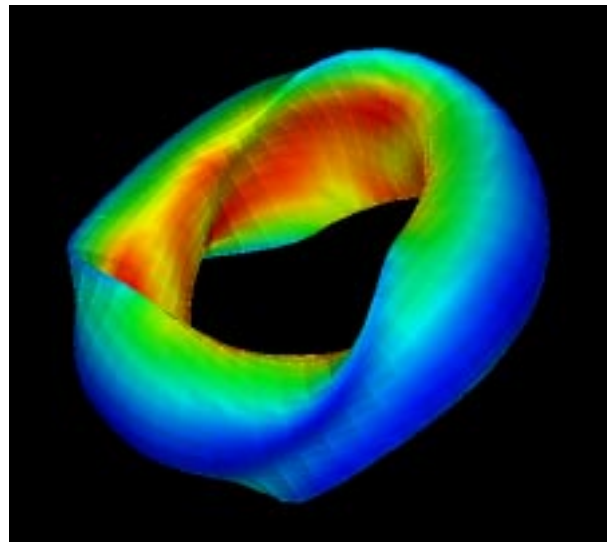


**Calculation  
of the Neoclassical Radial Electric Field  
using the Global Toroidal Code**

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## Motivation

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- Neoclassical radial electric field is important for stellarator transport (e.g. collisionless particle dynamics; zonal flow physics; transport barriers)
  - Departure of axi-symmetry can be weak (i.e. QA concept): standard calculation of  $E_r$  can be difficult
  - Well-established gyro-kinetic particle simulation techniques offer alternative possibility to determine  $E_r$
  - Method can be generalized to viscous flow damping in fully three-dimensional, non-axisymmetric geometries

## The Method

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Write confining  $\mathbf{B}$  field in Boozer coordinates  $(\psi, \theta, \zeta)$  as

$$\begin{aligned}\mathbf{B} &= \iota(\psi) \nabla\zeta \times \nabla\psi + \nabla\psi \times \nabla\theta \\ \mathbf{B} &= g(\psi) \nabla\zeta + I(\psi) \nabla\theta + \beta_* \nabla\psi\end{aligned}\quad (1)$$

with  $g(\psi) \propto$  poloidal current;  $I(\psi) \propto$  toroidal current.

Jacobian of transformation  $\mathcal{J} \equiv [\nabla\psi \cdot (\nabla\theta \times \nabla\zeta)]^{-1}$  satisfies

$$\mathcal{J}B^2 = g(\psi) + \iota(\psi) I(\psi) \equiv f(\psi) \quad (\text{flux surface quantity}) \quad (2)$$

Ion Momentum Balance

$$\rho \frac{d\mathbf{V}}{dt} = -\nabla \cdot \mathbf{P} + en \left( \mathbf{E} + \frac{\mathbf{V} \times \mathbf{B}}{c} \right) + \mathbf{F} + \mathbf{R} \quad (3)$$

where  $\mathbf{V}$  : fluid velocity;  $\rho$  : mass density,  $\mathbf{R}$  : collisional drag;  $\mathbf{F}$  : (external) applied force and  $\mathbf{P} = P_{\parallel} \widehat{\mathbf{b}}\widehat{\mathbf{b}} + (\mathbf{I} - \widehat{\mathbf{b}}\widehat{\mathbf{b}}) P_{\perp}$  : the pressure tensor; also  $P_{\parallel}$  ( $P_{\perp}$ ) : parallel (perpendicular) pressure.

Take scalar product of Eq.(2) with  $\mathbf{e}_{\zeta} \equiv \partial\mathbf{r}/\partial\zeta$  (with  $\mathbf{r}$  is the position vector) and operate with  $\langle \dots \rangle = \iint \dots \mathcal{J}(\psi, \theta, \zeta) d\theta d\zeta$  one obtains

$$\frac{\iota(\psi)}{c} \frac{dQ}{dt} - \left\langle \frac{dL_{\zeta}}{dt} \right\rangle = \left\langle \frac{\partial \widehat{P}}{\partial \zeta} \right\rangle - T_{\zeta} \quad (4)$$

where  $\widehat{P} \equiv (P_{\parallel} + P_{\perp})/2$ ,  $T_{\zeta} = \langle (\mathbf{R} + \mathbf{F}) \cdot \mathbf{e}_{\zeta} \rangle$  is the torque due to applied forces and collisional drag;  $L_{\zeta}$  is the toroidal component of the canonical momentum  $\mathbf{L} = \rho \mathbf{V} + e \mathbf{A}/c$  where  $\mathbf{A} = \psi \nabla \theta - \chi \nabla \zeta$  is the vector potential and  $2\pi\chi$  is the poloidal flux.

To derive Eq.(4), note that

$$\left\langle e n \mathbf{e}_{\zeta} \cdot \left( \frac{\mathbf{V} \times \mathbf{B}}{c} \right) \right\rangle = \frac{e}{c} \iint n \mathbf{V} \cdot (\mathbf{B} \times \mathbf{e}_{\zeta}) \mathcal{J} d\theta d\zeta = \frac{e}{c} \iota(\psi) \int \mathbf{\Gamma} \cdot d\boldsymbol{\sigma}_n = \frac{\iota(\psi)}{c} \frac{dQ}{dt}$$

Here  $Q$  is the total charge,  $\mathbf{\Gamma} = n \mathbf{V}$  is the particle flux, and  $d\boldsymbol{\sigma}_n \equiv \mathcal{J} \nabla \psi d\theta d\zeta$  is an area element normal to the magnetic surface  $\psi = \text{const}$  and pointing outwards.

For zero applied force and after a few ion-ion collision times, toroidal balance equation reads

$$\frac{\iota(\psi)}{c} \frac{dQ}{dt} = S \equiv \left\langle \frac{\partial \widehat{P}}{\partial \zeta} \right\rangle \quad (5)$$

Knowing the parallel and perpendicular pressures on the magnetic surface (velocity moments of  $\delta f$ ), one obtains a measure of the radial particle flux on that surface through Eq.(5)

## Calculation of Parallel & Perpendicular Pressures PPPL

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Write perpendicular pressure as

$$P_{\perp} = \sum_{m,n} (P_{\perp})_{m,n} \exp [i (m\theta + nN_p\zeta)] \quad (6)$$

where  $N_p$  is the number of field periods of the configuration and the Fourier coefficients are calculated according to

$$(P_{\perp})_{m,n} = \frac{\int_0^{2\pi} d\theta \int_0^{2\pi} d\zeta (mv_{\perp}^2/2) \delta f \exp [-i (m\theta + nN_p\zeta)] d^3v}{\int_0^{2\pi} d\theta \int_0^{2\pi} d\zeta} \quad (7)$$

Guiding center motion and collisions will spread the particles toward equal density in pitch and over the magnetic surface

Then, in a small layer  $\delta\psi \ll \psi_b$  (boundary), one notes that

$$\int \int d\theta d\zeta \implies \int \mathcal{J}^{-1} (\delta\psi)^{-1} d^3x \implies [F (\bar{\psi}) \delta\psi]^{-1} \int B^2 d^3x$$

and the  $(P_{\perp})_{m,n}$  Fourier components become

$$(P_{\perp})_{m,n} = \int \frac{\int d^3x (mv_{\perp}^2/2) \delta f B^2 \exp [-i (m\theta + nN_p\zeta)]}{\int d^3x B^2} d^3v \quad (8)$$

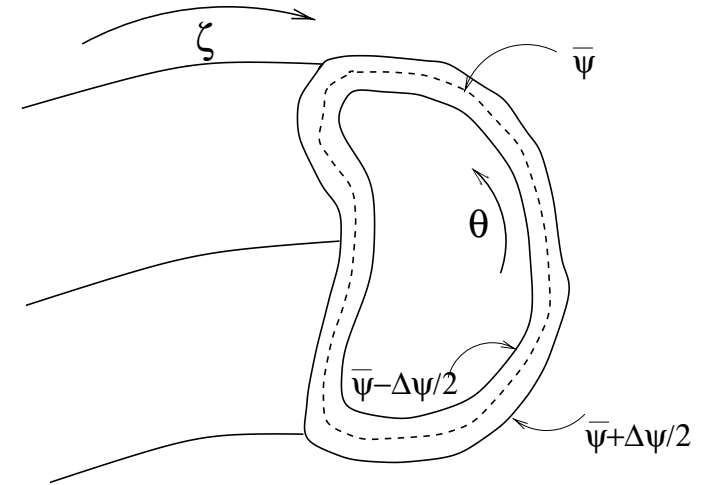
Same method applies for  $P_{\parallel}$ .

# Numerical Method

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Particles are initially randomized in  $\theta$  and  $\zeta$ , and between  $\bar{\psi} - \Delta\psi/2$  and  $\bar{\psi} + \Delta\psi/2$ .

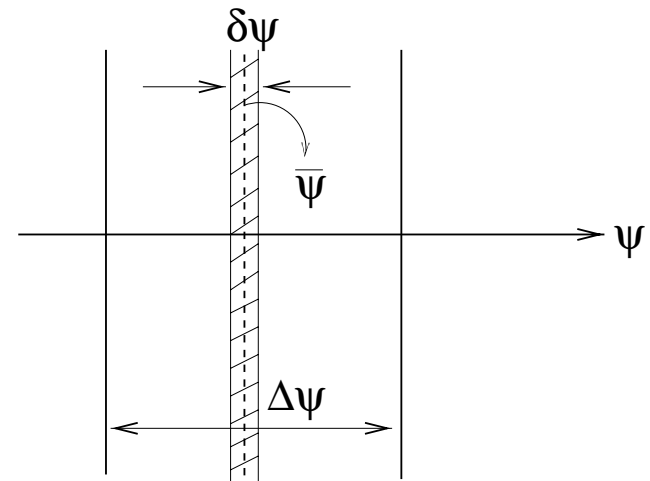
Introduce radial coordinate  $r = \sqrt{\psi/B_0}$  so that  $\delta r \approx \Delta\psi / (2\sqrt{B_0\psi})$ ; typical drift time is  $\tau_d \approx \Delta r / V_d$  where  $V_d \approx V_{th} (\rho_{th}/R_c)$  is the typical radial curvature drift velocity.



We must have  $\tau_d \gg \tau_r$ , where  $\tau_r$  is the relaxation time (typically a few ion-ion collision times).

Calculation of  $P_{||}$  and  $P_{\perp}$  are carried out within an annulus  $\delta\psi \ll \Delta\psi \ll \psi_b$  centered around  $\bar{\psi}$ .

Parallel and perpendicular pressures calculated on different processor element (PE) are collected on a single PE (PE=0), on which the Fourier components  $(P_{\perp})_{m,n}$  and  $(P_{||})_{m,n}$  are evaluated.



## Divergence of Pressure Tensor

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Write Pressure Tensor  $\mathbf{P}$  as

$$\mathbf{P} = \tilde{P}\mathbf{B}\mathbf{B} + P_{\perp}\mathbf{I} + P_{\perp}\mathbf{I} \quad (9)$$

where  $\tilde{P} \equiv (P_{\parallel} - P_{\perp}) / B^2$ . Noting that

$$\nabla B^2 / 2 = \mathbf{B} \times (\nabla \times \mathbf{B}) + (\mathbf{B} \cdot \nabla) \mathbf{B} \quad (10)$$

and

$$\nabla \cdot \mathbf{P} = \mathbf{B} (\mathbf{B} \cdot \nabla \tilde{P}) + \tilde{P} \nabla \cdot (\mathbf{B}\mathbf{B}) + \nabla P_{\perp} \quad (11)$$

and using Ampere's law and the radial force balance equation, we obtain

$$\nabla \cdot \mathbf{P} = \mathbf{B} (\mathbf{B} \cdot \nabla \tilde{P}) + \tilde{P} \left( \frac{1}{2} \nabla B^2 + 4\pi \nabla P_0 \right) + \nabla P_{\perp} \quad (12)$$

where  $P_0 = P_0(\psi)$  is the equilibrium pressure. Taking the scalar product of Eq.(12) with  $\mathbf{e}_{\varphi} \equiv \partial \mathbf{r} / \partial \varphi$  where  $\mathbf{r}$  is the position vector and  $\varphi = \{\theta, \zeta\}$  one gets

$$\mathbf{e}_{\varphi} \cdot (\nabla \cdot \mathbf{P}) = B_{\varphi} (\mathbf{B} \cdot \nabla \tilde{P}) + \frac{\tilde{P}}{2} \frac{\partial}{\partial \varphi} B^2 + \frac{\partial P_{\perp}}{\partial \varphi} \quad (13)$$

Taking the flux-surface average  $\langle \bullet \rangle \equiv \iint \mathcal{J} (\bullet) d\theta d\zeta$  of Eq.(13) yields

$$\langle \mathbf{e}_\varphi \cdot (\nabla \cdot \mathbf{P}) \rangle = \frac{1}{2} \left\langle \frac{\partial}{\partial \varphi} (P_{\parallel} + P_{\perp}) \right\rangle , \quad (14)$$

since  $\langle \mathbf{B} \cdot \nabla F \rangle = 0$  for any function  $F = F(\psi, \theta, \zeta)$  and

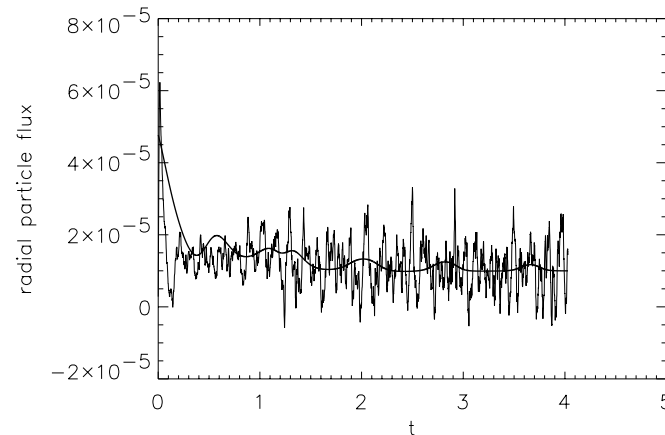
$$\begin{aligned} \left\langle \tilde{P} \frac{\partial B^2}{\partial \varphi} \right\rangle &= \left\langle \frac{P_{\parallel} - P_{\perp}}{B^2} \frac{\partial B^2}{\partial \varphi} \right\rangle \\ &= \left\langle B^2 (P_{\perp} - P_{\parallel}) \frac{\partial}{\partial \varphi} \left( \frac{1}{B^2} \right) \right\rangle \\ &= \left\langle \frac{f(\psi)}{\mathcal{J}} (P_{\perp} - P_{\parallel}) \frac{\partial}{\partial \varphi} \left( \frac{1}{B^2} \right) \right\rangle \\ &= f(\psi) \iint (P_{\perp} - P_{\parallel}) \frac{\partial}{\partial \varphi} \left( \frac{1}{B^2} \right) d\theta d\zeta \\ &= f(\psi) \iint \frac{1}{B^2} \frac{\partial}{\partial \varphi} (P_{\parallel} - P_{\perp}) d\theta d\zeta \\ &= \iint \mathcal{J} \frac{\partial}{\partial \varphi} (P_{\parallel} - P_{\perp}) d\theta d\zeta \\ &= \left\langle \frac{\partial}{\partial \varphi} (P_{\parallel} - P_{\perp}) \right\rangle . \end{aligned}$$



## Results & Conclusions

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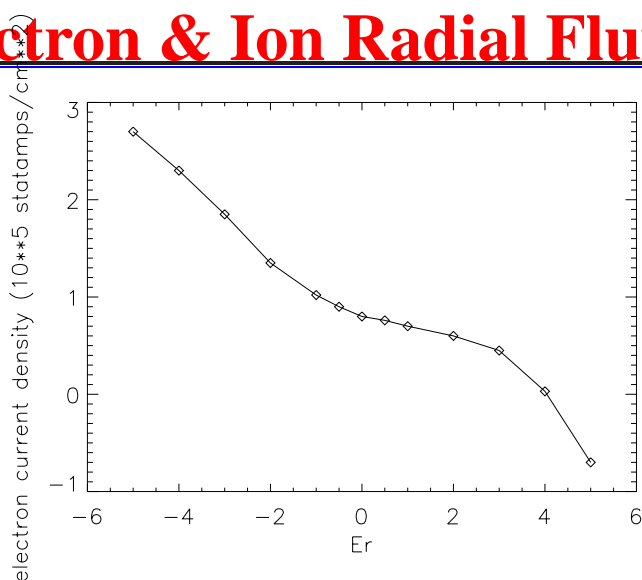
- Run for NCSX plasma (C82 configuration); with central ion temperature  $T_i(0) = 2.76$  KeV; central electron temperature  $T_e(0) = 2.14$  KeV; central plasma density  $n_0 = 6.73 \times 10^{13}$  cm $^{-3}$ . Magnetic surface of reference  $\psi/\psi_b = 0.7$ .
- Equilibrium  $B$  field is specified using 30 Fourier harmonics
- Trajectories of  $2 \times 10^5$  Lagrangian markers are integrated; time step  $\Delta t/\tau_{ii} = 4 \times 10^{-4}$ ; collisional effects are calculated every 10 time steps.
- Background distribution function  $f_0$  loaded as a Maxwellian with  $\langle V_{\parallel} \rangle = 0$ .



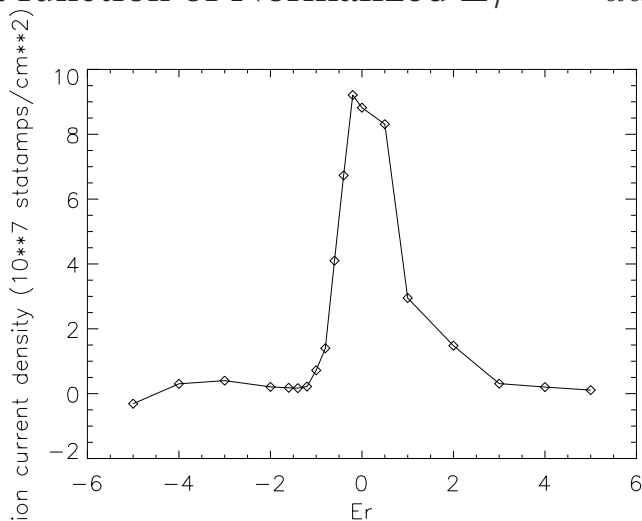
Direct measurement (broken line) and gyro-kinetic calculation (smooth curve) of  $\Gamma_r$  (a.u.)

# Electron & Ion Radial Fluxes

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Electron Current Density as a function of Normalized  $E_r = -ad/dr (e\Phi/T_i(0))$



Ion Current Density as a function of Normalized  $E_r = -ad/dr (e\Phi/T_i(0))$

Stable root found at  $E_r \simeq -26.2$  kV/m