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Structure of nonlocal gradient-drift instabilities in Hall $\mathbf{E} \times \mathbf{B}$ discharges

Ivan Romadanov,^{1,a)} Andrei Smolyakov,^{1,b)} Yevgeny Raitsev,^{2,c)} Igor Kaganovich,^{2,d)} Tang Tian,^{3,e)} and Sergei Ryzhkov^{4,f)}

¹Department of Physics and Engineering Physics, University of Saskatchewan, Saskatoon, Saskatchewan S7N 5E2, Canada

²Plasma Physics Laboratory, Princeton University, Princeton, New Jersey 08543, USA

³University of Science and Technology of China, Hefei, Anhui 230026, People's Republic of China

⁴Bauman Moscow State Technical University, Moscow 105005, Russia

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Gradient-drift (collisionless Simon-Hoh) instability is a robust instability often considered to be important for Hall plasma discharges supported by the electron current due to the $\mathbf{E} \times \mathbf{B}$ drift. Most of the previous studies of this mode were based on the local approximation. Here, we consider the nonlocal model which takes into account the electron inertia as well as the effects of the entire profiles of plasma parameters such as the electric, magnetic fields, and plasma density. Contrary to local models, nonlocal analysis predicts multiple unstable modes, which exist in the regions, where local instability criteria are not satisfied. This is especially pronounced for the long wavelength modes which provide larger contribution to the anomalous transport. *Published by AIP Publishing.* [<http://dx.doi.org/10.1063/1.4971816>]

I. INTRODUCTION

Plasmas with crossed electric \mathbf{E} and magnetic \mathbf{B} fields are often used as ion sources in various applications. For moderate values of the magnetic field, the discharge is maintained by the electron $\mathbf{E} \times \mathbf{B}$ drift, while allowing the extraction, separation, and acceleration of unmagnetized ions. Such regimes, generally referred here as Hall plasmas, are widely employed in material processing (magnetrons^{1–3}) and electric propulsion (Hall thrusters^{4,5}) devices. The electron current across the magnetic field in such systems often exceeds the collisional value by orders of magnitude. Presumably, this anomalously large electron current is a result of plasma instabilities which are present in these devices for a wide range of plasma parameters and operational regimes. Despite long history of Hall plasmas applications, the understanding of the nature and mechanisms of plasma instabilities and plasma turbulence is still lacking.

Among different types of unstable modes, relevant to $\mathbf{E} \times \mathbf{B}$ Hall plasmas, gradient-drift modes have long been considered as a possible source of fluctuations.^{6,7,18} These modes are closely related to the so called anti-drift mode⁸ that exists in Hall plasmas with density gradient,

$$\omega = \frac{k_{\perp}^2 c_s^2}{\omega_*}, \quad (1)$$

where $c_s^2 = T_e/m_i$ is the ion-sound velocity and

$$\omega_* = k_y v_* = -\frac{k_y T_e}{e B_0 L_n}$$

is the electron diamagnetic drift frequency due to density gradient; L_n is the characteristic length scale of the density gradient, $L_n^{-1} = \nabla \ln n_0(x)$; $k_{\perp}^2 = k_y^2 + k_x^2$ is the wave vector perpendicular to the magnetic field, $\mathbf{B}_0 = B_0 \hat{\mathbf{z}}$. Note that the standard drift waves do not exist in Hall plasmas with unmagnetized ions. The frequency of the anti-drift mode (1) in fact does not depend on the electron temperature and the dispersion relation can also be written in the form $\omega = -\omega_{ci} k_{\perp}^2 L_n / k_y$, where $\omega_{ci} = e B_0 / m_i$ is the ion cyclotron frequency. The condition $\omega > k_z v_{Te}$, where k_z is along the magnetic field lines, is required so that the mode is propagating almost perpendicular to the magnetic field. When the latter condition is not satisfied, the mode converts into the ion sound mode, $\omega^2 = k^2 c_s^2$, which may propagate at a finite angle with respect to the magnetic field.⁹

When the stationary electron current due to the $\mathbf{E} \times \mathbf{B}$ drift is present, the anti-drift mode becomes unstable due to coupling with the ballistic mode $\omega = k_y V_0$.^{10,11} Resulting gradient drift instability is described by the dispersion equation,

$$-\frac{\omega_{ci} k_{\perp}^2 L_n}{k_y} = \frac{\omega^2}{\omega - \omega_0}, \quad (2)$$

where $\omega_0 = \mathbf{k} \cdot \mathbf{V}_0$ is the azimuthal (closed drift) flow of electrons in crossed $\mathbf{E} \times \mathbf{B}$ fields and $\mathbf{E}_0 = E_0 \hat{\mathbf{x}}$. This is the reactive instability of negative energy perturbations with a phase velocity below the stationary $V_0 = \mathbf{E} \times \mathbf{B} / B^2$ velocity. This mode is referred here as the collisionless Simon-Hoh instability.^{10–12} The condition $\mathbf{E} \cdot \nabla n > 0$ is required for the instability.¹³

The gradient-drift instability was identified in early Hall thruster experiments^{6,14} as related to violent large scale structures (spokes). It was shown⁷ that properly profiled magnetic field improves the stability of the Hall thruster. The modified condition for the instability was derived in the form $\mathbf{E} \cdot \nabla(n/B) > 0$.

^{a)}Electronic mail: ivr509@mail.usask.ca

^{b)}Electronic mail: andrei.smolyakov@usask.ca

^{c)}Electronic mail: yraitses@pppl.gov

^{d)}Electronic mail: ikaganovich@pppl.gov

^{e)}Electronic mail: tt961108@gmail.com

^{f)}Electronic mail: svryzhkov@bmstu.ru

The gradient drift instability in plasmas with inhomogeneous magnetic field was revisited in Refs. 15 and 16 where the full compressibility was included resulting in the modified dispersion relation

$$\frac{\omega_* - \omega_D}{\omega - \omega_0 - \omega_D} = \frac{k_{\perp}^2 c_s^2}{\omega^2}, \quad (3)$$

where

$$\omega_D = k_y v_D = -\frac{2k_y T_e}{eB_0 L_n}$$

is the magnetic drift velocity due to the axial gradient of the magnetic field; $L_B^{-1} = \nabla \ln B_0(x)$ is the characteristic length of such a gradient. One consequence of this modification was that the magnetic field gradient enters in the combination $\nabla(n/B^2)$ rather than $\nabla(n/B)$ in the original work.⁷ In a finite temperature plasma, there also exists an additional contribution due to magnetic field gradient because of the compressibility of the electron diamagnetic velocity¹⁸ (the term with ω_D in the denominator on the left hand side). It was also shown in Ref. 15 that finite temperature fluctuations are also important in situations with a strong magnetic field gradient (complete dispersion relation for this case is given in Ref. 16).

The collisionless Simon-Hoh instability has been considered among the most important sources of plasma fluctuations, in particular, as a cause of large scale structures or spokes.^{19–22} Much of the previous work on gradient-drift instabilities,^{23–25} see also other works cited in Ref. 23, was done within the local approximation neglecting the effects of the density and electric field profiles. The authors in Ref. 26 have considered the nonlocal solution to compare the eigenmode structure and eigenmode frequency with the data from Hall thruster experiments.²⁶ However, the theoretical model used in Ref. 26 (following to Refs. 27 and 28) does not apply to the collisionless Simon-Hoh instability in which ions are not magnetized. The global analysis has been employed in Refs. 23, 29, and 30 for linear stability studies of Hall thruster for rather general fluid model that included effects of ionization, collisions, and heat flow. This model involved several different unstable modes and destabilization mechanisms; therefore, it is somewhat difficult to directly compare the results of the local and nonlocal analysis.

The standard collisionless Simon-Hoh instability is considered to be a low frequency mode with the eigenfrequency well below the $\mathbf{E} \times \mathbf{B}$ frequency.¹⁶ However, it was shown recently¹⁷ that depending on the value of the k_x wave vector (along the electric field direction), the mode frequency becomes comparable to the $\mathbf{E} \times \mathbf{B}$ frequency, $\omega_0 = \mathbf{k} \cdot \mathbf{V}_0$, and thus, for the short-wavelengths, may become comparable with the low hybrid frequency $\omega_{LH} = (\omega_{ce}\omega_{ci})^{1/2}$ which required to be accounted for electron inertia.

The goal of this work is to develop a nonlocal model for the collisionless Simon-Hoh instability and related higher frequency lower-hybrid modes and investigate the role of density and electric field profiles on the eigenmode structure, real part of the frequency, and growth rate in conditions

typical for Hall plasma experiments. In case of the general profiles of the electric field and plasma density, the non-local stability problem is reduced to the eigenvalue problem which is solved here by the spectral method with Chebyshev polynomials.³⁵ The results from the spectral code have been verified by a finite-difference differentiation matrix, and shooting method. Both spectral and finite-difference methods show good convergence and precision for low mode numbers. However, higher modes require significant increase in number of polynomials or number of steps for finite difference method to achieve the accurate solution.

For our simulations, we used parameters relevant to two Hall plasma devices: Penning discharge and Hall thruster. Typical plasma parameters and profiles for such devices were taken from the experiments at Princeton Plasma Physics Laboratory.^{31–34}

Penning discharge is a device with cylindrical geometry, in which the applied magnetic field is along the z -axis, and the electric field is in the radial direction. Plasma is created by the electron beam injected along the z -axis, from either filament cathode or plasma cathode such as a hollow cathode or RF plasma cathode. Magnetic field is created by Helmholtz coils. The experiments with the Penning discharges demonstrate range of instabilities and structures (spokes) similar to the Hall thruster^{33,34} and provide easier access for plasma diagnostics. Typical experimental parameters are: electric field $E = -50$ V/m, magnetic field $B = 50$ G, $v_0 = 1 \times 10^4$ m/s, $L_n = -0.1, \dots, -0.04$ m, system length $L = 0.1$ m, $\omega_{ci} = 3.7 \times 10^3$ s⁻¹, $\omega_{LH} = (\omega_{ce}\omega_{ci})^{1/2} = 1.8 \times 10^6$ s⁻¹. For calculations in Sections III A and III B, the parameters of Penning discharge were considered in slab geometry, with the magnetic field along the z -axis, and electric potential and density gradient along the x -axis.

Hall thruster is another device of interest in which the gradient-drift instabilities play an important role. Typical device has coaxial or cylindrical, for the Cylindrical Hall Thruster (CHT),^{31,32} geometry. The discharge in the Hall thruster is created in the channel between the anode, which can be used as a gas distributor, and the cathode-neutralizer. A gas propellant, typically argon or xenon, is supplied to the channel through anode and then is ionized by high energy electrons from the cathode and as well electrons heated in the discharge. Magnetic field is sufficiently strong to make the electron Larmor radius much smaller than the chamber characteristic length thus creating the strong azimuthal $\mathbf{E} \times \mathbf{B}$ current of electrons. Collisions allow the electrons move to the anode in the axial direction, across the magnetic field and in the direction of the stationary electric field. The azimuthal current due to the $\mathbf{E} \times \mathbf{B}$ drift can be tens times larger than the axial electron current. The axial field accelerates the ions towards the channel exhaust, where they are neutralized by electrons from the cathode neutralizer. The thrust is a reaction force to this electrostatic acceleration, applied to the magnetic circuit. Even though azimuthal current is responsible for the ion thrust, axial current is important as well, because it completes the electric circuit.

The axial electron current in the Hall thruster is typically an order of magnitude larger than the collisional value. This anomalous current is created by fluctuations

driven by the electric field and gradients of the magnetic field, plasma density, and electron temperature which are strongly inhomogeneous along the thruster axis. For the Hall thruster configuration, the electric field, magnetic field, and density gradients are in the x - (axial) direction; the y -direction is periodic (azimuthal) direction. Typical magnetic field, electric field, and density profiles will be presented in Section IV.

II. BASIC EQUATIONS

Here, we formulate the eigenmode problem for the slab geometry with uniform axial magnetic field $\mathbf{B}_0 = B_0 \hat{\mathbf{z}}$, and inhomogeneous plasma density $n_0 = n_0(x)$ and equilibrium potential $\phi_0 = \phi_0(x)$. We consider the nonlocal modes in the form $\tilde{\Phi} = \tilde{\phi}(x) \exp(-i\omega t + ik_y y)$. The y direction is assumed to be periodic, a finite length interval is considered in the radial direction x . This geometry is an approximation for the cylindrical geometry of the Penning discharge with the axial magnetic field as in Ref. 31, in which case y is an azimuthal direction and x is radial.

The linearized ion momentum equation has the form

$$-i\omega m_i \mathbf{v}_i = -e \nabla \tilde{\phi}. \quad (4)$$

Here, we assume cold ions and neglect the effect of the magnetic field on ion motion. The linear continuity equation for ions is

$$-i\omega n_i + \nabla \cdot (n_i \mathbf{v}_i) = 0. \quad (5)$$

After substitution of \mathbf{v}_i into the ion continuity equation, one finds

$$-i\omega n_i - \frac{ie}{m_i \omega} (n_0 \nabla^2 \phi + \nabla n_0 \cdot \nabla \phi) = 0. \quad (6)$$

In neglect of the electron inertia, electron motion along the magnetic field, and electron temperature, the electron velocity is a sum of the $\mathbf{E} \times \mathbf{B}$ drifts due to the stationary

electric field $\mathbf{E}_0 = -\nabla \phi_0$, and the perturbed potential $\tilde{\phi}$, with velocity $\tilde{\mathbf{v}}_E = \nabla \tilde{\phi} \times \mathbf{B}$. The resulting density perturbation is

$$\tilde{n}_e = -i \frac{\tilde{\mathbf{v}}_E \cdot \nabla n_0}{(\omega - \omega_0)}. \quad (7)$$

Using the quasineutrality, one gets the eigenvalue equation in slab geometry

$$\frac{\partial^2 \phi}{\partial x^2} - k_y^2 \phi + \frac{n'_0}{n_0} \frac{\partial \phi}{\partial x} - \frac{\omega^2}{(\omega - \omega_0)} \frac{k_y}{\omega_{ci}} \frac{\nabla n_0}{n_0} \phi = 0, \quad (8)$$

where ω is an eigenvalue problem for prescribed boundary conditions on ϕ . In general case, ω and ϕ are complex. This eigenvalue problem is solved with the following boundary conditions: $\tilde{\phi}(a) = \tilde{\phi}(-a) = 0$ or $\tilde{\phi}(a) = \tilde{\phi}(0) = 0$, depending on the geometry type.

It is convenient to use the transformation

$$\phi(x) = \frac{1}{(n_0(x))^{1/2}} \psi(x). \quad (9)$$

Then Equation (8) is reduced to the form

$$\frac{\partial^2 \psi}{\partial x^2} - k_y^2 \psi + \left[-\frac{\omega^2}{(\omega - \omega_0)} \frac{k_y}{\omega_{ci}} \frac{n'_0}{n_0} + \frac{1}{4} \frac{n_0''}{n_0^2} - \frac{1}{2} \frac{n_0''}{n_0} \right] \psi = 0. \quad (10)$$

For the constant gradient profile,

$$n_0(x) = n_0 \exp\left(\frac{x}{L_n}\right), \quad (11)$$

one has the following equation:

$$\frac{\partial^2 \psi}{\partial x^2} - k_y^2 \psi + \left[-\frac{\omega^2}{(\omega - \omega_0)} \frac{k_y}{L_n \omega_{ci}} - \frac{1}{4L_n^2} \right] \psi = 0. \quad (12)$$

For constant L_n and ω_0 , from (12) one gets the following local dispersion relation:

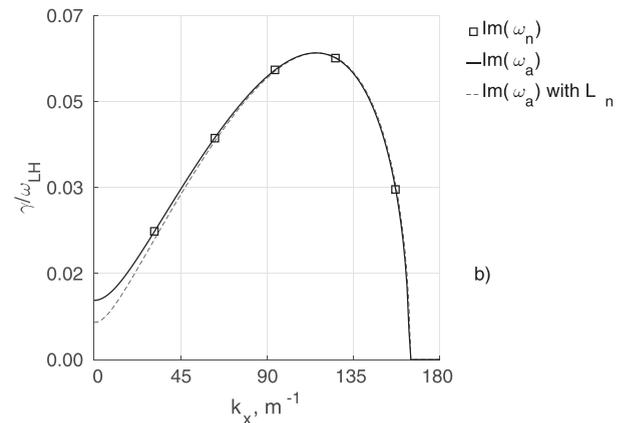
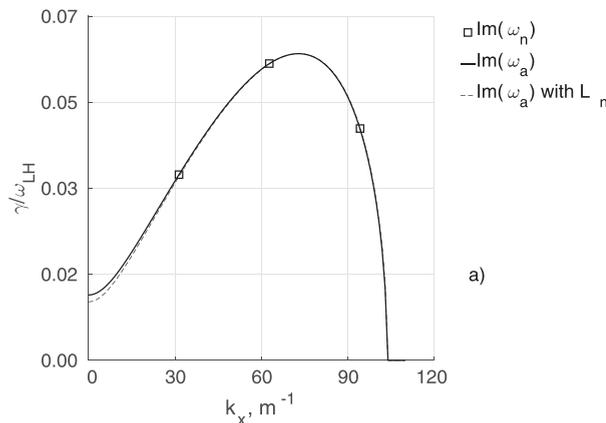


FIG. 1. The growth rate as a function of k_x , and $k_y = 10 \text{ m}^{-1}$; (a) for $L_n = -0.1 \text{ m}$, and (b) for $L_n = -0.04 \text{ m}$. Bold lines—growth rates from Eq. (13); thin dashed lines—growth rates from Eq. (2); squares—growth rates from Eq. (8).

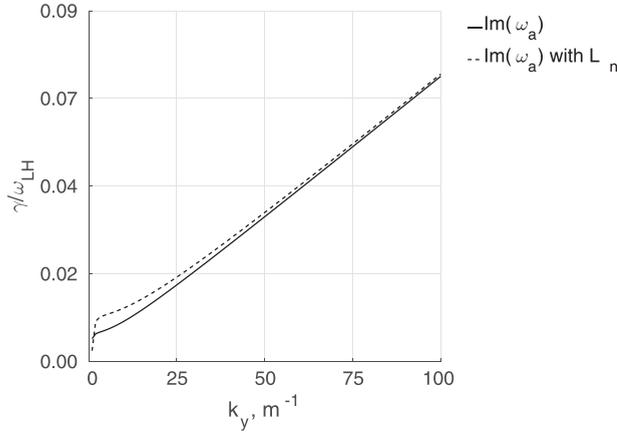


FIG. 2. The growth rate as a function of k_y , for $k_x = 31 \text{ m}^{-1}$ and $L_n = -0.04 \text{ m}$. Bold line—growth rates from Eq. (2); thin dashed line—growth rates from Eq. (13).

$$\left(k_{\perp}^2 + \frac{1}{4L_n^2}\right) \frac{c_s^2}{\omega^2} = \frac{\omega_*}{\omega - \omega_0}. \quad (13)$$

This equation, basically, is a modified Eq. (2). From this equation, one can see that density gradient increases the effective value of k_{\perp} modifying the spectrum of unstable modes. This is illustrated in Fig. 1(a) for $L_n = -0.1 \text{ m}$, and $L_n = -0.04 \text{ m}$ in Fig. 1(b). Profile of the ω_0 is uniform and system size $L = 0.1 \text{ m}$. Bold lines are growth rate calculated from the local Eq. (13) with the effect of L_n . Thin dashed lines indicate the growth rate from Eq. (2). Density gradient effect leads to a slight increase of growth rates at low k_x values. Square marks are the discrete unstable growth rates obtained from numerical solution of Eq. (8) for the system size $L = 0.1 \text{ m}$.

It is important to stress that k_x values are not arbitrary and are defined by the size of the system. Another thing to note is that stronger density gradients (smaller absolute values of L_n) significantly increase a number of unstable modes and range of unstable values of k_x ; however, the maximal value of the growth rate is not changed.

Dependency of growth rate on k_y wavenumber with and without the effect of L_n is shown in Fig. 2. Results were obtained for $L_n = -0.04 \text{ m}$. Consideration of L_n term leads to the constant shift of growth rate. Therefore, for moderate gradient values, this effect does not lead to significant change of the maximum growth rate. One can see from Figs. 1 and 2 that for the considered plasma parameters, the effect of $n'_0/n_0(\partial\phi/\partial x)$ term is not essential.

However, Eq. (8) does not take into account the effect of electron inertia. The important result presented in Fig. 1 is that the maximum of the eigenfrequency is close to the ω_0 frequency. In fact, one can show¹⁷ that for continuous k_x , the eigenmode with the maximum growth rate is $\omega = \omega_0 + i\omega_0$. Therefore, it may become comparable with the lower-hybrid frequency and electron inertia has to be included. The respective equation can be written as in Ref. 17 (similar equation was also given in Refs. 37 and 38)

$$\frac{\partial^2 \phi}{\partial x^2} - k_y^2 \phi - F(\omega) \frac{1}{(\omega - \omega_0)} \frac{k_y}{\omega_{ci}} \frac{\nabla n_0}{n_0} \phi = 0, \quad (14)$$

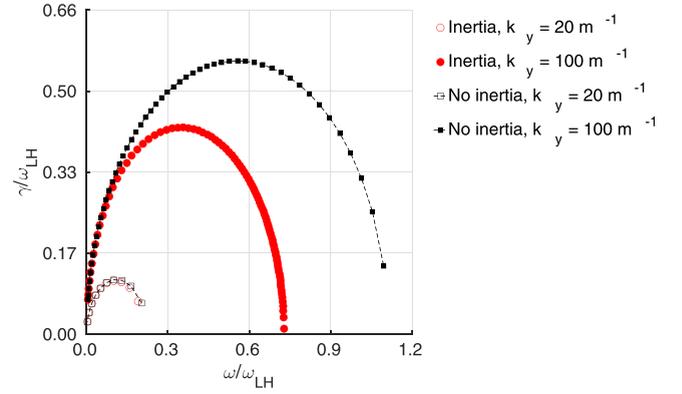


FIG. 3. The full spectrum of unstable solutions with (circles) and without (squares) electron inertia effect for $L_n = -0.04 \text{ m}$: empty markers— $k_y = 20 \text{ m}^{-1}$; filled markers— $k_y = 100 \text{ m}^{-1}$.

where

$$F(\omega) = \frac{\omega^2}{1 - \omega^2/\omega_{LH}^2}.$$

The full spectra of unstable solutions for two values of k_y with and without the effect of electron inertia are shown in Fig. 3. These spectra were obtained by solving Eqs. (8) and (14) for constant ω_0 profile and $L_n = -0.04 \text{ m}$. For low values of k_y (empty circles and squares), unstable frequencies are much lower than ω_{LH} , so inertia effect does not play a significant role; however, for the case of high k_y values (solid circles and squares) electron inertia leads to the significant decrease of unstable frequency and growth rate values.

Another critical effect of the electron inertia is the stabilization of the instability at large values of the k_y . These results are shown in Figs. 4(a) and 4(b). The results for the fixed value of the k_x are shown in Fig. 4(a); $k_x = 31 \text{ m}^{-1}$ is the value chosen as a minimal possible for this system. As was shown above, the value of the k_x at which the growth rate is maximal changes as a function of k_y . The dependence of the maximum growth rate on the k_y is shown in Fig. 4(b). The value of k_x is different at each k_y value.

Dependency on k_x obtained from Equations (8) and (14) is presented in Fig. 5 for two different k_y and L_n values. As before, stronger density gradients lead to the increase in the number of unstable modes. For low k_y values, inertia does not bring significant effect, but for higher k_y values inertia increases the range of unstable modes and causes the appearance of unstable modes with much higher k_x values.

Therefore, electron inertia causes qualitative change in the behavior of unstable solution with the growth of k_y , bringing the cutoff of the instability at high k_x values. Changes in the dependency on k_x are more quantitative, and appear mostly at high k_y values.

III. NONLOCAL EIGENMODES FOR SOME MODEL PROFILES

Here, we analyze the properties of the eigenmodes for different model profiles of $\mathbf{E} \times \mathbf{B}$ velocity. These simple profiles illustrate some characteristic features of the eigenmodes which also persist in more realistic cases.

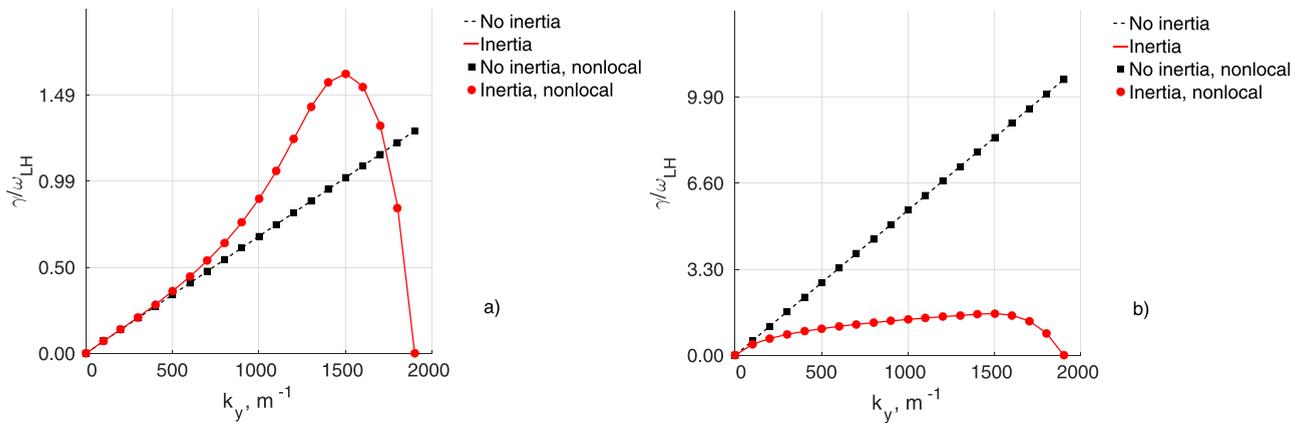


FIG. 4. The growth rate as a function of k_y at $L_n = -0.04$ m. Dashed line—growth rate without electron inertia effect; solid line—growth rate with electron inertia effect; squares and circles—growth rates from nonlocal model without and with electron inertia effect, respectively. (a) Fixed $k_x = 31$ m^{-1} ; (b) maximal growth rate at each k_y value, k_x changes with k_y .

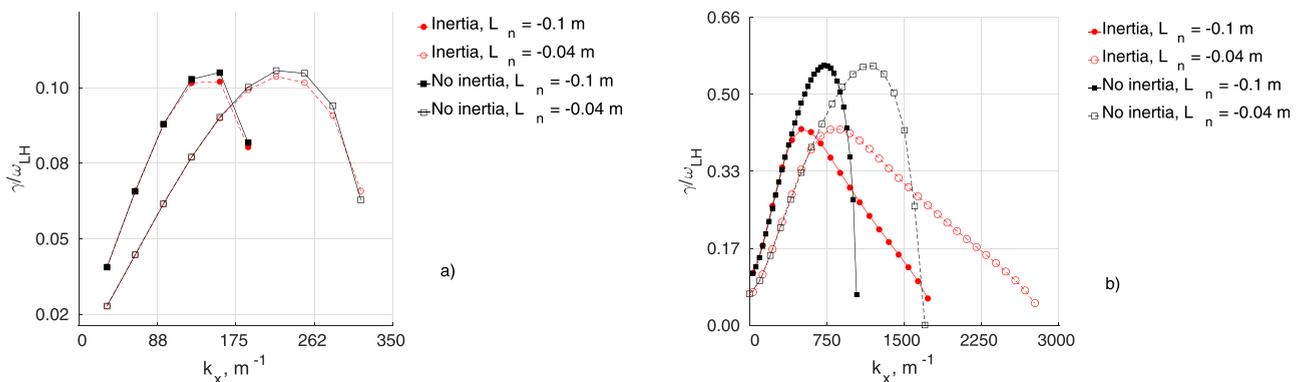


FIG. 5. The growth rate as a function of k_x for $k_y = 20$ m^{-1} (a), and $k_y = 100$ m^{-1} (b). Circles—results with the inertia; squares—results without the inertia. Filled markers—results for $L_n = -0.1$ m; empty markers—results for $L_n = -0.04$ m.

A. Step-like profile of the $\mathbf{E} \times \mathbf{B}$ velocity

Eigenvalue problem (8) with a step-like profile of the $\mathbf{E} \times \mathbf{B}$ drift frequency ω_0 is shown in Fig. 6. Solution inside of each region can be obtained in the form of the harmonic functions, $\phi \sim A_i \cos(\lambda_i x) + B_i \sin(\lambda_i x)$, and the dispersion relation is easily obtained from matching conditions. A formal dispersion relation is not very illuminating. Here, we just emphasize the main features of such solutions and compare the nonlocal modes with the predictions of the local theory.

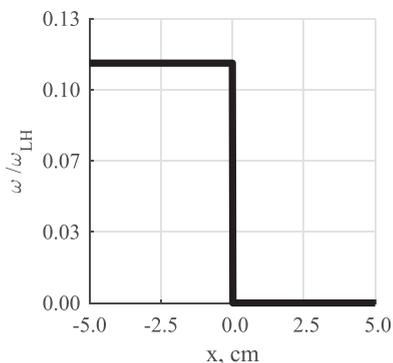


FIG. 6. Profile of ω_0 .

First, we consider a case with the low k_y wavenumber values. As was noted above, the k_x range of the unstable modes is decreasing for lower values of k_y . Since the allowed wavenumber k_x is discrete, the number of unstable eigenmodes decreases with decreasing k_y . The effect of density gradient L_n on nonlocal unstable eigenmodes is similar to the effect of k_y ; thus, the density gradient is fixed and taken $L_n = -0.04$ m for one wavenumber $k_y = 20$ m^{-1} . A characteristic feature of the ω_0 profile, shown in Fig. 6, is that the local theory predicts the instability only in the region $-5 < x < 0$, while the region $0 < x < 5$ should be stable. The nonlocal solutions may exist in the locally stable region.

Full spectra of unstable solutions are shown in Fig. 7(a). Three characteristic modes are chosen to show the shape of eigenfunctions in Fig. 7(b): point (1) with the lowest value of the real part of the frequency $\omega = (0.01 + 0.04i) \cdot \omega_{LH} s^{-1}$; point (2) with the largest growth rate $\omega = (0.11 + 0.1i) \cdot \omega_{LH} s^{-1}$; and point (3) with smallest growth rate $\omega = (0.14 + 0.01i) \cdot \omega_{LH} s^{-1}$. All solutions have the nonlocal structure across the whole domain $-5 < x < 5$ including the region $0 > x > 5$ which is locally stable.

The tendency toward the local theory increases for larger k_y , as is expected. The full spectrum of eigenvalues in phase space for step-like profile of ω_0 with $k_y = 100$ m^{-1} is shown in Fig. 8(a). Some characteristic eigenfunctions are

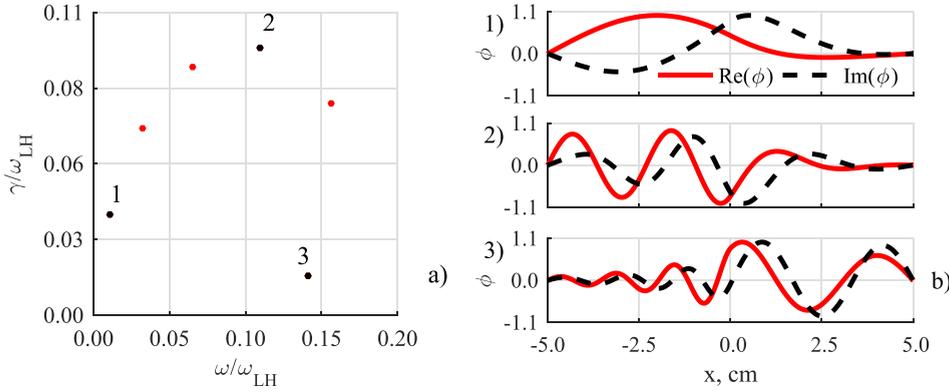


FIG. 7. Results for step-like profile of the ω_0 with $k_y = 20 \text{ m}^{-1}$ and $L_n = -0.04 \text{ m}$. (a) The full spectra of unstable eigenvalues. (b) The eigenfunctions are for eigenvalues (from top to bottom): 1 – $\omega = (0.01 + 0.04i) \cdot \omega_{LH} \text{ s}^{-1}$, 2 – $\omega = (0.11 + 0.1i) \cdot \omega_{LH} \text{ s}^{-1}$, and 3 – $\omega = (0.14 + 0.01i) \cdot \omega_{LH} \text{ s}^{-1}$.

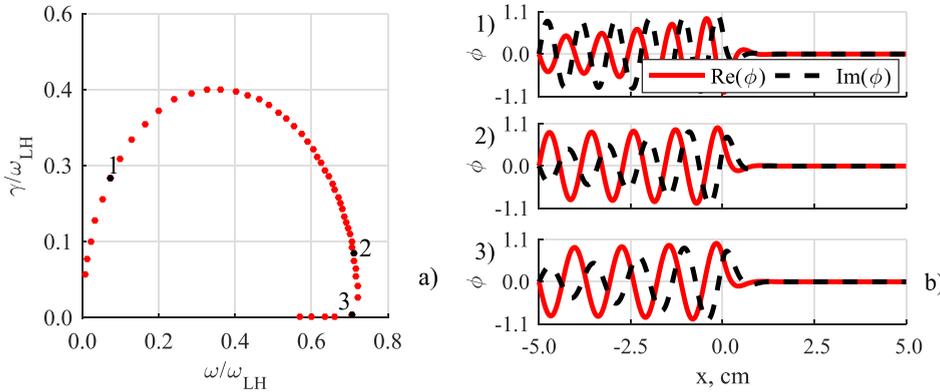


FIG. 8. (a) Full spectrum of the eigenvalues for $k_y = 100 \text{ m}^{-1}$. (b) The eigenfunctions are for eigenvalues (from top to bottom): 1 – $\omega = (0.07 + 0.25i) \cdot \omega_{LH} \text{ s}^{-1}$, 2 – $\omega = (0.71 + 0.12i) \cdot \omega_{LH} \text{ s}^{-1}$, and 3 – $\omega = (0.71 + 0.004i) \cdot \omega_{LH} \text{ s}^{-1}$.

shown in Fig. 8(b); one can see that there are solutions extending into the locally stable region.

General tendency is that the solutions with the lower growth rates are more deeply extended into the locally stable region.

B. Linear profile of the $\mathbf{E} \times \mathbf{B}$ velocity

Here, we investigate the eigenmode problem for the parabolic potential profile $\phi = \alpha x^2$ corresponding to the case of a constant shear of the $\mathbf{E} \times \mathbf{B}$ velocity and exponential density profile $n = n_0 \exp(-x/a)$ which makes the density gradient length scale constant, $L_n^{-1} = n_0^{-1} \partial n_0 / \partial x = \text{const}$. Such profiles may occur in Penning discharge configurations where ions are confined radially by the inward radial electric field.³⁴

We assume that $\omega_0 = k_y v_0 x / L$, where L is the width of the region in the radial direction x . Several unstable

eigenfunctions for $k_y = 20 \text{ m}^{-1}$ and $L = 20 \text{ cm}$ are shown in Fig. 9. The full spectrum of the unstable solutions is presented in Fig. 9(a).

For the $\mathbf{E} \times \mathbf{B}$ velocity with a constant shear, there exists a discrete spectrum of multiple unstable modes. Similar to the local theory results, the ground state is not the most unstable solution. The eigenmode with the largest growth rate for $k_y = 20 \text{ m}^{-1}$ is shown in Fig. 10(b) with respect to the profile of the $\mathbf{E} \times \mathbf{B}$ frequency, $\omega_0(x)$. Lines with circles and squares represent the growth rate obtained from local theory for $k_x = 0$ and $k_x = 63 \text{ m}^{-1}$, respectively. As it is seen, with the increase of k_x value, unstable region shifts to the right.

The eigenfunctions with the largest growth rate are approximately localized close to the resonance point $\text{Re}(\omega) \simeq \omega_0$; however, there is an asymmetrical shift due to the finite growth rate. Similar to the local theory, the growth rate increases with the effective value of the radial wave number k_x .

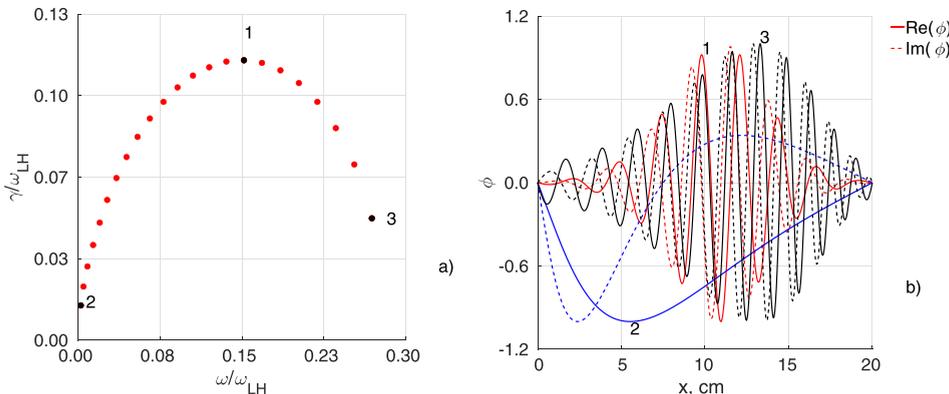


FIG. 9. (a) Full spectrum of the unstable eigenvalues; (b) unstable eigenfunctions for the constant $\mathbf{E} \times \mathbf{B}$ shear profile; wavenumber $k_y = 20 \text{ m}^{-1}$. The eigenfunction with the largest growth rate, $\omega = (0.15 + 0.11i) \cdot \omega_{LH} \text{ s}^{-1} - 1$ (red); the ground state unstable eigenfunction with $\omega = (0.002 + 0.014i) \cdot \omega_{LH} \text{ s}^{-1} - 2$ (blue); the eigenfunction with largest real frequency $\omega = (0.27 + 0.05i) \cdot \omega_{LH} \text{ s}^{-1} - 3$ (black).

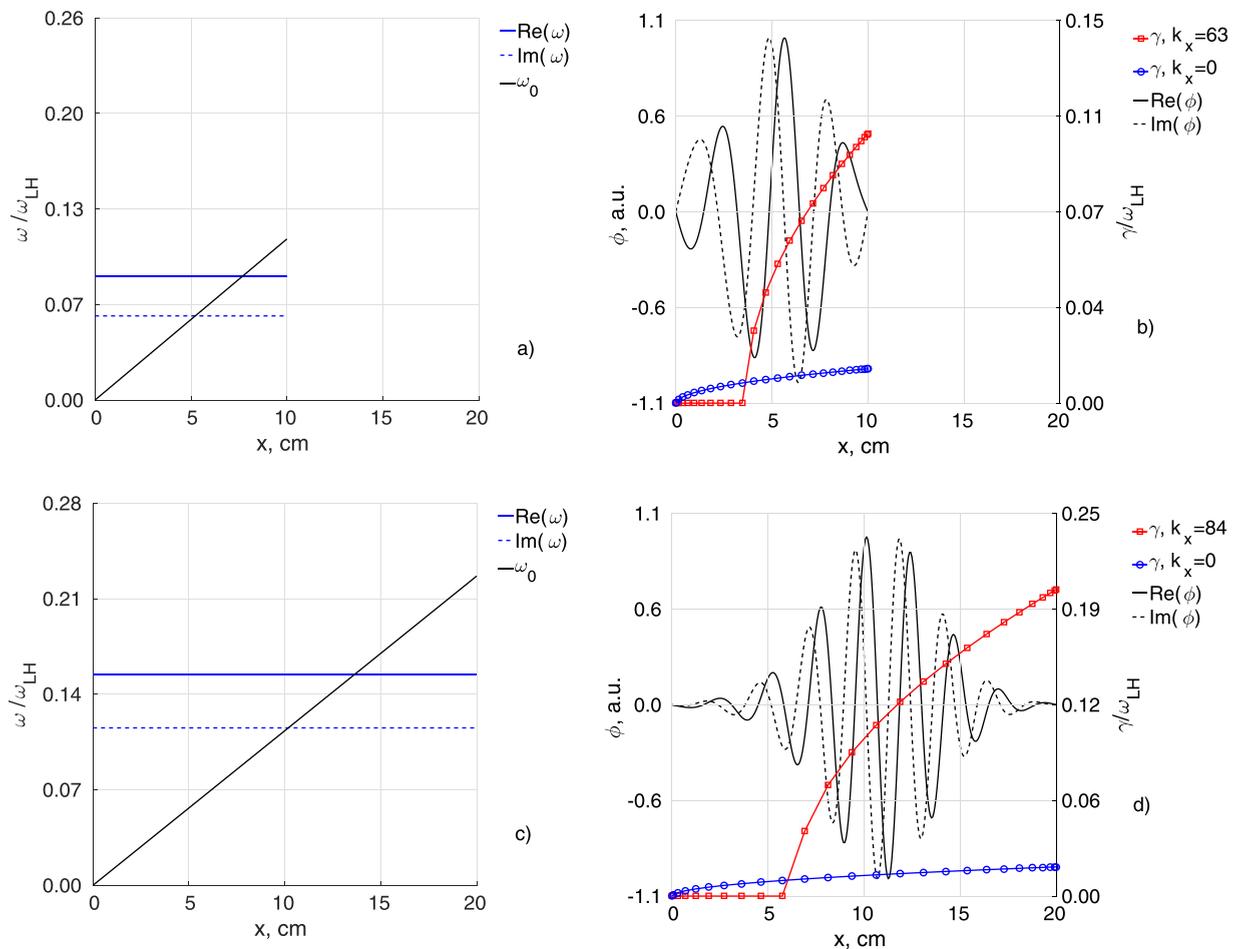


FIG. 10. Left side: the eigenfunction with the largest growth rate $\omega = (0.09 + 0.6i) \cdot \omega_{LH} s^{-1}$ (b), with respect to the profile of the $\mathbf{E} \times \mathbf{B}$ frequency (a), $k_y = 20 m^{-1}$ and $L = 10 cm$. The growth rates from the local theory for the same parameters and $k_x = 0$ —squares (red), $k_x = 63 m^{-1}$ —circles (blue). Right side: the eigenfunction with the largest growth rate $\omega = (0.07 + 0.06i) \cdot \omega_{LH} s^{-1}$ (d) in the extended domain: $L = 20 cm$, for $k_y = 20 m^{-1}$, with respect to the $\mathbf{E} \times \mathbf{B}$ frequency profile (c). The growth rates from the local theory for the same parameters and $k_x = 0$ —squares (red), $k_x = 84 m^{-1}$ —circles (blue).

This is illustrated by the eigenfunction for the same value of $k_y = 20 m^{-1}$ as in Fig. 10(b) but for the extended domain with the length $L = 20 cm$, see Fig. 10(d). In the extended domain, the most unstable eigenfunction has larger growth rate, higher effective k_x , and the localization region shifts to the right (toward higher local ω_0), compare Figs. 10(b) and

10(d). Again, there is very little resemblance between the results of the local theory and the nonlocal solution. The local growth rates for comparison are shown in Figs. 10(b) and 10(d) for two different values of the radial wave number k_x . The local theory predicts that the mode with $k_x = 0$ is unstable in the wide region. For the $k_x = 84 m^{-1}$, the

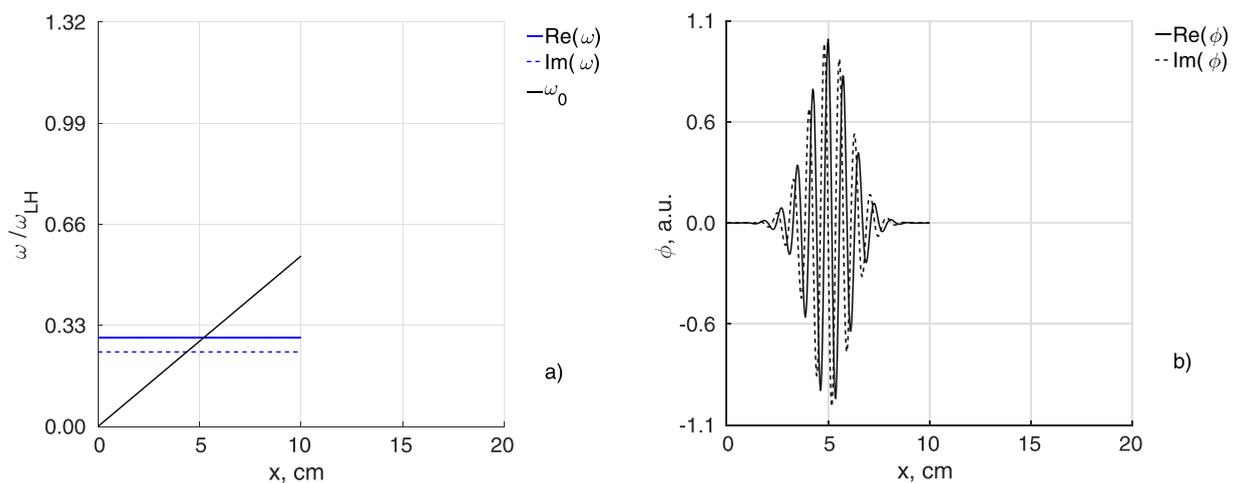


FIG. 11. The eigenfunction with largest growth rate $\omega = (0.29 + 0.24i) \cdot \omega_{LH} s^{-1}$, $k_y = 100 m^{-1}$, $L = 10 cm$.

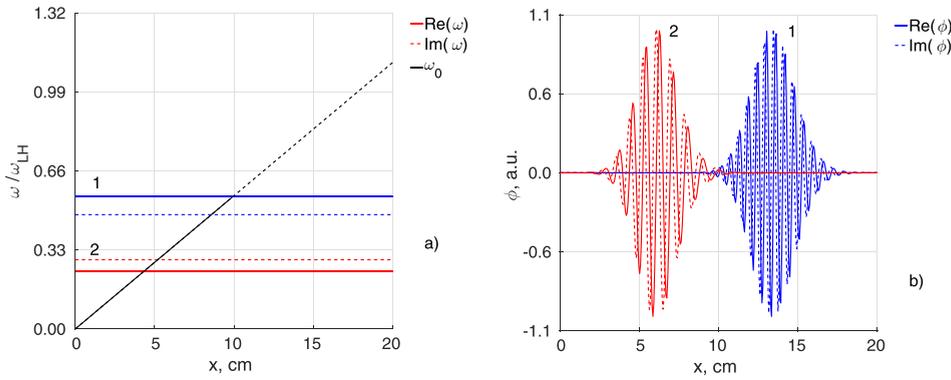


FIG. 12. Several eigenfunctions are shown for the extended domain $L = 20$ cm, $k_y = 100$ m $^{-1}$. The eigenfunction with the largest growth rate for the extended domain, $\omega = (0.56 + 0.48i) \cdot \omega_{LH} \text{ s}^{-1} - 1$ (blue); $\omega = (0.24 + 0.29i) \cdot \omega_{LH} \text{ s}^{-1} - 2$ (red). The eigenfunction with the eigenvalue $\omega = (0.24 + 0.29i) \cdot \omega_{LH} \text{ s}^{-1} - 2$ (red) has approximately the same frequency and localization as mode on with Fig. 11, for $L = 10$ cm.

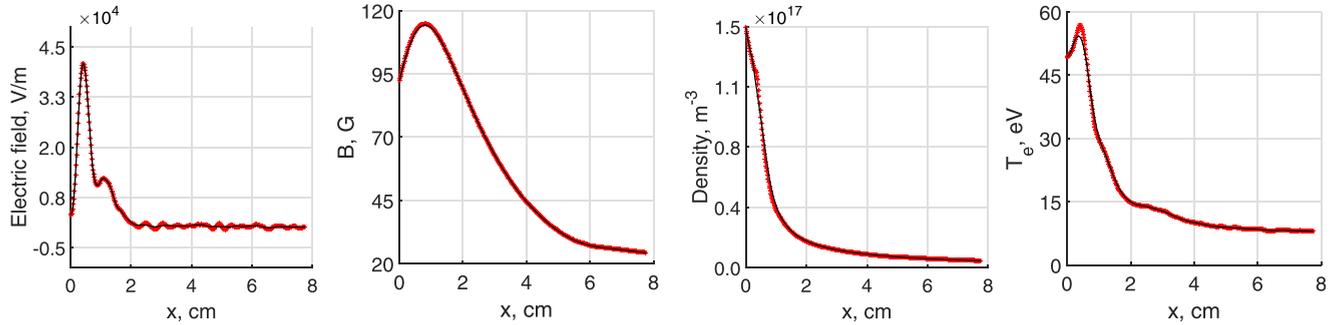


FIG. 13. Initial profiles of electric, magnetic fields, electron density, and electron temperature for Hall thruster experiment.³⁹

instability exists locally only in the region for $x > 6$ cm. In fact, the nonlocal solution has the effective k_x which is higher than $k_x = 84$ m $^{-1}$ and for which the local theory predicts no instability in the whole domain $0 < x < 20$ cm.

The growth rate increases almost linearly with the wave-number k_y , and the localization region becomes narrower around the resonance point $Re(\omega) \simeq \omega_0$. The most unstable eigenfunction for $k_y = 100$ m $^{-1}$ and system length $L = 10$ cm is shown in Fig. 11(b) with respect to the profile of the $\omega_0(x)$.

Several unstable eigenfunctions in the extended domain $L = 20$ cm with the same value of $k_y = 100$ m $^{-1}$ are shown in Fig. 12. One can see that these modes are relatively local in the sense that they are weakly dependent on the boundary conditions at $x = 0$ and $x = L$. The eigenfunction with $\omega = (0.29 + 0.24i) \cdot \omega_{LH} \text{ s}^{-1}$ and localized at $x \simeq 5$ cm, shown in Fig. 11, corresponds to almost the same frequency $Re(\omega)$ and localized at the same $x \simeq 5$ cm as mode (2) which is shown in Fig. 12(b). However, localization of mode (1) is far from resonant point for this eigenmode, see line (1) in Fig. 12. This is due to the effect of the electron inertia.

Therefore, another property of nonlocal modes can be understood from the example with the linear profile of ω_0 . For unstable eigenmodes with absolute values of frequencies and growth rates well below of the ω_{LH} , the localization region is approximately determined by the point of the resonance $Re(\omega) = \omega_0$. For higher frequencies this is not true, in general.

IV. NONLOCAL EIGENMODES IN PLASMAS WITH TEMPERATURE, DENSITY, AND MAGNETIC FIELD PROFILES

Inhomogeneous magnetic field plays an important role in operation of Hall thrusters^{4,5} and magnetrons.¹⁻³

Modification of the instability criteria due to the magnetic field was noted in Refs. 4 and 5. An additional effect of the magnetic field gradient occurs due to the finite temperature.^{15,18} The appropriate eigenvalue equation³⁶ is

$$\frac{\partial^2 \phi}{\partial x^2} - k_y^2 \phi + \frac{1}{1 - \omega^2/\omega_{LH}^2} \frac{\omega^2(\omega_* - \omega_D)}{(\omega - \omega_0 - \omega_D)} \frac{1}{c_s^2} \phi = 0. \quad (15)$$

For simulations, we use typical parameters from experiments in a Hall thruster.^{39,40} The density n_0 , electron

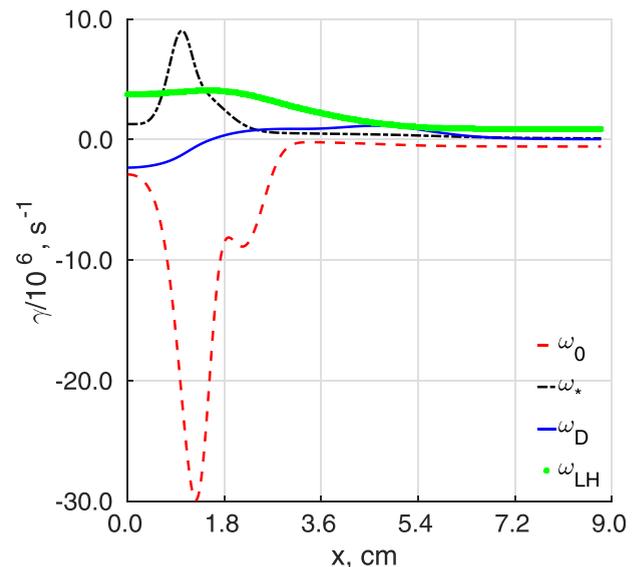


FIG. 14. Characteristic frequencies profiles and unstable solutions for $k_y = 8.1$ m $^{-1}$.

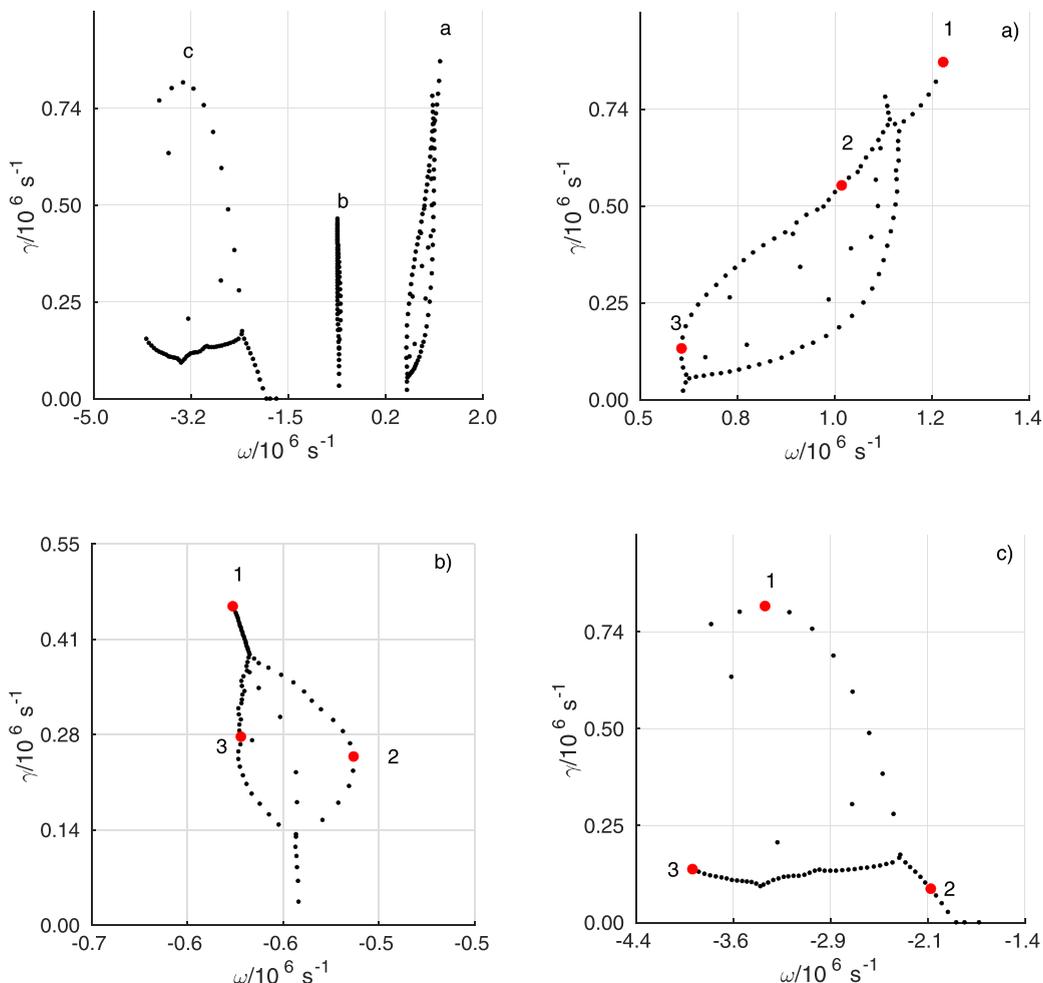


FIG. 15. Full spectra of unstable solutions for Hall thruster profiles at $k_y = 8.1 \text{ m}^{-1}$. Three groups of unstable solutions are marked as (a)–(c). Eigenvalues for numbered dots are presented in Table I.

temperature T_e , electric field E , and magnetic field B profiles along the thruster axis (x -direction) are shown in Fig. 13.

The profiles of the drift frequencies ω_0 , ω_* , ω_D , and ω_{LH} are shown in Fig. 14. Calculations were done for parameters from Fig. 13 and $k_y = 8.1 \text{ m}^{-1}$. For the complex profiles of plasma parameter, the nonlocal solution in general requires high number of polynomials in the spectral method. To verify the solutions obtained with spectral and shooting methods, and confirm the convergence, we have also used the integral relations that follow from Eq. (15)

TABLE I. Eigenvalues for Fig. 15.

Fig. 15(a)	1	$\omega = (1.2 + 0.9) \times 10^6 \text{ s}^{-1}$
	2	$\omega = (1.0 + 0.6) \times 10^6 \text{ s}^{-1}$
	3	$\omega = (0.6 + 0.1) \times 10^6 \text{ s}^{-1}$
Fig. 15(b)	1	$\omega = (-0.6 + 0.5) \times 10^6 \text{ s}^{-1}$
	2	$\omega = (-0.5 + 0.2) \times 10^6 \text{ s}^{-1}$
	3	$\omega = (-0.6 + 0.3) \times 10^6 \text{ s}^{-1}$
Fig. 15(c)	1	$\omega = (-3.4 + 0.8) \times 10^6 \text{ s}^{-1}$
	2	$\omega = (-2.1 + 0.9) \times 10^6 \text{ s}^{-1}$
	3	$\omega = (-4.0 + 0.1) \times 10^6 \text{ s}^{-1}$

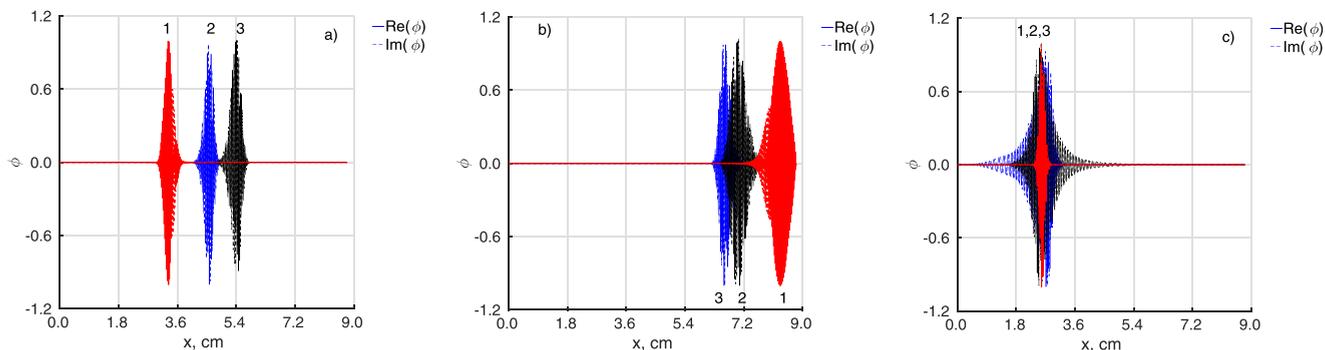


FIG. 16. Unstable eigenfunctions from each group of unstable solutions at $k_y = 8.1 \text{ m}^{-1}$.

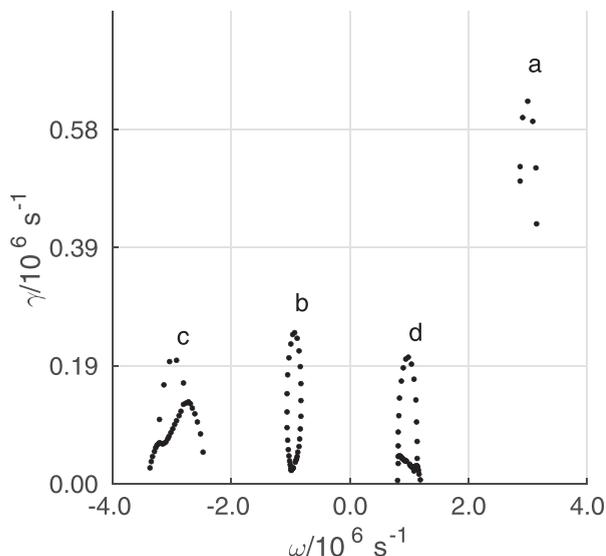


FIG. 17. Full spectra of unstable solutions for Hall thruster profiles at $k_y = 81 \text{ m}^{-1}$. Four distinct groups of the unstable solutions can be identified.

$$\int (|\phi|^2 + k_y^2 |\phi|^2) dx = \int \text{Re} \left(\frac{1}{1 - \omega^2/\omega_{LH}^2} \frac{\omega^2(\omega_* - \omega_D)}{(\omega - \omega_0 - \omega_D)} \right) \frac{1}{c_s^2} |\phi|^2 dx \quad (16)$$

and

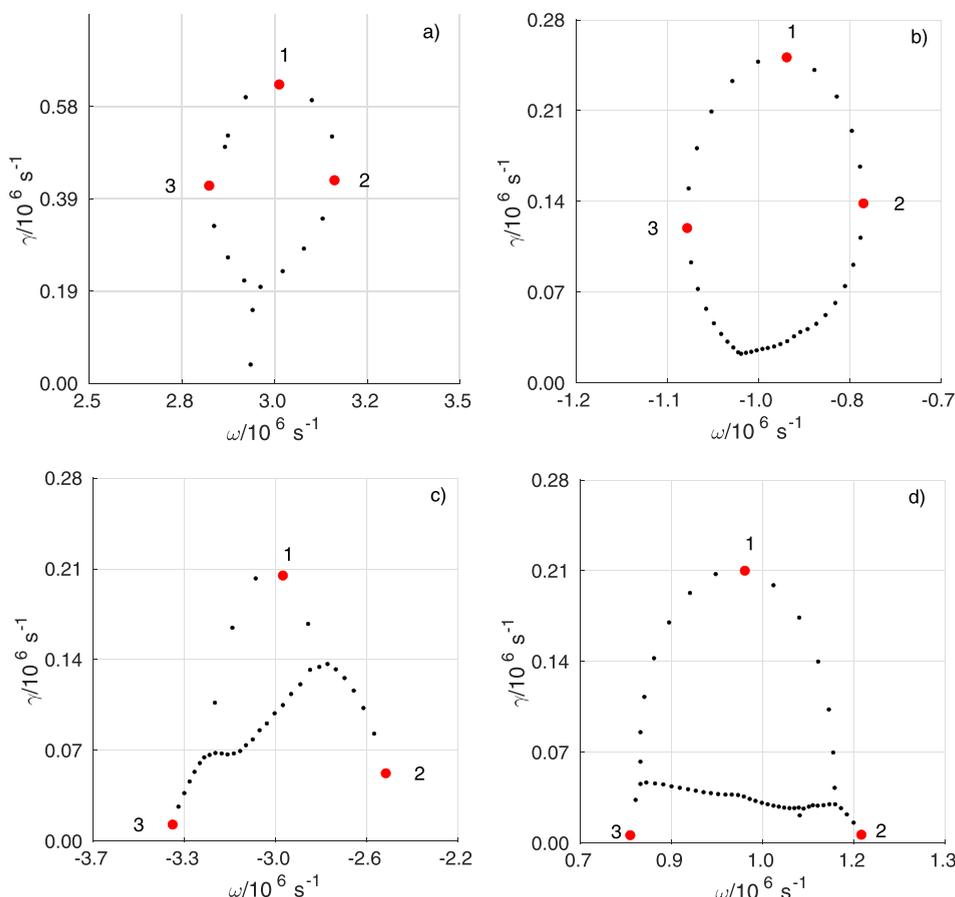


FIG. 18. Each group of unstable solution from Fig. 17. Eigenvalues for numbered dots are presented in Table II.

$$0 = \int \text{Im} \left(\frac{1}{1 - \omega^2/\omega_{LH}^2} \frac{\omega^2(\omega_* - \omega_D)}{(\omega - \omega_0 - \omega_D)} \right) \frac{1}{c_s^2} |\phi|^2 dx. \quad (17)$$

The ratio of left hand side and right hand side integrals was tracked with the increase in the number of polynomials. Desired accuracy was 10^{-5} .

Similar to general trend, for a given k_y , there exist multiple eigenmodes with different growth rates and different localization regions. However, full spectra of unstable solutions look very different from spectra for simple profiles. An example of such spectra for $k_y = 8.1 \text{ m}^{-1}$ is presented in Fig. 15. Three distinct groups of unstable solutions can be identified.

An important difference is the sign of the real part of the frequency. There exist unstable modes with positive and negative frequencies. Note that the negative sign of the real part of the frequency correspond to the rotation in the $\mathbf{E} \times \mathbf{B}$ direction. Corresponding eigenvalues and eigenfunctions for the numbered point in each group are given in Table I and Fig. 16, respectively.

Unstable eigenfunctions for marked points are shown in Fig. 16. Eigenfunctions, corresponding to different groups, have different localization. Contrary to the case of simple profiles, narrowly localized modes appear even for low k_y value. Modes in group (c) have almost exactly the same localization, thus local theory should work very well in this region. However, modes in group (a) are located in the region, where the local theory does not predict any instability (see in Fig. 20).

TABLE II. Eigenvalues for Fig. 18.

Fig. 19(a)	1	$\omega = (3.0 + 0.6) \times 10^6 \text{ s}^{-1}$
	2	$\omega = (3.1 + 0.4) \times 10^6 \text{ s}^{-1}$
	3	$\omega = (2.8 + 0.4) \times 10^6 \text{ s}^{-1}$
Fig. 19(b)	1	$\omega = (-0.9 + 0.23) \times 10^6 \text{ s}^{-1}$
	2	$\omega = (-0.8 + 0.13) \times 10^6 \text{ s}^{-1}$
	3	$\omega = (-1.0 + 0.1) \times 10^6 \text{ s}^{-1}$
Fig. 19(c)	1	$\omega = (-2.9 + 0.14) \times 10^6 \text{ s}^{-1}$
	2	$\omega = (-2.4 + 0.03) \times 10^6 \text{ s}^{-1}$
	3	$\omega = (-3.4 + 0.02) \times 10^6 \text{ s}^{-1}$
Fig. 19(d)	1	$\omega = (1.2 + 0.01) \times 10^6 \text{ s}^{-1}$
	2	$\omega = (1.0 + 0.17) \times 10^6 \text{ s}^{-1}$
	3	$\omega = (0.8 + 0.02) \times 10^6 \text{ s}^{-1}$

For higher value of k_y , full spectra of unstable solutions is modified, there appeared fourth group of unstable solutions, see Fig. 17. Group (d) appeared in between of groups (a) and (b) from Fig. 15.

A detailed view of each group is presented in Fig. 18. Corresponding eigenvalues and eigenfunctions for the numbered point in each group are given in Table II and Fig. 19, respectively.

Unstable eigenfunctions are presented in Fig. 19. Modes for groups (a) and (b) become local, as it is expected for higher k_y values. However, unexpected result is that the modes in groups (c) and (d) become highly nonlocal.

Absolute values of frequencies are comparable in these two groups, but growth rates are lower in the region farther from the anode (group (d)).

Moreover, the real and imaginary parts of the most unstable eigenvalue almost do not change with the k_y wavenumber.

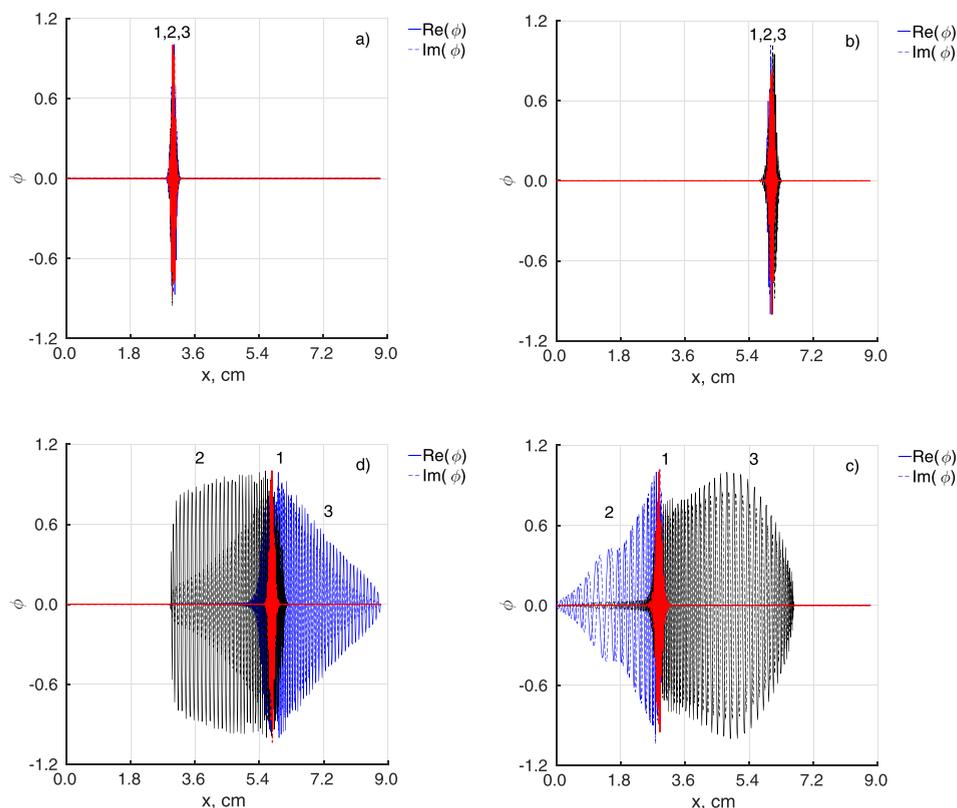


FIG. 19. Unstable eigenfunctions from each group of unstable solutions at $k_y = 81 \text{ m}^{-1}$.

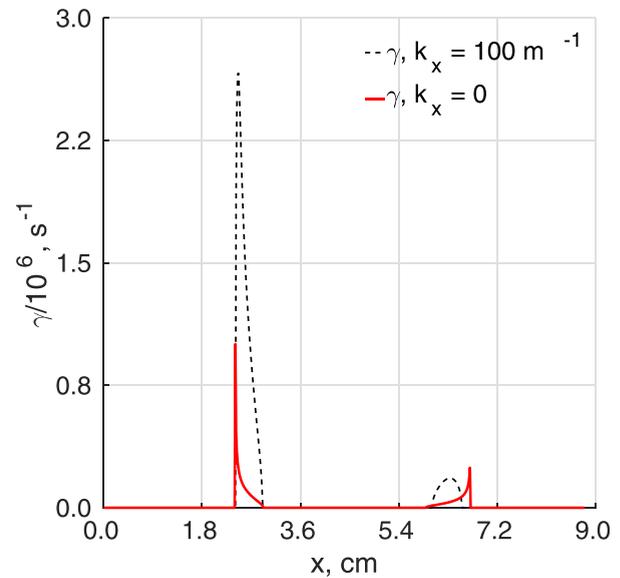


FIG. 20. The growth rate from local model are shown as a function of distance along the thruster axis for $k_y = 8.1 \text{ m}^{-1}$, and $k_x = 0 \text{ m}^{-1}$, $k_x = 100 \text{ m}^{-1}$.

Growth rate of the unstable mode mostly depend on mode localization. Real part of frequencies for fastest modes is comparable with low-hybrid frequency $\omega_{LH} = 3.4 \times 10^6 \text{ s}^{-1}$.

The nonlocal modes (with the least number of nodes) have the lower growth rates compared to the localized (with the higher effective higher k_x) modes. The nonlocal unstable modes, presented in Figs. 16 and 19, are present in the regions where local modes are stable (shown in Fig. 20). It is important to note that the fast instabilities are absent in the acceleration region where the electric field is large.

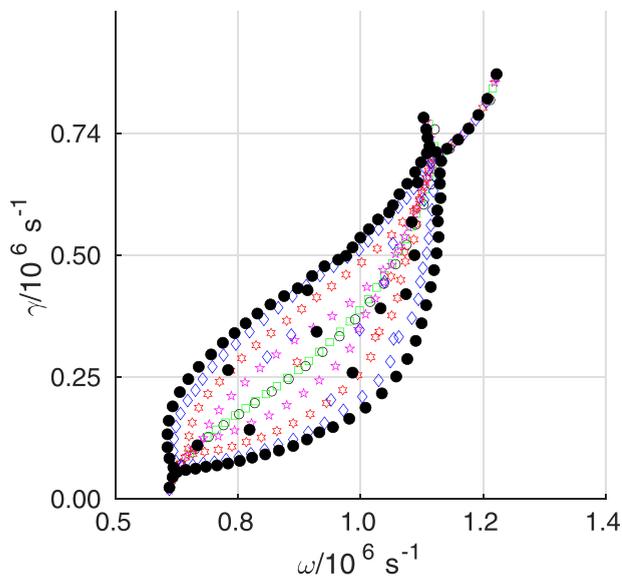


FIG. 21. Convergence of group (b) in Fig. 15 with the increase in the number of polynomials. Filled black circles—the stage when convergence is reached.

The integral forms, Eqs. (16) and (17), were used to track the convergence. An example of the convergence for one group is shown in Fig. 21. Each marker and color represents different number of polynomials. Green markers are solutions for low polynomials number. With the increase in the number of polynomials, eigenvalues reach asymptotically the limit values which are shown as solid black circles.

V. CONCLUSION

Linear, nonlocal model for gradient-drift modes including the collisionless Simon-Hoh and lower-hybrid instabilities was developed and used to study the nonlocal structure of unstable mode in Hall plasma with inhomogeneous profiles of plasma density, temperature, electric and magnetic fields. The eigenvalue problem was solved numerically by using spectral and shooting methods and verified by the integral relations.

Our model includes the collisionless Simon-Hoh destabilization mechanism¹³ $\mathbf{E} \cdot \nabla n > 0$ and electron inertia that couples this mode to the lower-hybrid mode. A number of factors were not included in the model, such as collisional,⁴¹ ionization²³ effects, and additional terms that explicitly include the shear flow effects.^{17,41} In general, neglect of these effects may limit the applicability of our model to some practical Hall devices. Our emphasis, however, was on the higher frequency modes with frequencies much higher than some typical low frequency (kHz range) oscillations, such as breathing mode due to ionization process. We have considered the eigenmode structure in the axial direction of the Hall thruster, and the radial direction for the Penning discharge. One of the important results of our analysis is demonstration of the existence of multiple modes with the eigenmode frequencies comparable to the electron drift frequency.⁴²

There are significant differences between predictions of the local and nonlocal models, as it has been noted

earlier.^{23,30} The nonlocal model for the step-like ω_0 profiles shows that for the low wavenumbers k_y nonlocal solutions can propagate into the region of the domain where local instability criteria are not satisfied. For higher k_y modes, the results of nonlocal model become similar to the local model where the unstable mode is highly localized. However, there are still unstable modes that extend into the region of stability and grow there even for high k_y values.

Simulations for the $\mathbf{E} \times \mathbf{B}$ flow with a constant shear and constant density gradient show that there exist multiple unstable eigenmodes with different growth rates and different localization regions. Mode localization becomes more evident with increase of k_y . Moreover, such modes are almost independent on boundary conditions at $x=0$ and $x=L$.

We have also studied the eigenmode structure for complex profiles of the magnetic field, electric field, and electron temperature relevant to Hall thruster experiments.³⁹ In general, the results demonstrate characteristic features similar to the outlined above. There exist multiple eigenmodes with different growth rates and different localization regions. The most unstable modes tend to localize in the region with higher gradients; however, modes with lower growth rates are present in the whole domain. There are significant differences between the predictions of the local and nonlocal models. This discrepancy is especially important for the modes with low $k_y(m)$ and low effective k_x (longer wavelength).

Often, the modes with largest growth rate are thought to dominate the dynamics on the longer time scales. However, in case of complex profiles, different eigenmodes may be localized in different spatial regions, even in regions which are locally stable. For example, in case of Hall thruster profiles there are modes which reside at the middle of the domain, where the local model predicts no instability. Some modes propagate from locally unstable regions to locally stable. The anomalous current in these conditions is due to the electron convection in the direction of the equilibrium electric field, $\sim \langle \tilde{n} \tilde{V}_x \rangle$.⁴³ Then one can estimate the diffusion transport using a mixing length formula, $D_a \simeq \gamma/k_y^2$, where γ is the mode growth rate, k_y is the azimuthal wave number. Hence, significance of multiple modes is especially important for low k_y modes which have the nonlocal character and whose stability criteria could be very different from the predictions of the local theory and which provide larger contribution to the anomalous transport ($D_a \simeq \gamma/k_y^2$). These features of nonlocal solutions will be important for nonlinear transport calculations. In view of different stability criteria for multiple modes, especially for low k_y values, the predictions of the local theory may be misleading.

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- ¹P. Kelly and R. Arnell, *Vacuum* **56**, 159 (2000).
- ²T. Ito, C. V. Young, and M. A. Cappelli, *Appl. Phys. Lett.* **106**, 254104 (2015).
- ³S. N. Abolmasov, *Plasma Sources Sci. Technol.* **21**, 035006 (2012).
- ⁴A. I. Morozov, *Plasma Phys. Rep.* **29**, 235 (2001).
- ⁵A. I. Morozov and V. V. Savelyev, "Fundamentals of stationary plasma thruster theory," in *Reviews of Plasma Physics* (Springer US, Boston, MA, 2000), pp. 203–391.
- ⁶A. I. Morozov, Y. V. Esipchuk, A. M. Kapulkin, V. A. Nevrovskii, and V. A. Smirnov, *Sov. Phys. Tech. Phys.* **17**, 482 (1972).
- ⁷Y. V. Esipchuk, A. I. Morozov, G. N. Tilinin, and A. V. Trofimov, *Sov. Phys. Tech. Phys.* **43**, 1466 (1973).
- ⁸A. M. Fridman, *Sov. Phys. Dokl.* **9**, 75 (1964).
- ⁹A. I. Smolyakov, W. Frias, I. D. Kaganovich, and Y. Raitses, *Phys. Rev. Lett.* **111**, 115002 (2013).
- ¹⁰A. Simon, *Phys. Fluids* **6**, 382 (1963).
- ¹¹F. C. Hoh, *Phys. Fluids* **6**, 1184 (1963).
- ¹²Y. Q. Tao, R. W. Conn, L. Schmitz, and G. Tynan, *Phys. Plasmas* **1**, 3193 (1994).
- ¹³Y. Sakawa, C. Joshi, P. K. Kaw, F. F. Chen, and V. K. Jain, *Phys. Fluids B* **5**, 1681 (1993).
- ¹⁴L. A. Artsimovich, I. M. Andronov, Y. V. Esipchuk, I. A. Bersukov, and K. N. Kozubskii, *Kosm. Issled.* **12**, 451 (1974).
- ¹⁵W. Frias, A. Smolyakov, I. D. Kaganovich, and Y. Raitses, *Phys. Plasmas* **20**, 052108 (2013).
- ¹⁶W. Frias, A. Smolyakov, I. D. Kaganovich, and Y. Raitses, *Phys. Plasmas* **21**, 062113 (2014).
- ¹⁷A. Smolyakov, O. Chapurin, W. Frias, O. Koshkarov, I. Romadanov, T. Tang, M. Umansky, Y. Raitses, I. Kaganovich, and V. Lakhin, *Plasma Phys. Controlled Fusion* **59**, 1, 14041 (2016); see <http://stacks.iop.org/0741-3335/59/i=1/a=014041>.
- ¹⁸A. Kapulkin and M. M. Guelman, *IEEE Trans. Plasma Sci.* **36**, 2082 (2008).
- ¹⁹J. P. Boeuf and B. Chaudhury, *Phys. Rev. Lett.* **111**, 847 (2013).
- ²⁰M. J. Seakerak, B. W. Longmier, A. D. Gallimore, D. L. Brown, R. R. Hofer, and J. E. Polk, *IEEE Trans. Plasma Sci.* **43**, 72 (2015).
- ²¹C. L. Ellison, Y. Raitses, and N. J. Fisch, *Phys. Plasmas* **19**, 013503 (2012).
- ²²A. Kapulkin and M. Guelman, "Lower-hybrid instability in Hall thruster," in Proceedings of the 29th International Electric Propulsion Conference (Electric Rocket Propulsion Society, 2005).
- ²³D. Escobar and E. Ahedo, *Phys. Plasmas* **22**, 102114 (2015).
- ²⁴Y. V. Esipchuk and G. N. Tilinin, *Sov. Phys. Tech. Phys.* **21**, 417 (1976).
- ²⁵A. Kapulkin and M. Guelman, "Low frequency instability and enhanced transfer of electrons in near-anode region of Hall thruster," in Proceedings of the 30th International Electric Propulsion Conference (Electric Rocket Propulsion Society, 2007).
- ²⁶B. A. Jorns and R. R. Hofer, *Phys. Plasmas* **21**, 053512 (2014).
- ²⁷R. F. Ellis, E. Marden-Marshall, and R. Majeski, *Plasma Phys.* **22**, 113 (1980).
- ²⁸F. F. Chen, *Phys. Fluids* **10**, 1647 (1967).
- ²⁹S. Barral, Y. Jayet, S. Mazouffre, M. Dudeck, E. Vron, and P. Echegut, "Hall effect thruster with an AlN discharge channel," in Proceedings of the 29th International Electric Propulsion Conference (Electric Rocket Propulsion Society, 2005).
- ³⁰D. Escobar and E. Ahedo, *IEEE Trans. Plasma Sci.* **43**, 149 (2015).
- ³¹Y. Raitses, P. Baele, and V. M. Donnelly, in Proceedings of the 64th Gaseous Electronics Conference (2011).
- ³²A. Smirnov, Y. Raitses, and N. J. Fisch, *J. Appl. Phys.* **95**, 2283 (2004).
- ³³Y. Raitses, J. B. Parker, E. Davis, and N. J. Fisch, "Background gas pressure effects in the cylindrical Hall thruster," in Proceedings of the 46th AIAA/ASME/SAE/ASEE Joint Propulsion Conference and Exhibit (American Institute of Aeronautics and Astronautics, 2010), Paper No. AIAA-2010-6775.
- ³⁴Y. Raitses, I. Kaganovich, and A. Smolyakov, "Effects of the gas pressure on low frequency oscillations in $E \times B$ discharges," in Proceedings of the Joint Conference of 30th ISTS, 34th IEPC, and 6th NSAT (2015).
- ³⁵L. N. Trefethen, *Spectral Methods in MATLAB* (Society for Industrial and Applied Mathematics, Philadelphia, PA, USA, 2000).
- ³⁶W. Frias, A. I. Smolyakov, I. D. Kaganovich, and Y. Raitses, *Phys. Plasmas* **19**, 072112 (2012).
- ³⁷A. M. DuBois, T. Edward, W. Amatucci, and G. Ganguli, *Phys. Plasmas* **21**, 062117 (2014).
- ³⁸G. Ganguli, Y. C. Lee, and P. J. Palmadesso, *Phys. Plasmas* **31**, 823 (1988).
- ³⁹Y. Raitses, D. Staack, M. Keidar, and N. J. Fisch, *Phys. Plasmas* **12**, 057104 (2005).
- ⁴⁰A. A. Litvak and N. J. Fisch, *Phys. Plasmas* **8**, 648 (2001).
- ⁴¹A. A. Litvak and N. J. Fisch, *Phys. Plasmas* **11**, 1379 (2004).
- ⁴²A. Lazurenko, V. Krasnoselskikh, and A. Bouchoule, *IEEE Trans. Plasma Sci.* **36**, 1977 (2008).
- ⁴³R. Spektor, "Quasi-linear analysis of anomalous electron mobility inside a Hall thruster," in Proceedings of the 30th International Electric Propulsion Conference (Electric Rocket Propulsion Society, 2009).