

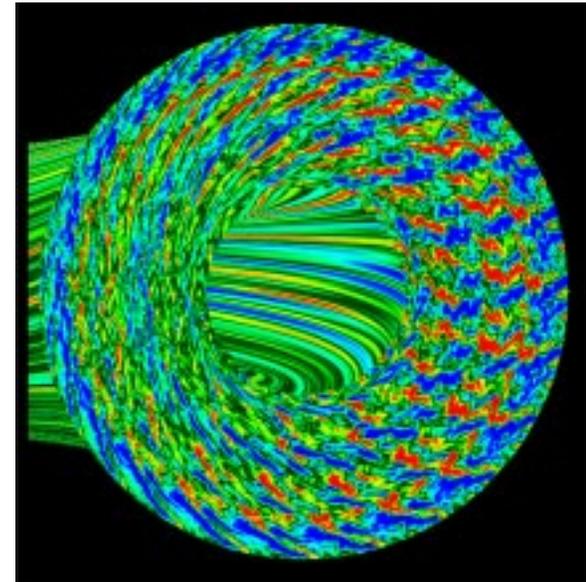
# **THEORY-BASED MODELS OF TURBULENCE AND ANOMALOUS TRANSPORT IN FUSION PLASMAS**

**G.W. Hammett, Princeton Plasma Physics Lab  
w3.pppl.gov/~hammett  
APS Centennial, Atlanta, March, 1999**

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**Close collaboration with M.A. Beer (PPPL),  
W. Dorland (Univ. of Maryland), M. Kotschenreuther  
(Univ. of Texas), R.E. Waltz (General Atomics).**

**Acknowledgments: A. Dimits, G.D. Kerbel (LLNL),  
T.S. Hahm, Z. Lin, P.B. Snyder (PPPL),  
S.E. Parker (U. Colorado), and many others.**



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part of the Numerical Tokamak Turbulence Project national collaboration, a DOE HPCCI Grand Challenge.**

# **THEORY-BASED MODELS OF TURBULENCE AND ANOMALOUS TRANSPORT IN FUSION PLASMAS**

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## **I. Simple picture of plasma microinstabilities**

**Inverted pendulum → Rayleigh-Taylor → Magnetic curvature instability.**

**Difference between MHD and micro-instabilities/drift-waves.**

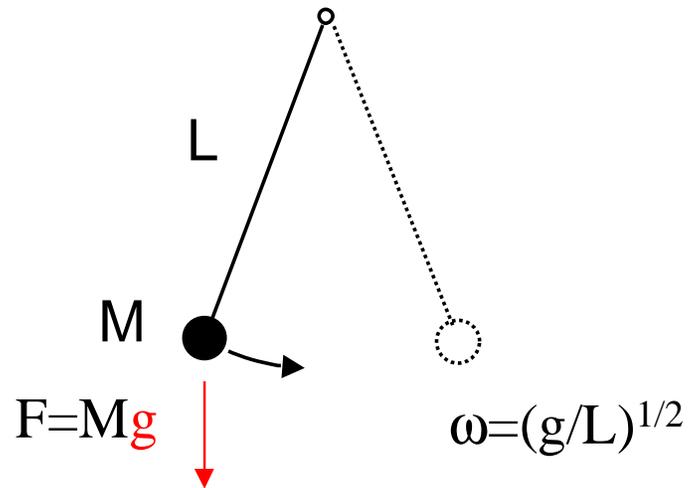
## **II. Complexity and challenge of plasma turbulence**

**nonlinear, chaotic, wide-range of space and time scales**

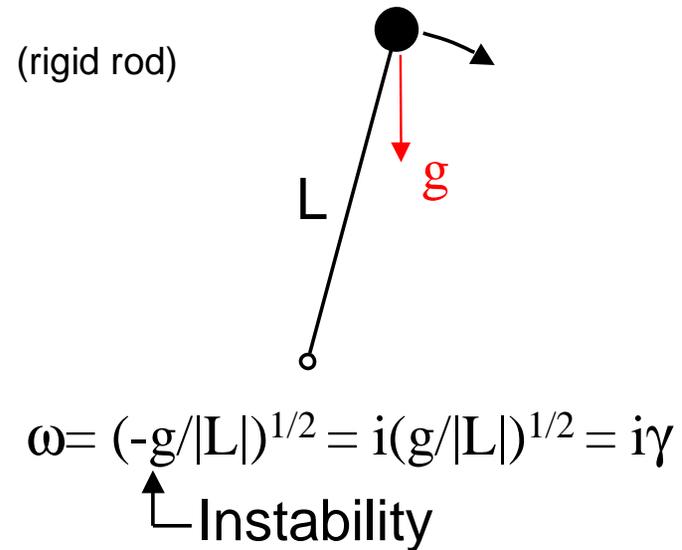
**theoretical and computational advances made in tackling these problems.**

## **III. Comparisons with experiments, remaining challenges.**

## Stable Pendulum

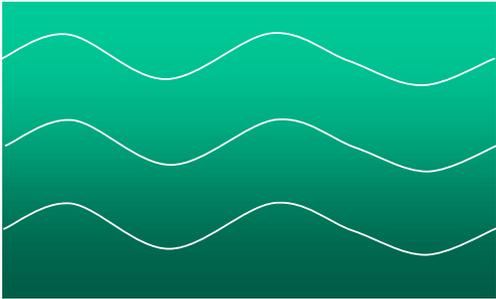


## Unstable Inverted Pendulum



## Density-stratified Fluid

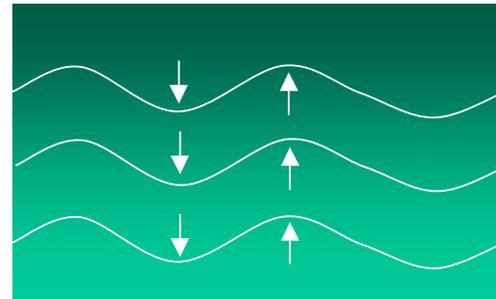
$$\rho = \exp(-y/L)$$



stable  $\omega=(g/L)^{1/2}$

## Inverted-density fluid ⇒ Rayleigh-Taylor Instability

$$\rho = \exp(y/L)$$

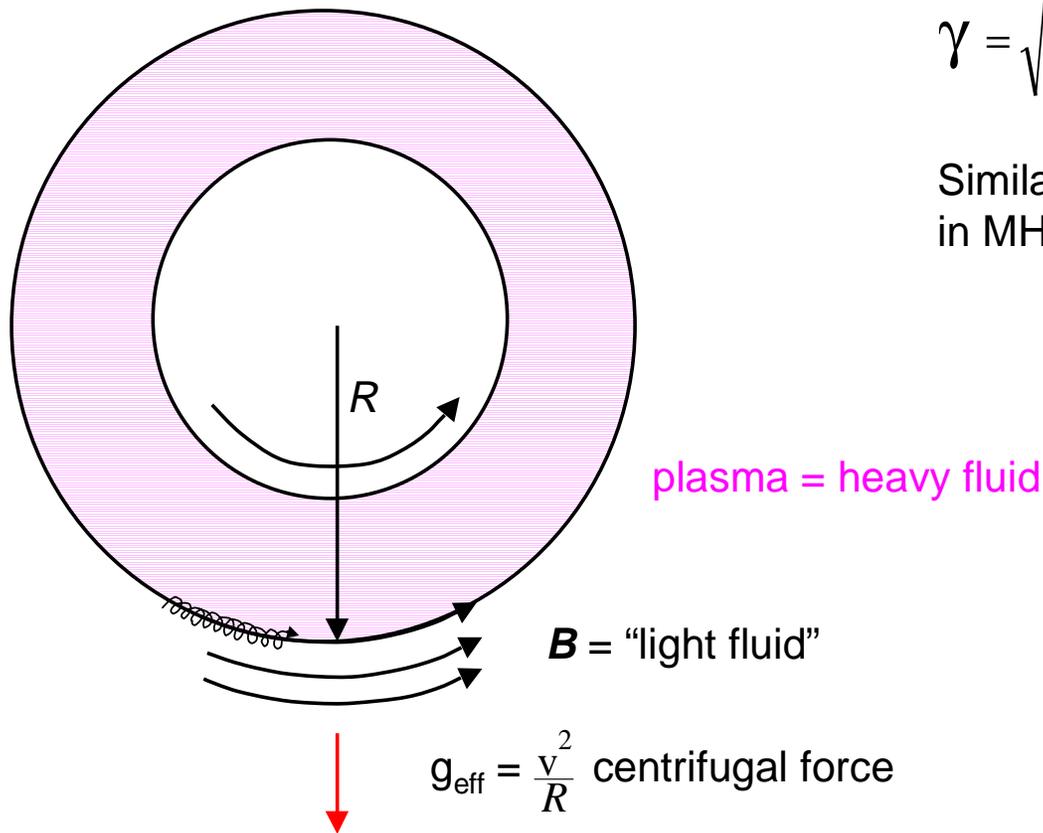


Max growth rate  $\gamma=(g/L)^{1/2}$

# “Bad Curvature” instability in plasmas ≈ Inverted Pendulum / Rayleigh-Taylor Instability

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Top view of toroidal plasma:



Growth rate:

$$\gamma = \sqrt{\frac{g_{\text{eff}}}{L}} = \sqrt{\frac{v_t^2}{RL}} = \frac{v_t}{\sqrt{RL}}$$

Similar instability mechanism  
in MHD & drift/microinstabilities

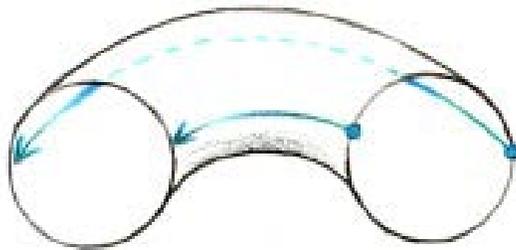
$1/L = \nabla \rho / \rho$  in MHD,  
 $\propto$  combination of  $\nabla n$  &  $\nabla T$   
in microinstabilities.

# The Secret for Stabilizing Bad-Curvature Instabilities

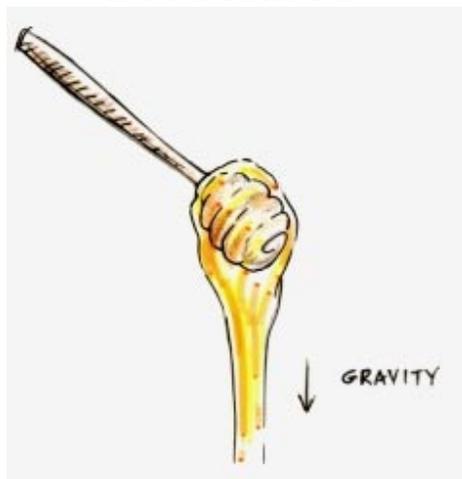
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Twist in  $\mathbf{B}$  carries plasma from bad curvature region to good curvature region:

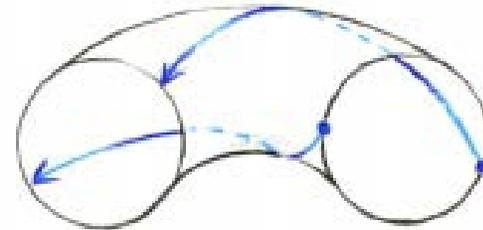
PURELY TOROIDAL  $\mathbf{B}$



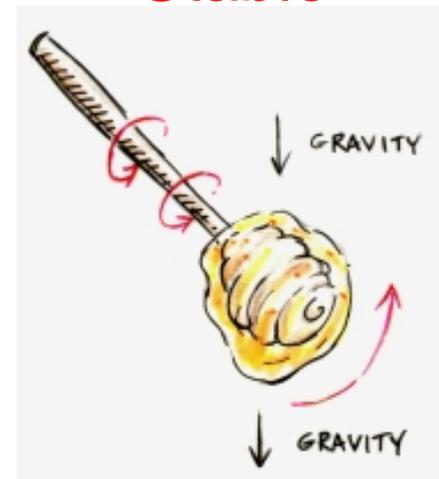
Unstable



TWISTING  $\mathbf{B}$



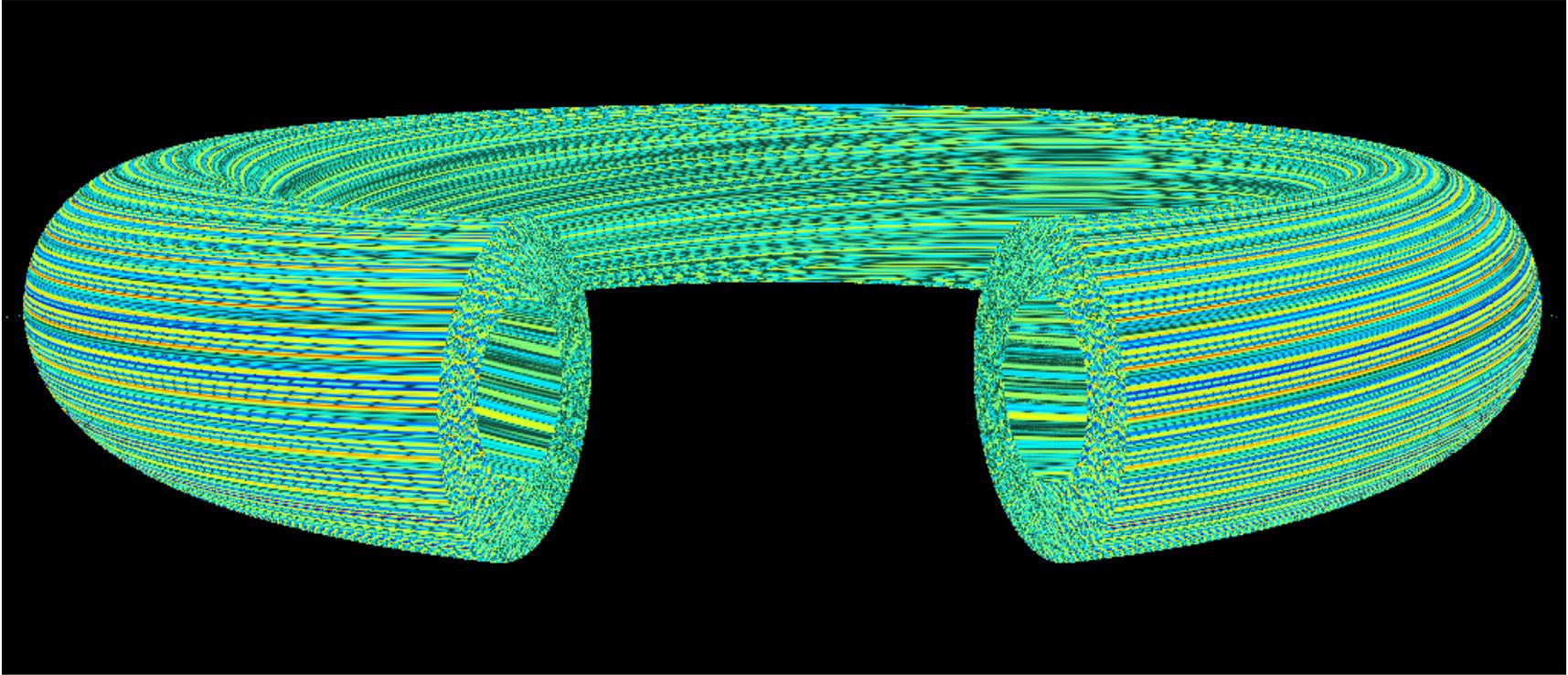
Stable



Similar to how twirling a honey dipper can prevent honey from dripping.

# Cut-away view of tokamak turbulence simulation

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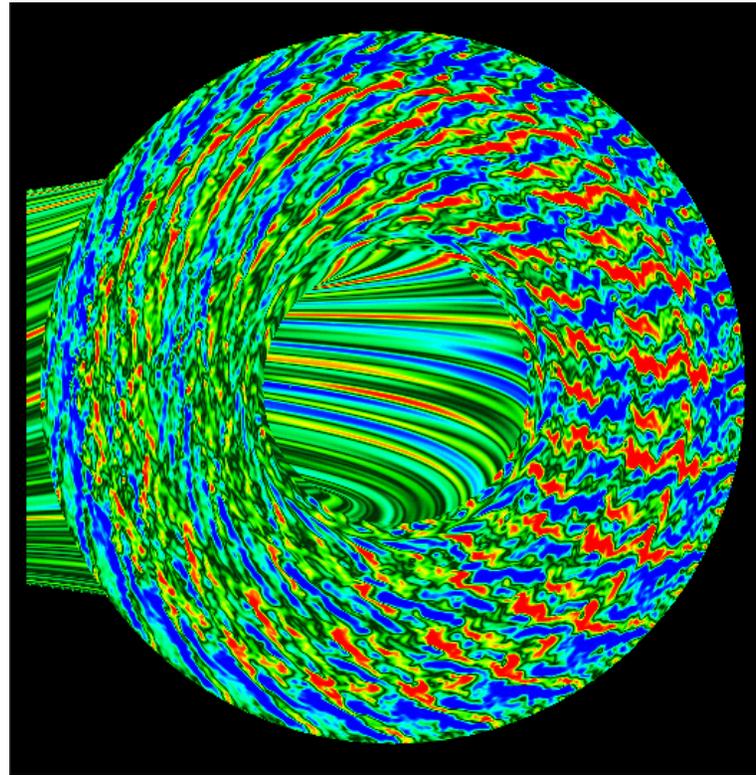


Waltz (General Atomics), Kerbel (LLNL), et.al., gyrofluid simulations. Similar pictures from gyrokinetic particle simulations.

Lots more pictures at [www.acl.lanl.gov/GrandChal/Tok/gallery.html](http://www.acl.lanl.gov/GrandChal/Tok/gallery.html).

# Simulations of Tokamak Plasma Turbulence

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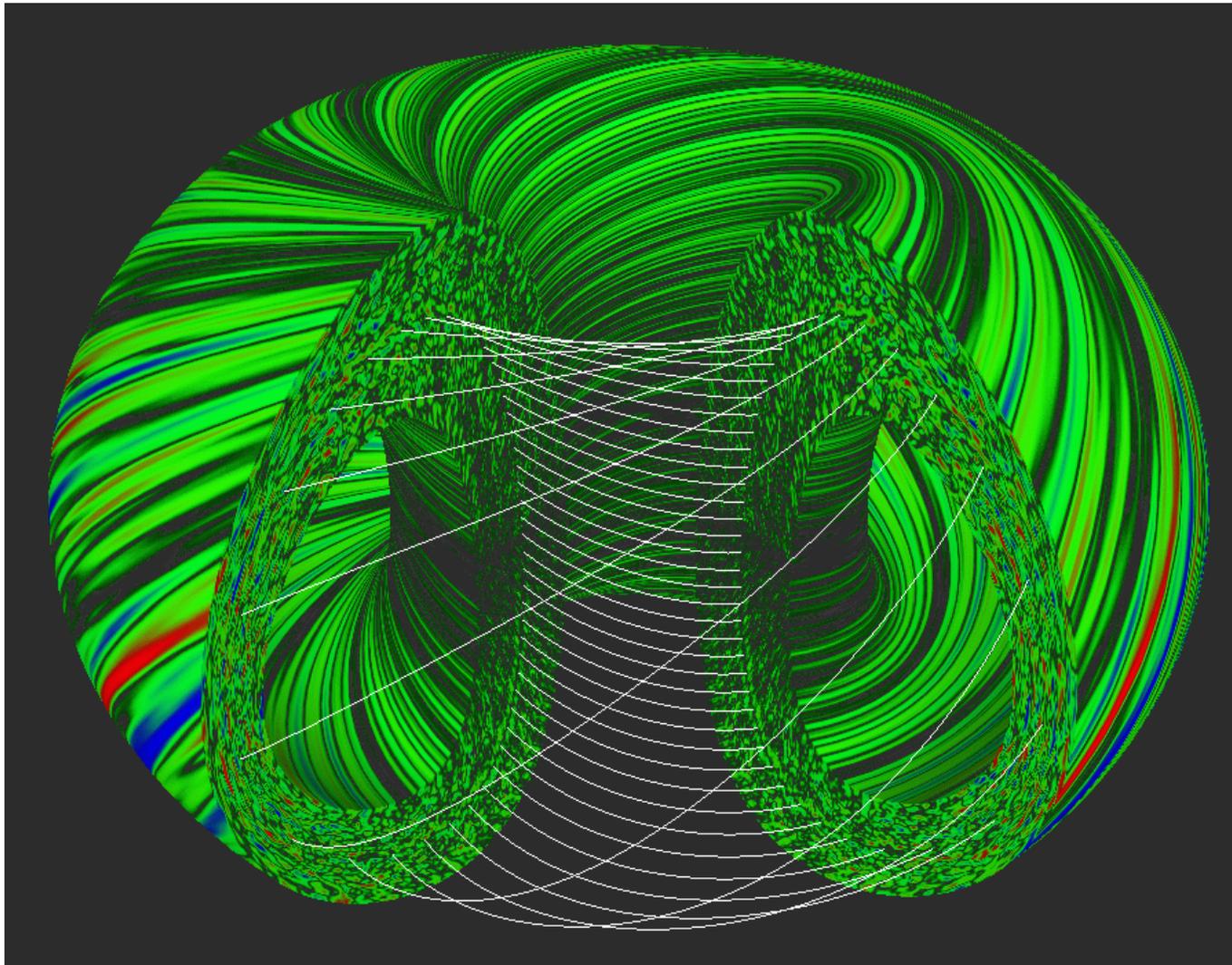


- **Realistic simulations made possible by advances in plasma theory, experimental insights, and parallel supercomputers.**
- **Fundamental science: fascinating physics of plasma turbulence.**
- **Applications: studying ways to reduce turbulence and the cost of a fusion energy power plant.**

General Atomics (San Diego), NERSC (Livermore/Berkeley), PPPL (Princeton), IFS (U.Texas, Austin), ACL (Los Alamos), part of the Numerical Tokamak Project, a DoE/HPCC Computational Grand Challenge.

# Simulations can handle realistic non-circular geometry

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**Turbulence can be reduced by strong plasma shaping in advanced tokamaks, spherical torii, etc.**

General Atomics (San Diego), NERSC (Livermore/Berkeley), PPPL (Princeton), IFS (U.Texas, Austin), ACL (Los Alamos), part of the Numerical Tokamak Project, a DoE/HPCC Computational Grand Challenge.

# "Defrosting Magnetic Field Lines"\*

## MHD (macrostability) vs. Drift waves (microstability)

Electron fluid force Eq:

$$\rho_e \frac{d\vec{u}}{dt} = \underbrace{-\nabla p_e}_{\text{Pressure Force}} - \underbrace{\rho_e \frac{e}{m} \left( \vec{E} + \frac{\vec{u}_e \times \vec{B}}{c} \right)}_{\text{Lorentz } \vec{E} \neq \vec{B} \text{ Forces}} - \underbrace{\nu_{ei} \rho_e (\vec{u}_e - \vec{u}_i)}_{\text{e-i collisional drag}}$$

Drift-waves.      Ideal MHD      + resistive MHD

Compare terms 2 & 4

$$\frac{\nabla p_e}{en_e e B/c} \sim k \rho_i \ll 1 \text{ in MHD}$$

Keeping term 2  $\Rightarrow$  new small-scale modes not in regular MHD theory. "Drift waves", "Ion Temperature Gradient modes", ... "microinstabilities"

$$\omega \approx \omega_* \sim \frac{v_E}{L} k_y \rho_i$$

\*Horton and Berk

Ideal MHD

$$\vec{E} + \frac{\vec{u}_{\text{fluid}} \times \vec{B}}{c} \equiv \vec{E}' = 0$$

Plasma = perfect conductor

$\vec{E}' = 0$  in frame moving with fluid.

Magnetic field is "frozen" in to plasma: Plasma &  $\vec{B}$  move together.

## Component $\parallel$ to $\underline{B}$ :

$$\begin{aligned} \nabla_{\parallel} p_e &= -en E_{\parallel} \\ &= en \left( \nabla_{\parallel} \Phi + \frac{1}{c} \frac{\partial A_{\parallel}}{\partial t} \right) \end{aligned}$$

Drift-modes allow electrostatic modes.  $\underline{E} \times \underline{B}$  move plasma without bending  $\underline{B}$ .

(As  $\beta \uparrow$ , must include magnetic component.)

Ideal MHD  $E_{\parallel} = 0$   
 $\Phi$  must induce  $\underline{B}$



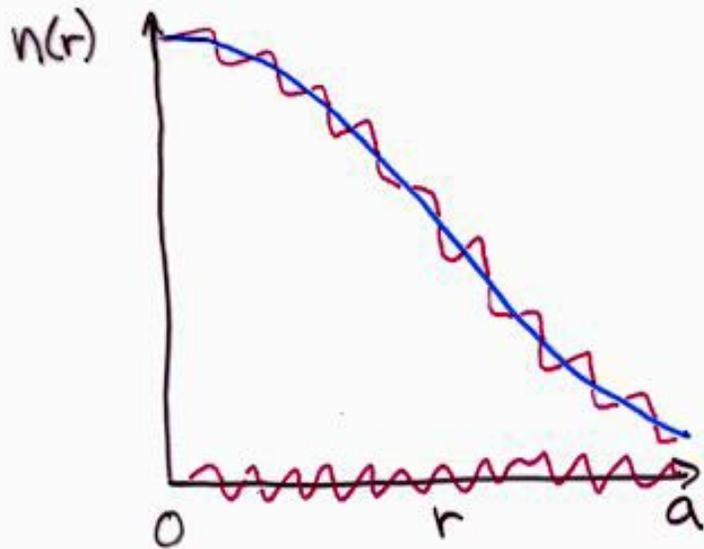
Drift frequency  $\omega \sim \omega_* = \frac{v_{ti}}{L_n} (k_y \rho_i)$

Drift waves lower  $\omega$  than MHD:  $\frac{\omega}{k_{\parallel}} \sim v_{ti}$ ,  $\frac{\omega}{k_{\perp}} \sim$  drift speeds of ions & trapped electrons

Wave-particle resonances important  
 Landau-damping (d inverse)

Small (except for  $k_y \rho_i \sim 1$ )

## Microinstabilities are small-amplitude but still nonlinear



$$n = n_0(r) + \tilde{n}(\underline{x}, t)$$

$$n_0 \gg \tilde{n}$$

$$\text{but } \nabla n_0 \sim \nabla \tilde{n}$$

↑  
Can locally flatten  
or reverse total gradient  
that was driving instability.

\* Turbulence causes loss of plasma to the wall,  
but confinement still  $\times 10^5$  better than without  $\underline{B}$ .

$$\text{If no } \underline{B}, \text{ loss time } \sim \frac{a}{v_t} \sim 1 \text{ } \mu\text{sec}$$

$$\text{with } \underline{B}, \text{ expts. measure } \sim 0.1 - 1.0 \text{ sec.}$$

# Extreme Range of Time (& Space) Scales To Cope With in Plasmas

PPPL#91X0674

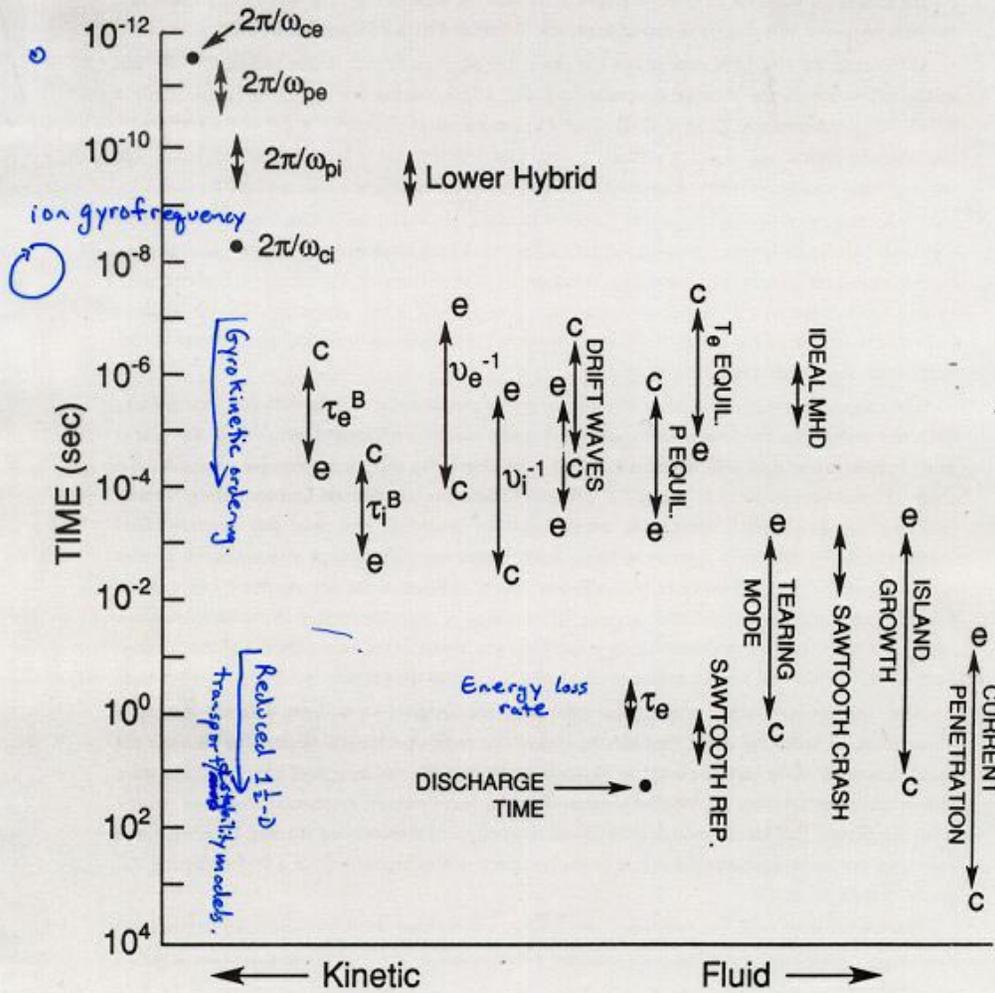


FIGURE III-1: TIME SCALES FOR TPX.  
e DENOTES EDGE AND c DENOTES CENTER

Vlasov, Boltzmann, Liouville Eq:

Particle Distribution

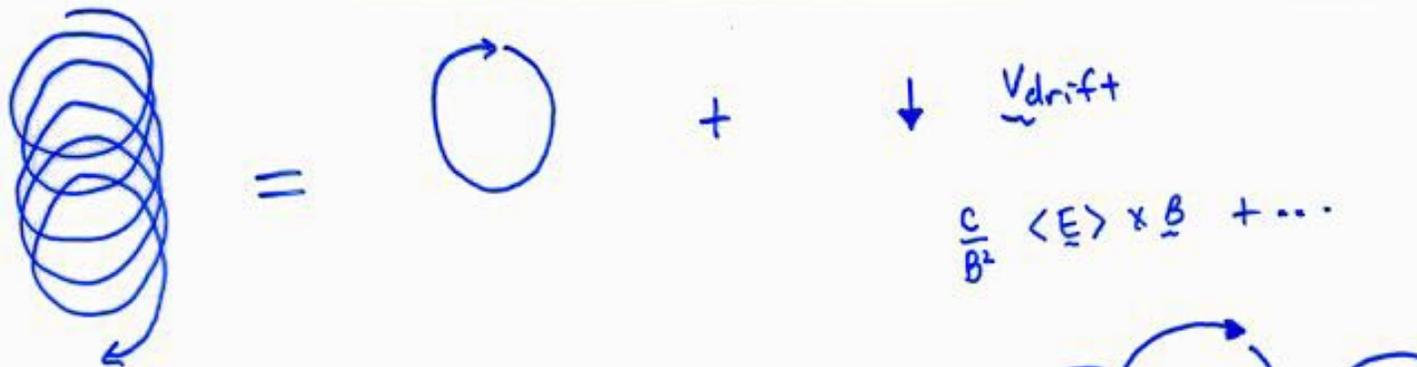
$$\frac{\partial f(\underline{x}, \underline{v}, t)}{\partial t} + \underline{v} \cdot \frac{\partial f}{\partial \underline{x}} + \frac{q}{m} \left( \underline{E} + \frac{\underline{v} \times \underline{B}}{c} \right) \cdot \frac{\partial f}{\partial \underline{v}} = C(f)$$

Nonlinear,  $\underline{E}$  &  $\underline{B}$  depend on  $f$  through Maxwell's Eqs.

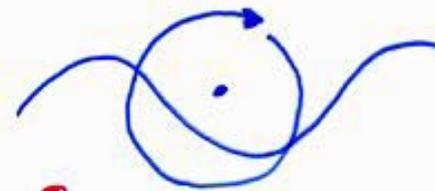
Nonlinear Gyrokinetic Eq. 1982-88

(Frieman & Chen, W.W. Lee, Dubin, Krommes, Hahm, Brizard...)

linear gyrokinetics  
1960's & 70's.



Possible to eliminate fast gyrofrequency  $\Phi$   
time scales & retain nonlinear dynamics  
&  $k_{\perp} \rho_i \sim 1$



# Major Theoretical & Algorithmic Speedups

For moderate size tokamak  $\frac{L}{\rho} \sim 400$

## Speedup

Nonlinear Gyrokinetic Eq. Avg. Fast Gyromotion

$$\frac{\Omega_{ci}}{w_*} \sim \frac{L}{\rho}$$

x400

## $\delta f$ Method

(Dimitis, Kotschenreuther, Parker, Lee, ...)

Reduce  $\frac{1}{\sqrt{N}}$  Noise

$$\left(\frac{\delta f}{f}\right) \sim \frac{f}{L} \times 10^3 - 10^5$$

$$f(\underline{x}, \underline{v}, t) = \underbrace{f_0}_{\text{Smooth}} + \delta f$$

Discrete particle w/ weights

Also useful for accelerators, ICF, ...

## Efficient Coordinates

(Cowley, Beer, Hammett, Dimitis)

Nonlinear extension of ballooning transform related to Bloch transform in solidstate

Coordinates aligned with twisting sheared  $\underline{B}$

$$\frac{\Delta_{\perp}}{\Delta_{\parallel}} \sim \frac{L}{\rho} \frac{q}{qR}$$

x40

Reduced simulation size flux-tube

$$\left(\frac{a}{L_{\perp \text{sim}}}\right)^2$$

x ~ 16

Gyro-Landau Fluid Closure Approx.

x 20 +

Bounce-avg electrons // implicit methods

x 10 - 50

Massively Parallel Supercomputers past 10 years

x 100

Caution compute time  $\propto \left(\frac{L}{\rho}\right)^3$

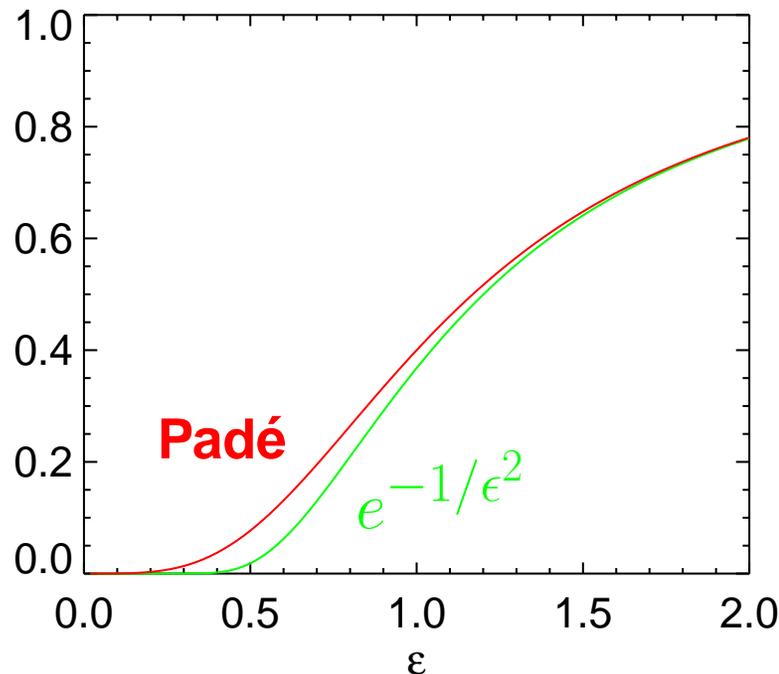
# Fluid closure approximations for collisionless limit

(Hammett & Perkins, Chang & Callen, Dorland, Beer, Waltz, ...)

$$\begin{aligned}
 f(\vec{x}, v_{\parallel}, v_{\perp}, t) &\rightarrow \int d^3v f v_{\parallel}^j v_{\perp}^k \\
 5D + t &\rightarrow 3D + t \times \mathbf{6 \text{ moments}} \\
 &\quad \text{(Density, avg. flow, parallel and perp pressures and heat fluxes)}
 \end{aligned}$$

Example closure: Heat conductivity:

$$\chi_{\parallel} = \frac{v_t^2}{\underbrace{\nu_{collisions}}_{\text{Chapman-Enskog Collisional Transport}} + \underbrace{|k_{\parallel}| v_t}_{\text{Collisionless limit n-pole approx. to Landau damping (and inverse) } 1/|k_{\parallel}| \rightarrow \text{non-local integral operator}}}$$



$$e^{-\omega^2/k^2 v_t^2} = e^{-1/\epsilon^2}$$

Taylor expansion vanishes to all orders, but Padé approx. useful:

$$\approx \frac{\epsilon^4}{0.5 + \epsilon^2 + \epsilon^4}$$

# IFS-PPPL Transport Model

Kotschenreuther, Dorland, Beer, Hammett '94

- Based on nonlinear gyrofluid simulations of ITG turbulence to map out structure of ion thermal conductivity  $\chi_i$ , & on linear gyrokinetic calc of growth rates and critical gradients.

$$\chi_i = \rho_i^2 \frac{v_{ti}}{R} \left( \frac{R}{L_T} - \frac{R}{L_{T,crit}(p_j)} \right)^{1,1/2} F(p_j) \underbrace{\left( 1 - \frac{\overbrace{\gamma_{shear}}}{\underbrace{\gamma_{lin}}}}_{\text{Waltz gyrofluid fit}}$$

Hahm-Burrell ExB shear

$$p_j = \left( \frac{R}{L_T}, \frac{R}{L_n}, \frac{T_i}{T_e}, q, \hat{s}, Z_{eff}, \nu_*, \frac{r}{R}, \dots \right)$$

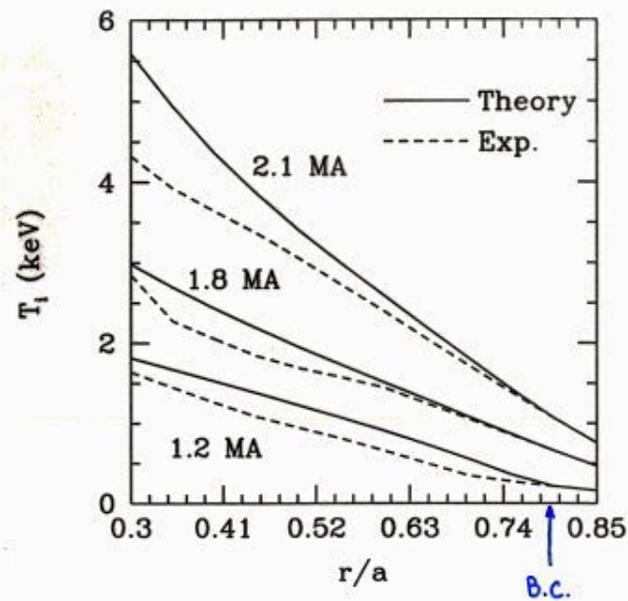
$\chi_e/\chi_i = \text{quasilinear}$

- Brought together scalings from many analytic theories into a single formula. Comprehensive enough to explain many observed trends in standard tokamak operating regimes, including some improved confinement regimes (given edge B.C.'s)

**IFS-PPPL transport model represented a significant advance. But a more complete model is needed:**

- **advanced tokamak regimes (negative shear, high  $\beta$ , strong shaping)**
- **internal transport barriers: suppress  $\chi_i$  &  $D_e$ , but large  $\chi_e$  ???!**
- **particle and momentum transport (presently just heat transport)**
- **edge turbulence**
- **better shear in equilibrium  $E \times B$ ,  $\omega_*(r)$ ,  $\eta_i(r)$**
- **better zonal flows, gyrofluid/gyrokinetic diffs**

## Comparison of measured & predicted ion temperature

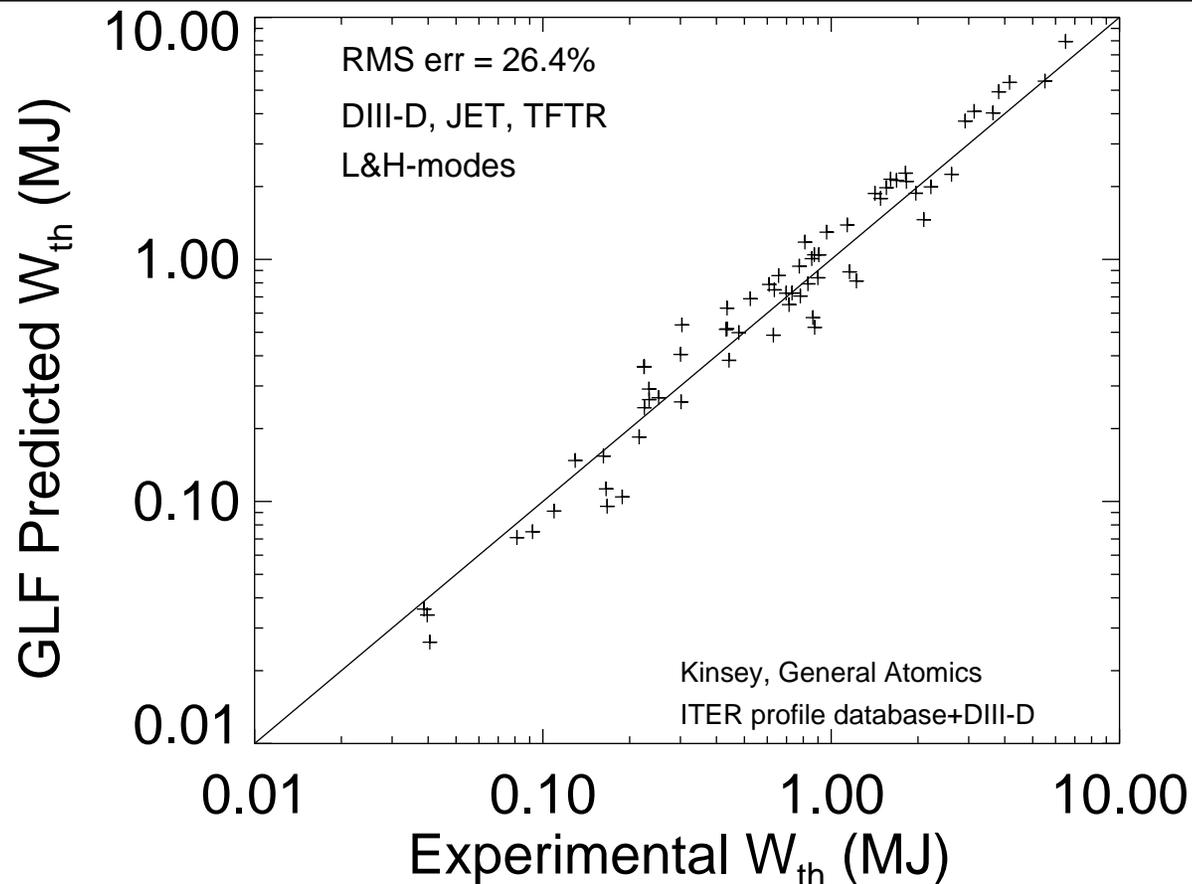


$$\text{Heat flux} = -n\chi_i \nabla T_i$$

Example of  $T_i$  predicted by IFS-PPPL model of ion thermal diffusivity  $\chi$  due to ITG-turbulence. Follows improvement with plasma current ( $\approx 30\%$  error, but sawteeth neglected..., better current scaling than most early models).

Kotschenreuther, Dorland, Beer, Hammett, Phys. Plas. 95

# Transport Model Based on Turbulence Simulations Follows Many Experimental Trends



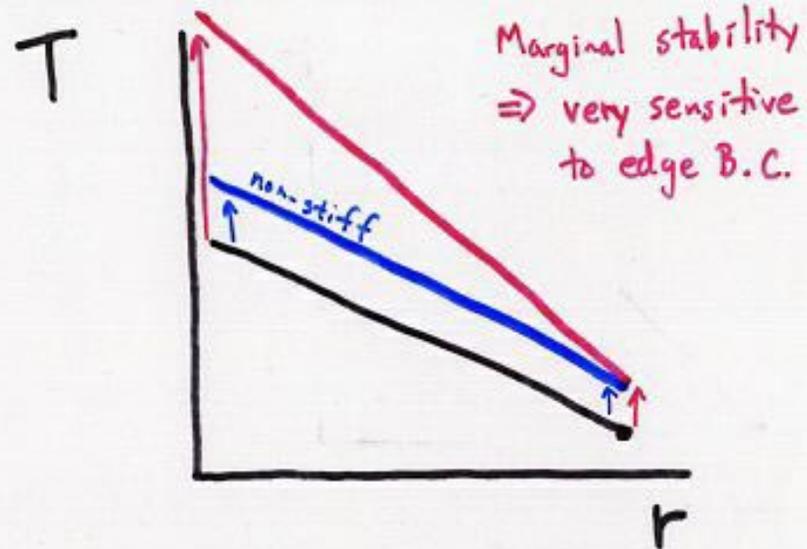
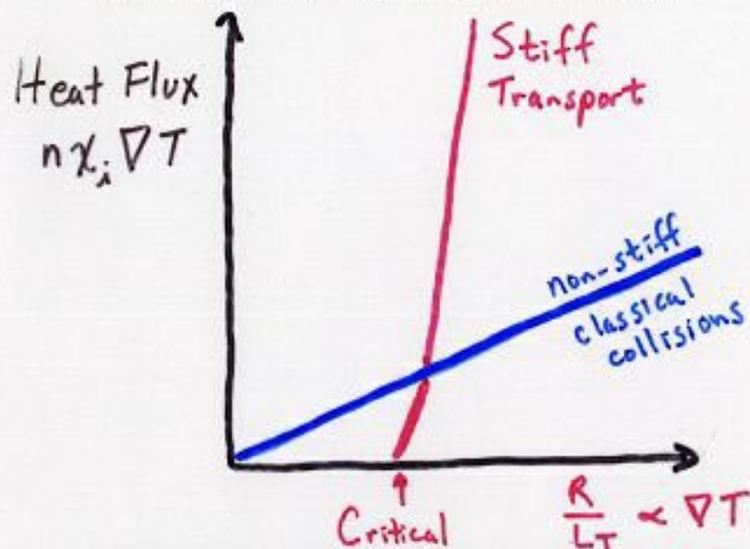
- **GLF23 transport model by Waltz et.al fit to Beer et.al. nonlinear 3-D gyrofluid simulations of ITG/trapped-electron turbulence.**

- Encouraging results so far, but many caveats: uses measured density and rotation profiles, uses measured temperatures at  $r/a = 0.9$ , electrostatic turbulence simulations need extension to magnetic fluctuations, gyrofluid/gyrokinetic discrepancy, etc... Much future work needed to be more accurate over a wider range of plasma parameters.

- **Rescaled GLF23,  $\downarrow \chi$  and  $E \times B$  shear, improves to RMS error  $\approx 19\%$ .**

# Large $\chi$ 's predicted by many 1980's analytic ITG theories lead to the proposal that temperature gradients would be forced to near marginal stability

(for example Biglari, Diamond, Rosenbluth, Phys. Fluids B1, 109 (1989), Horton et.al. Phys. Fluids B4, 953 (1992), Bateman PF B4, 634 (1992) and refs therein).



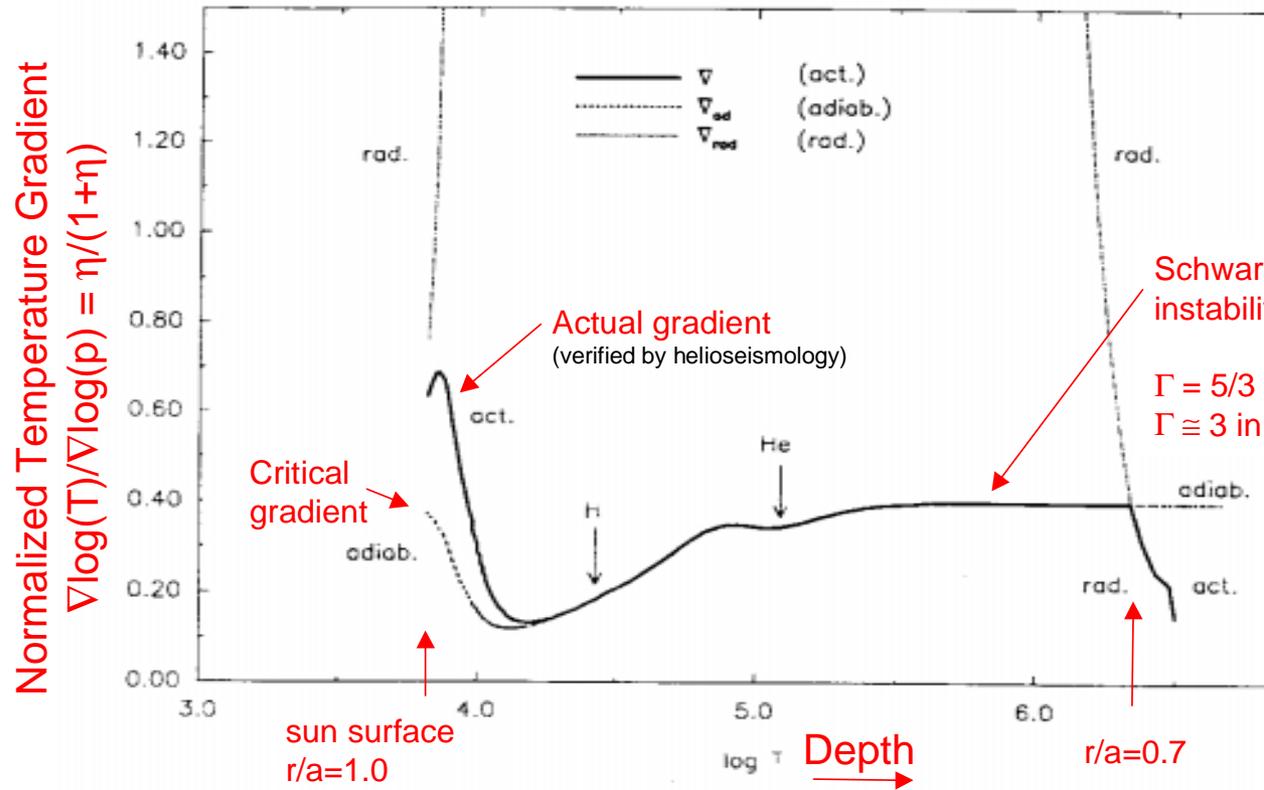
Resulting temperature profile is more sensitive to critical gradient than to magnitude of  $\chi$ . Core temperature becomes very sensitive to boundary condition if there is perfect marginal stability:

$$T = T_0 e^{-r/L_{Tcrit}}$$

$$T(r) = T_0 e^{-r/L_{Tcrit}}$$

Helps explain experimental sensitivity to edge boundary conditions (neutral recycling, wall conditions, supershots, edge transport barriers). Similar to the largest fusion reactor in the solar system....

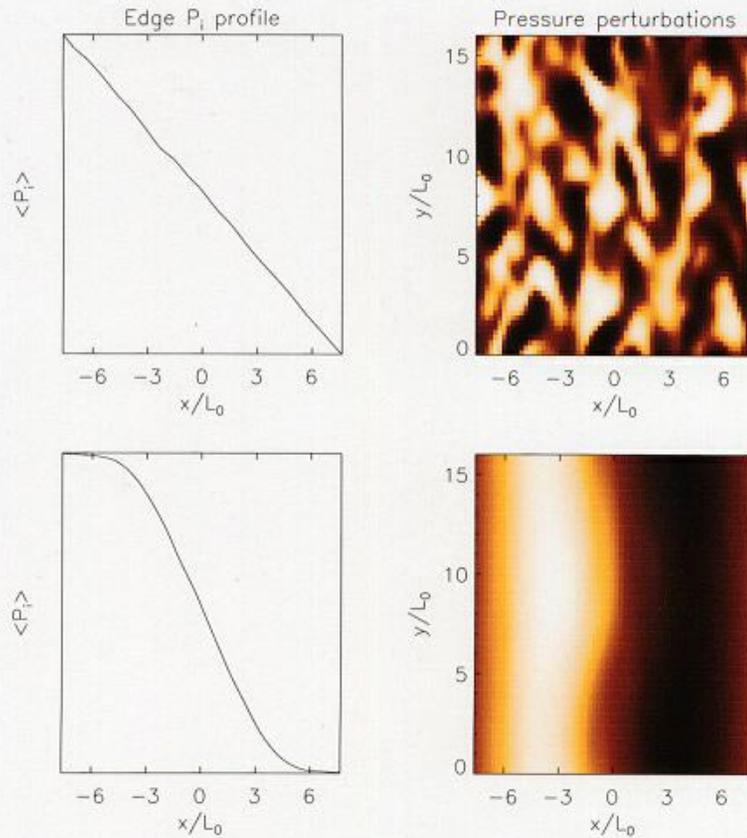
# Solar Convection Zone Near Marginal Stability



**Figure 2:** Temperature gradients  $\nabla_{rad}$ ,  $\nabla_{ad}$ , and  $\nabla$  as functions of  $\log T$  for a mixing length of the solar convection zone (Spruit 1977). At the bottom of the convection zone ( $\log$  depth  $\approx 2 \cdot 10^5$  km) the actual temperature gradient changes from  $\nabla \approx \nabla_{rad}$  to  $\nabla \approx \nabla_{ad}$ . Superadiabaticity becomes significant only in the surface layers ( $\log T < 4$ ) where the wind is driven. The ionization regions of hydrogen and helium are indicated; the adiabatic temperature gradient decreases there due to the effect of latent heat. (from Spruit, 1977 Ph.D., in Schüssler '92)

## Example of Edge Turbulence Simulations

EM effects play fundamental role in 3D simulations of L/H transition

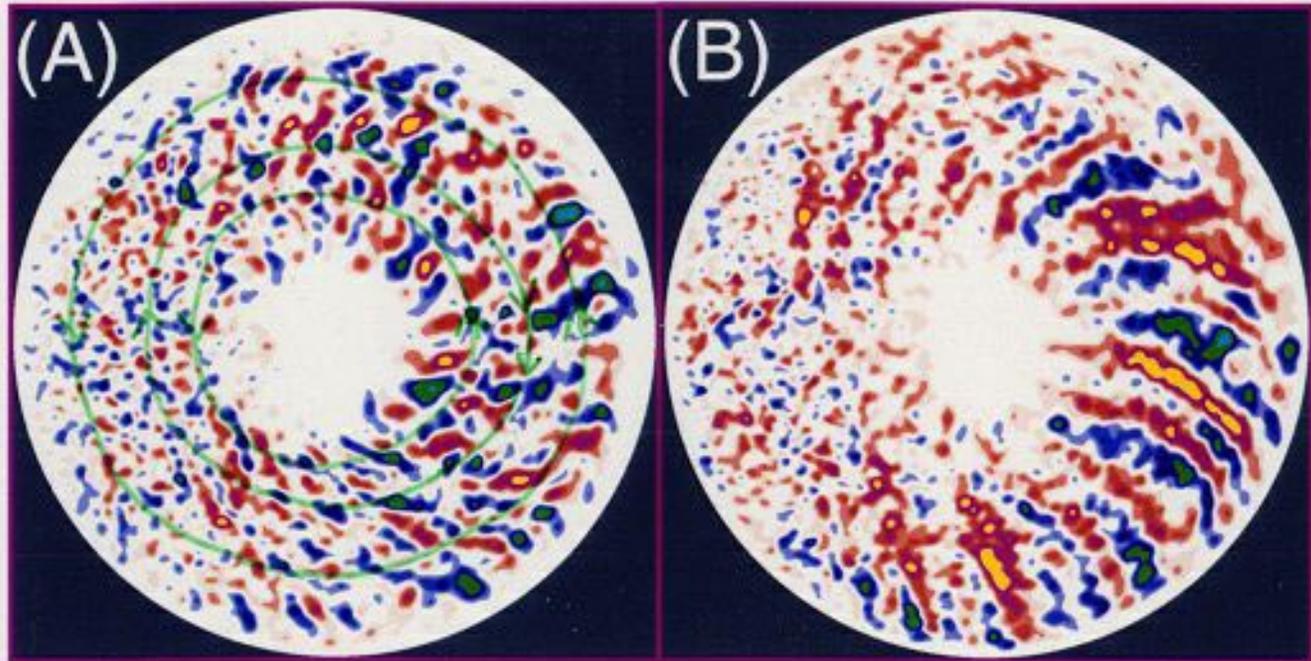


- Increasing edge pressure gradient in simulations leads to stronger magnetic fluctuations. Above critical threshold, these cause transport barrier to spontaneously form.

3D Drift-Braginskii  
Rogers & Drake

## Gyrokinetic Simulations of Plasma Microinstabilities: turbulence decorrelation by zonal flows

Sheared  
Zonal flows  
(generated by  
turbulence)  
causes eddies  
to break up  
radially  
⇒ ↓ transport.



With Flow

Without Flow

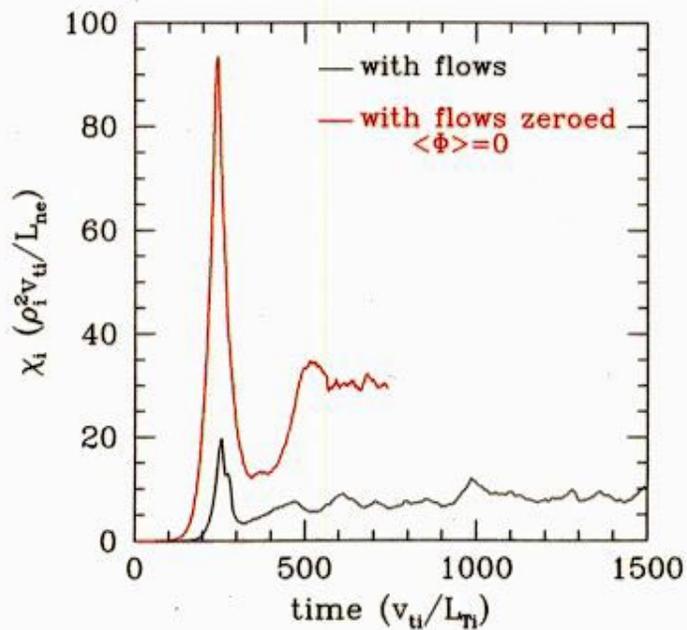
Turbulence reduction via sheared plasma flow (A),  
compared to case with flow suppressed (B).

[Z. Lin *et al.*, **Science** 281, 1835 (1998)]

Higher quality image available at [w3.pppl.gov/~zlin/gyrokinetic.html](http://w3.pppl.gov/~zlin/gyrokinetic.html)

## Turbulence-generated Zonal Flows

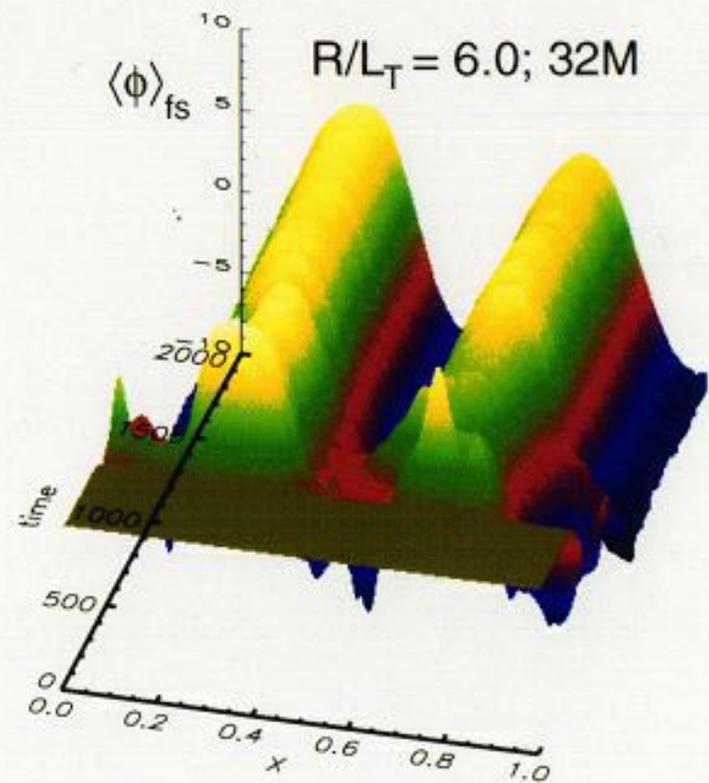
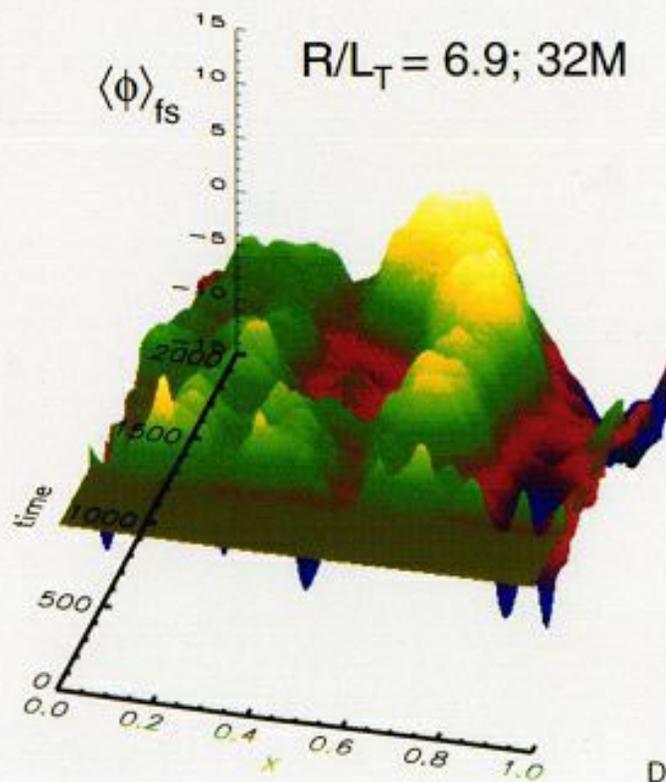
Hasegawa-Wakatani '87 3D resistive drift interchange  
Carreras et.al. '91 } suggests this happens only very  
Diamond + Kim '91 } close to the edge  
Dorland et.al. '93 } Show zonal flows very important  
Beer et.al. '93 } for core ITG turbulence also.



Rosenbluth - Hinton show there should be a component of the zonal flows that are linearly undamped (missed by original gyrofluid closure) +  $\propto \downarrow$  even more, as shown by Dimits ~~particle~~ particle simulations.

## FLUX-SURFACE-AVERAGED POTENTIALS

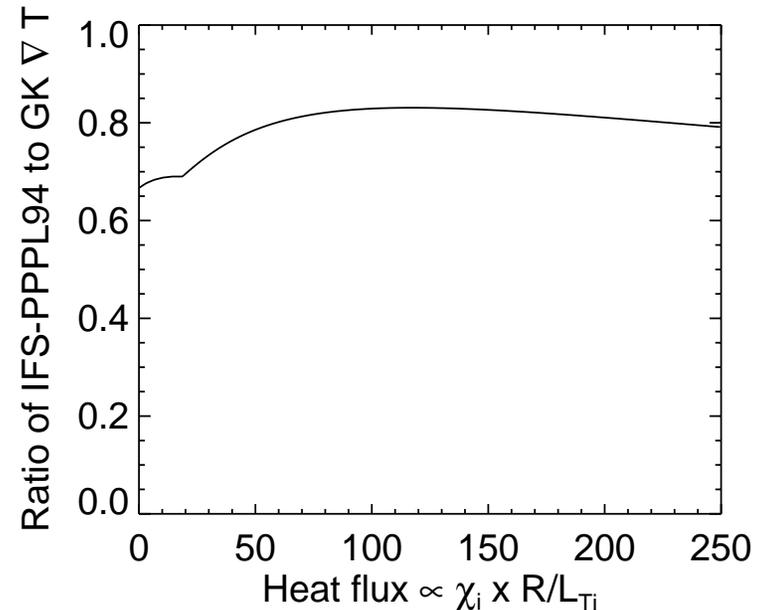
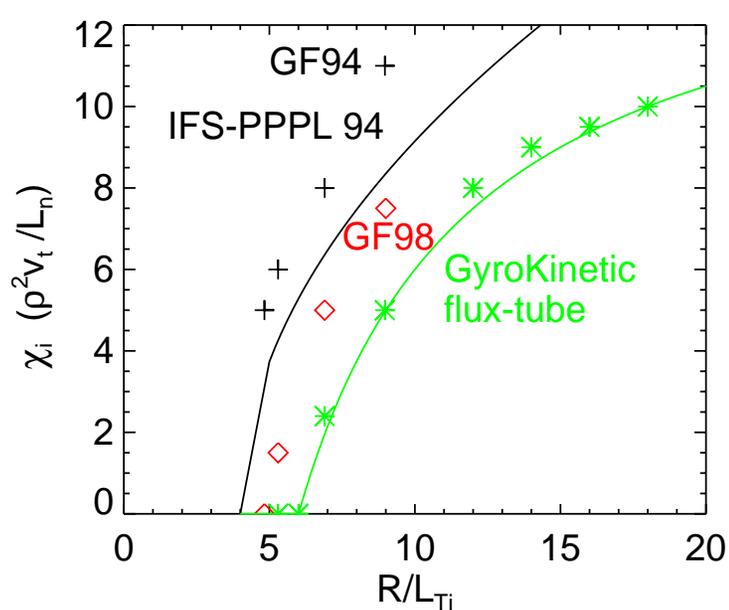
- Zero- $\chi_i$  states show stationary  $\langle\phi\rangle_{fs}$  structures, unlike nonzero- $\chi_i$  states.
- Max. ExB shear  $\approx 3\times\gamma_{max}$ ; ExB stabilization and radial transport barriers are playing a significant role.



Dimits et al., LLNL

# Gyrofluid/gyrokinetic (GF/GK) simulation differences

→ 20-33% change in predicted temperature gradient



- Dimits (LLNL): good convergence in his gyrokinetic particle simulations
- New neoclassical gyrofluid closure significantly improves GF/GK comparison.
- Turning this plot around, for a fixed amount of heat flux  $\propto \chi \nabla T$ , the temperature gradient predicted by the original gyrofluid-based IFS-PPPL model is 20-33% low. But  $P_{fusion} \propto T^2$ , and so may increase by  $\times 2$  or more.
- Nonlinear upshift in critical gradient may depend on: Rosenbluth-Hinton undamped zonal flows  $\uparrow$  with elongation (W. Dorland),  $\downarrow$  with weak collisions (Z. Lin),  $\downarrow$  ?? with non-adiabatic electrons [may limit inverse cascade that drives zonal flows (Diamond, Liang, Terry-Horton, Waltz, ...) and  $\uparrow$  turbulent viscosity].

# CONCLUSIONS

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**Major progress has been made during the past 10 years in direct 3D simulations of plasma turbulence and in reduced transport models.**

**Reasonable agreement with core temperature profiles ( $\sim 30\%$ ) in many cases, but more work needed to resolve significant uncertainties (edge turbulence, zonal flows, electron dynamics, ...).**

**Relatively complete simulations should be achievable soon...<sup>†</sup>**

**Also: many ways to reduce turbulence and improve performance (sheared flows, IBW, edge beams, density peaking, high beta advanced tokamak designs with strong Shafranov shift and shaping, ...)**

<sup>†</sup>**But needs a lot of hard work, more complete physics in codes, and new generation of computers.**

**P.S.: The content of the above slides is the same as I used in my talk at the Atlanta APS meeting (March, 1999), though I converted a few of them to be typeset instead of just scanning in handwritten slides.**