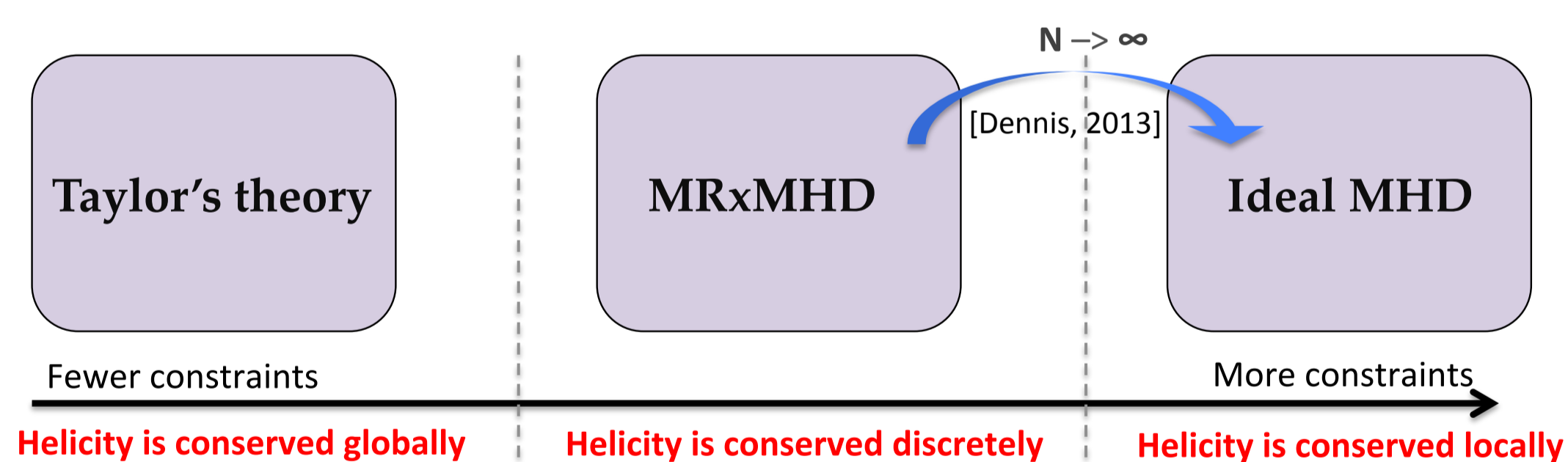


## 1. Motivation

- **Verification & Validation** are the milestones in the path towards predictive code capability [1].
  - Verification answers: *Are we solving the equations right?*
  - Validation answers: *Are we solving the right equations?*
- Fusion research years for **fast, robust, and reliable** codes describing 3D MHD equilibria.
- Challenge: intricate combination of magnetic **surfaces**, magnetic **islands**, and **chaos** [2].
- The Stepped Pressure Equilibrium Code (SPEC) was developed to fulfil this nontrivial task [3].
- SPEC was verified in axisymmetry and for slightly perturbed configurations [3-6].
- Here we present the first SPEC calculations of equilibria in stellarator geometries.

## 2. The SPEC code

- SPEC finds equilibria as extrema of the **Multiregion Relaxed MHD** energy functional [7].
- In MRxMHD, discrete topological constraints allow for **partial relaxation**:
  - Plasma is partitioned into  $N$  nested volumes,  $V_v$ , undergoing Taylor relaxation.
  - Volumes separated by  $N - 1$  interfaces,  $\partial V_v$ , constrained to remain magnetic surfaces.
  - Location and shape of surfaces determined self-consistently by force-balance condition.



- MRxMHD **equilibrium states satisfy**, for  $v = 1, \dots, N$ :

$$\begin{aligned} \nabla \times \mathbf{B} &= \mu_v \mathbf{B} \quad \text{in } V_v \\ \left[ \left[ p_v + \frac{B^2}{2} \right] \right]_v &= 0 \quad \text{in } \partial V_v \end{aligned}$$

- SPEC is a **fixed-boundary** code and requires specification of the boundary:

$$R = \sum_{mn} R_{mn} \cos(m\theta - n\varphi) \quad \text{and} \quad Z = \sum_{mn} Z_{mn} \sin(m\theta - n\varphi)$$

- SPEC also needs **two profiles**, e.g. pressure  $p(\psi_v)$  and transform  $\epsilon^{\pm}(\psi_v)$ .

- **Solution** in terms of the vector potential  $\mathbf{A} = A_\theta \nabla \theta + A_\varphi \nabla \varphi$ , written in the form

$$A_\alpha(s, \theta, \varphi) = \sum_{m,n,l} A_{\alpha,m,n,l} T_l(s) \cos(m\theta - n\varphi)$$

where  $T_l(s)$  are the Chebyshev polynomials.

- **Resolution parameters** are  $M_{pol} = \max(m)$ ,  $N_{tor} = \max(n)$ ,  $L_{rad} = \max(l)$ .

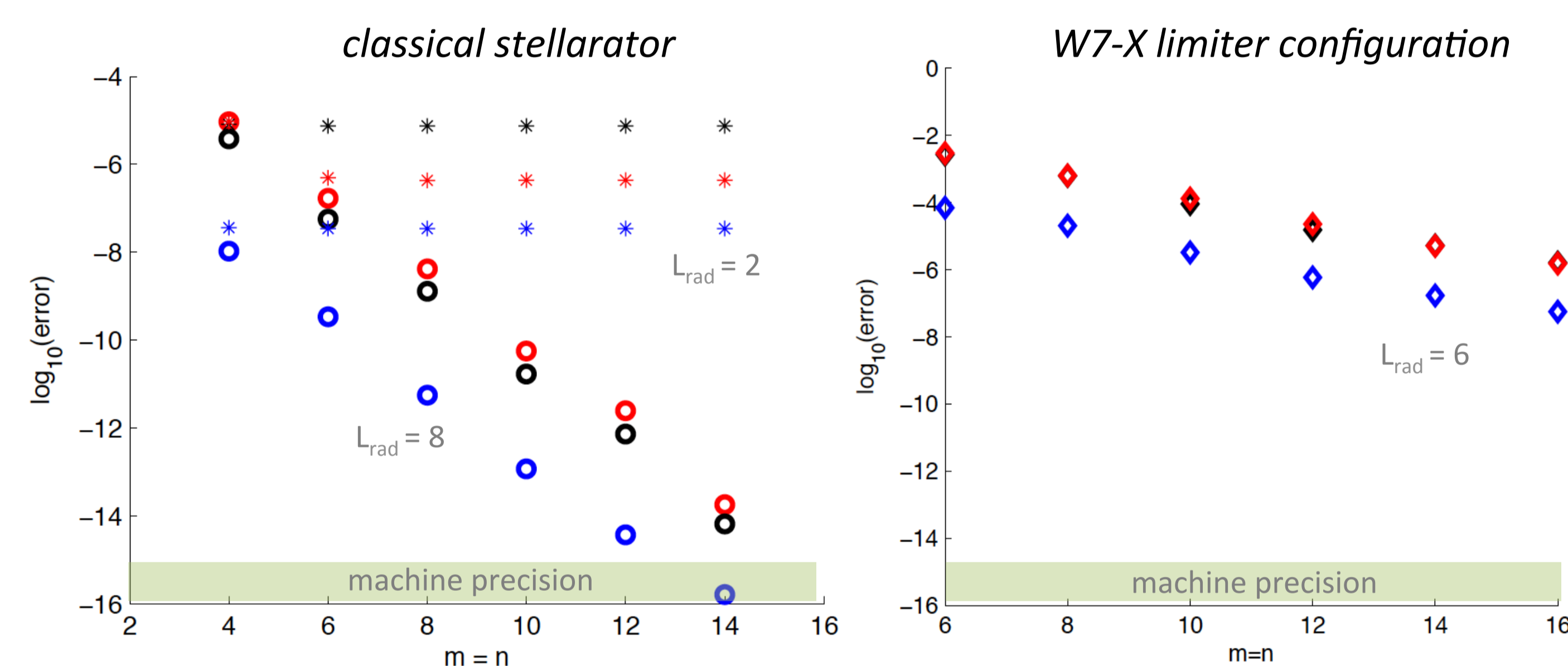
## 3. Stellarator vacuum fields

- There is a unique magnetic field  $\mathbf{B}$  (up to a scale factor) that satisfies

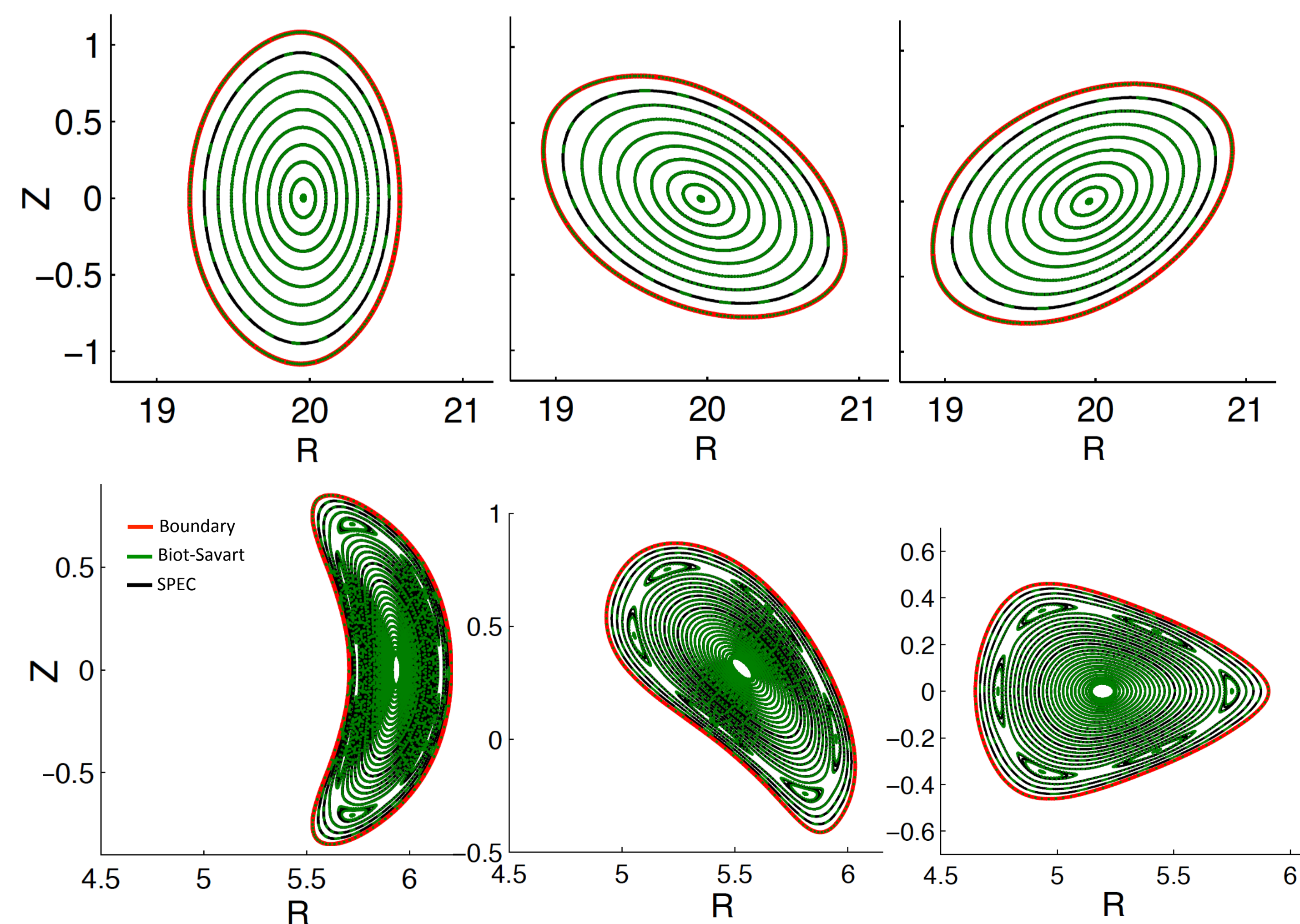
$$\begin{aligned} \nabla \times \mathbf{B} &= 0 \quad \text{in } V & (1) \\ \nabla \cdot \mathbf{B} &= 0 \quad \text{in } V & (2) \\ \mathbf{B} \cdot \hat{\mathbf{n}} &= 0 \quad \text{in } \partial V & (3) \end{aligned}$$

- SPEC can be used to calculate vacuum fields by setting  $N = 1$  and  $p = 0, \mu = 0$ .
- Two stellarator configurations are considered: a classical  $l = 2$  stellarator and W7-X OP1.1 configuration.
- Boundary harmonics are extracted from field-line-tracing performed on Biot-Savart calculations.
- **VERIFICATION**: numerical proof that Eqs. (1)-(3) are satisfied to arbitrary precision as resolution is increased.
- Equations (2) and (3) are satisfied by construction. For Eq. (1), we define

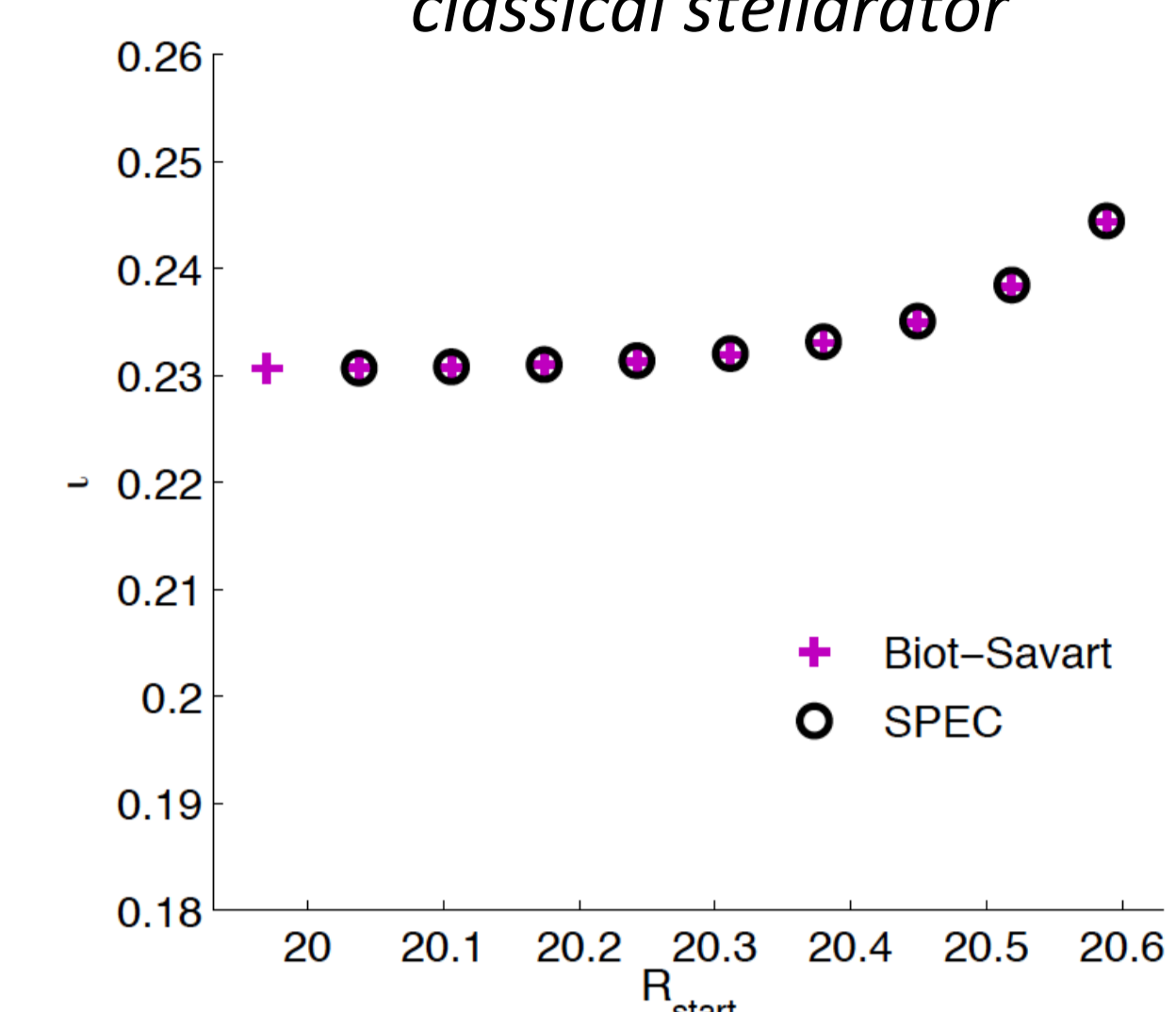
$$\epsilon_\alpha = \frac{1}{V} \int ds \oint d\theta \oint d\varphi \mathcal{J}(s, \theta, \varphi) |(\nabla \times \mathbf{B} - \mu \mathbf{B}) \cdot \nabla \alpha|$$



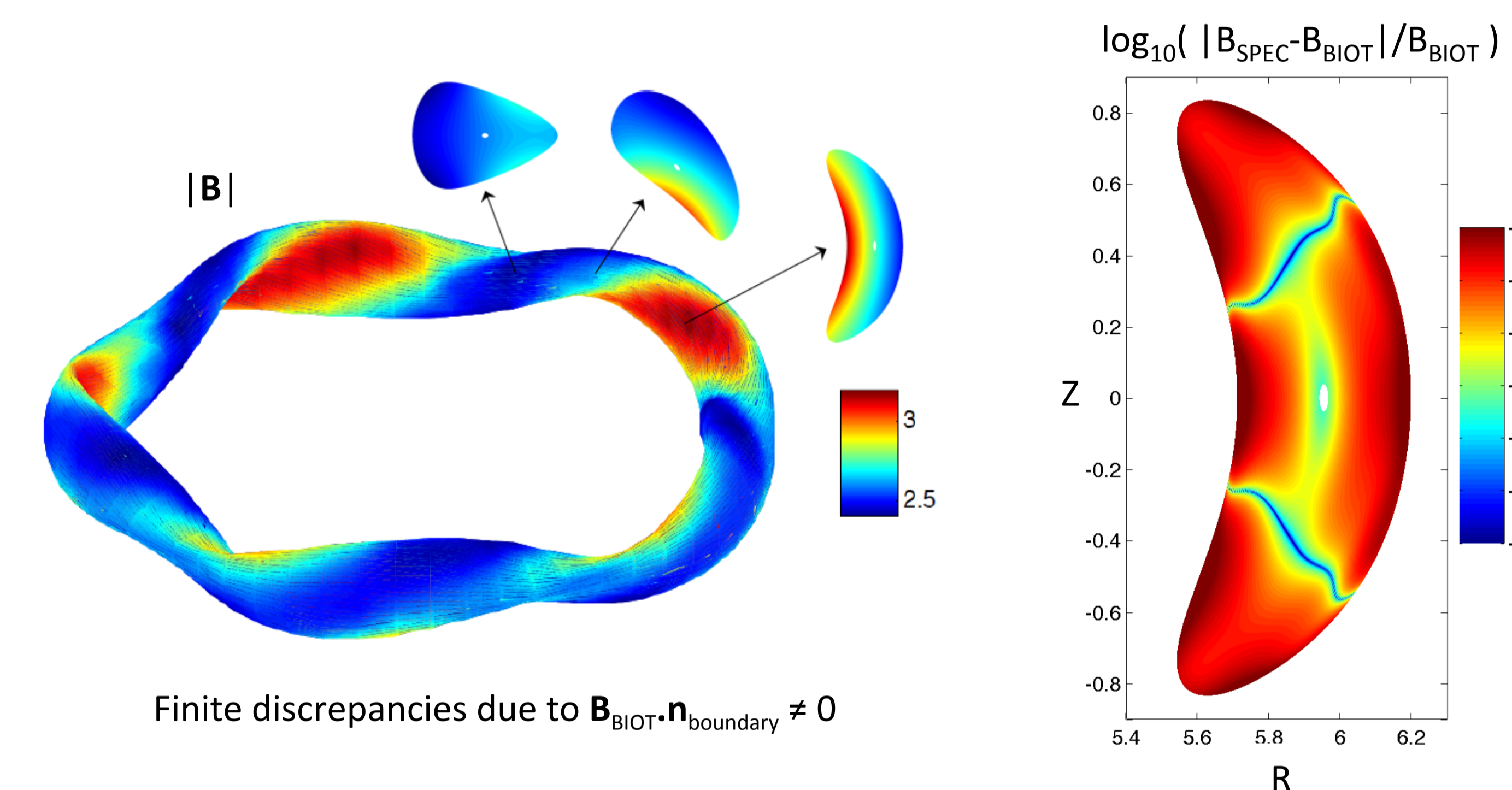
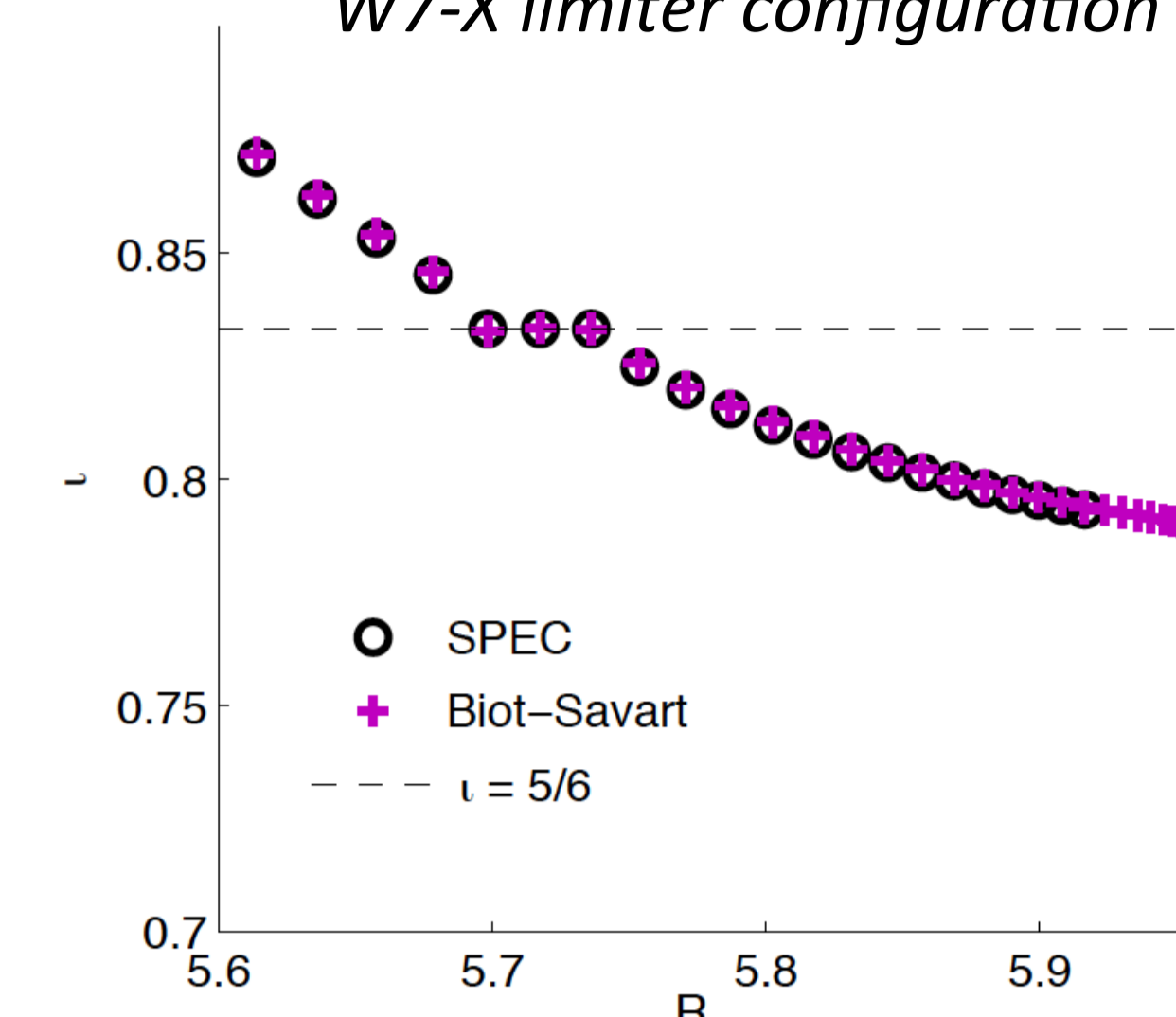
- **BEYOND VERIFICATION**: compare Poincaré plots, rotational transform, and  $|\mathbf{B}|$ , to Biot-Savart solution.



classical stellarator



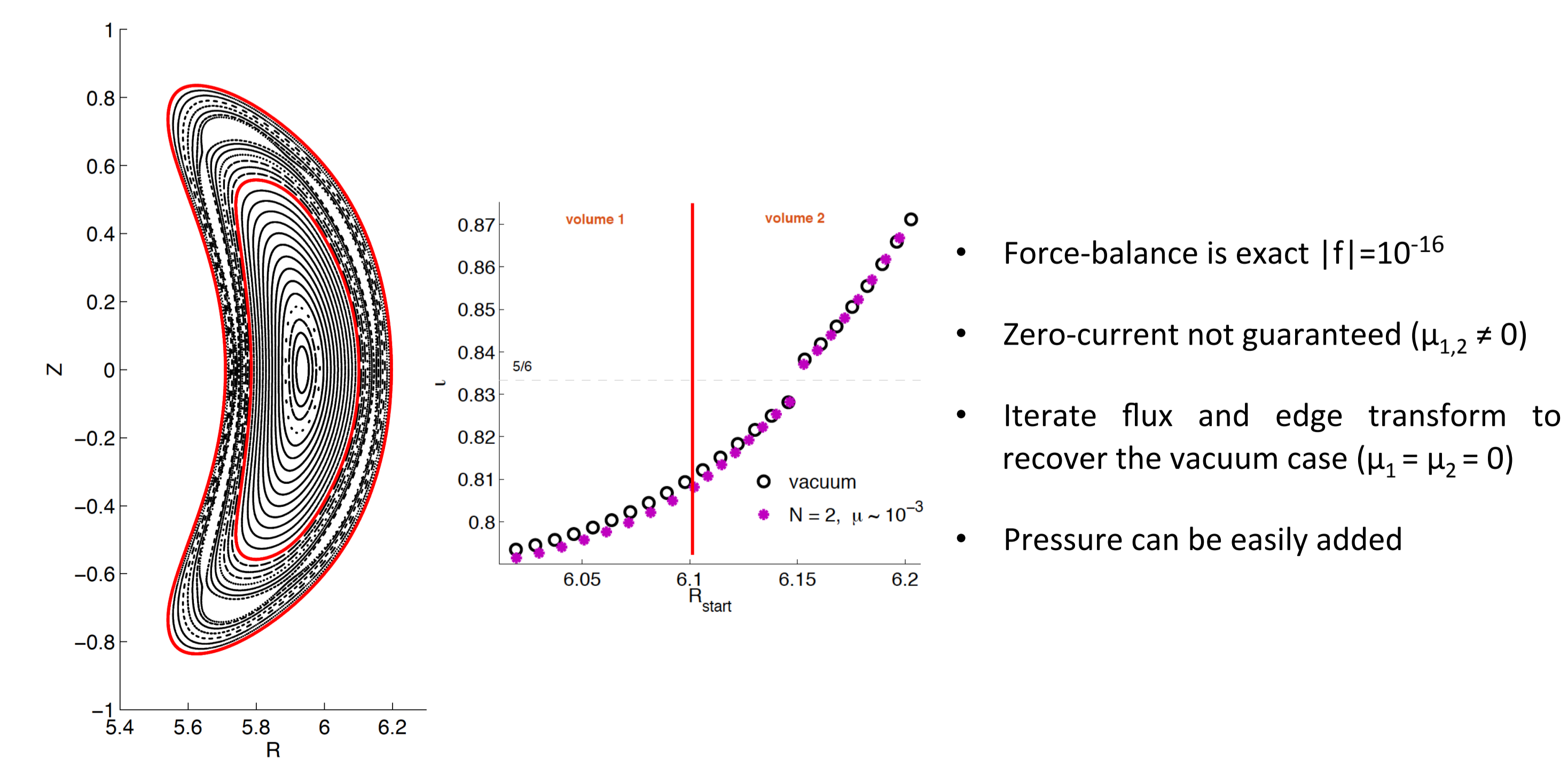
W7-X limiter configuration



Finite discrepancies due to  $\mathbf{B}_{BIOT} \cdot \mathbf{n}_{boundary} \neq 0$

## 4. Stellarator equilibria with KAM surfaces

- The **KAM theorem** ensures that the most robust surfaces are the most irrational (noble irrationals).
- Run SPEC with  $N = 2$  and a KAM surface with  $\epsilon^{\pm} = \epsilon_{noble}$ . As input, specify toroidal fluxes and edge transform.



- Force-balance is exact  $|f| = 10^{-16}$
- Zero-current not guaranteed ( $\mu_{1,2} \neq 0$ )
- Iterate flux and edge transform to recover the vacuum case ( $\mu_1 = \mu_2 = 0$ )
- Pressure can be easily added

## 4. References

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