

Edge Turbulence Imaging in Alcator C-Mod and NSTX

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Outline

- Motivations and goal
- Gas Puff Imaging (GPI) diagnostic
- C-Mod imaging results
 - Comparison of k-spectrum with simulations
- NSTX imaging results
 - Transition from L-H mode and density limit
- Tentative Conclusions
- Plans

Motivations

- Edge plasma conditions will be important for any magnetic fusion reactor
 - boundary condition for core stability (e.g. in H-mode)
 - flow through edge determines plasma-wall interaction
- Edge turbulence seems to be a dominant mechanism which determines edge plasma conditions
 - edge radial transport is normally \gg neoclassical
 - edge turbulence is normally very large ($\tilde{n}/n \approx 10\%$)

Goal

- Understand edge turbulence in present devices by measuring its 2-D space vs. time structure and comparing the results with non-linear edge turbulence simulations
- By “*understand*” we mean ideally:
 - measurements verified with different diagnostics
 - simulations checked by using different codes
 - experiments done on several different machines
 - experiments and simulation agree within $\approx 20\%$
 - simulation successfully predicts results in new regimes
 - results can be understood by simple theoretical models
- This is a challenging scientific goal and will also help to design and/or operate a “next-step” burning plasma experiment

Gas Puff Imaging Diagnostic

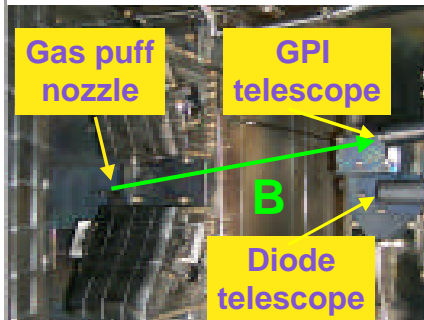
C-Mod

- Gas puff imaging (GPI) telescope views neutral line emission from He or D₂ gas puff along B field at the plasma edge (like BES but uses neutral gas instead of NBI)

$$S(\text{photons/cm}^3) = n_0 f(n_e, T_e) A$$

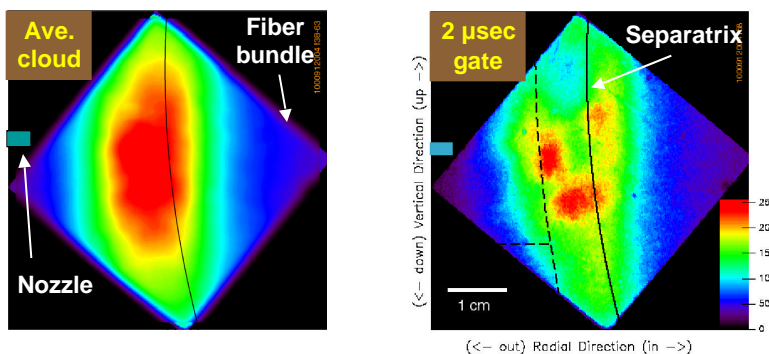
where the radiative decay rate is $A \gg 10^7 \text{ sec}^{-1}$ for these lines.

- Space and time variation of neutral light emission is measured with fast-gated cameras and PMs or PDs on discrete chords to determine edge turbulence structure (assumes $k_{\parallel} \ll k$)
- Gas puff changes plasma density by 1% in C-Mod and 10% in NSTX, but this probably does not perturb the edge turbulence significantly



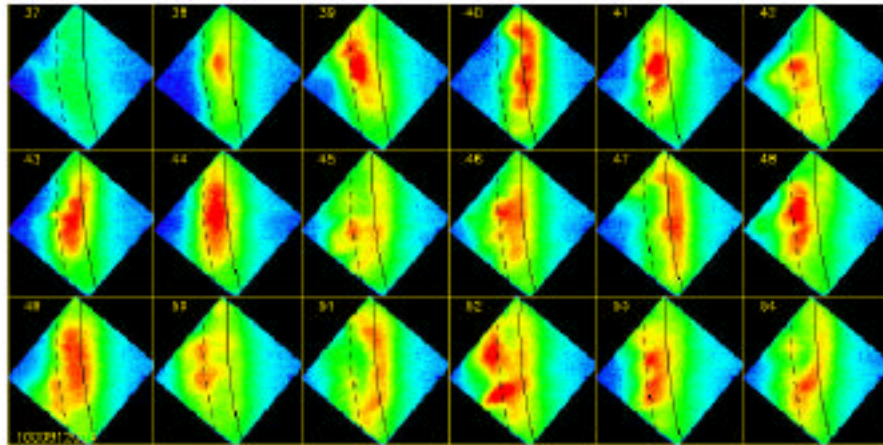
Example of 2-D GPI Image in C-Mod

- Camera images D (656 nm) light using optical system with 2-3 mm spatial resolution over 6 cm x 6 cm field
- Turbulence can be seen with camera gating of 2 μsec/frame as structure within time-averaged envelope of D cloud



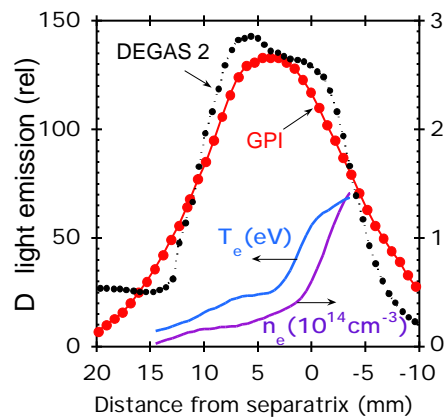
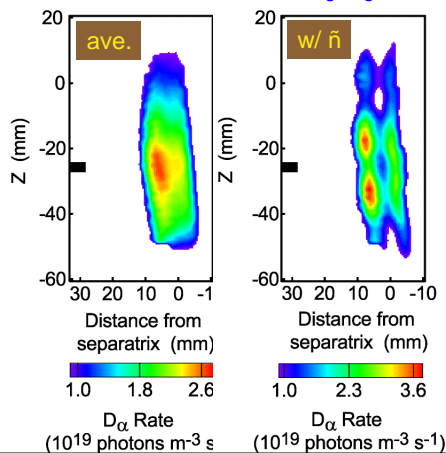
Typical Images During a Discharge

- Camera takes images with 2 μ sec gating @ 60 frames/sec
- Time between frames \gg turbulence autocorrelation time

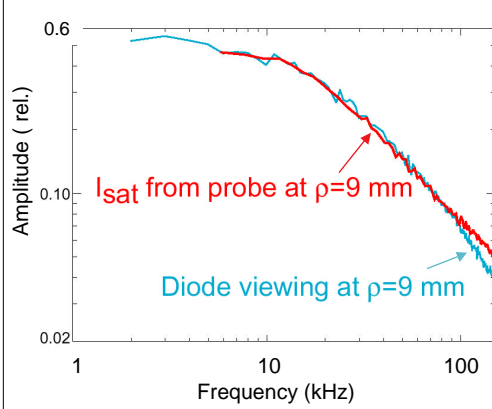


Interpretation of GPI Using DEGAS 2

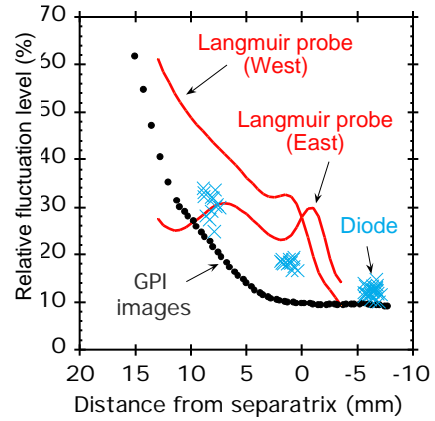
- 2-D neutral/atomic physics model of D_2 gas puff (Stotler)
 - $D \propto n_0 f(n_e, T_e) \propto n_e^{0.5} T_e^{0.5}$ just outside separatrix
- $\Rightarrow D \propto n_e/n_e$ at this radius (0.46 \rightarrow 0.83 vs. r/a)



Comparison with Langmuir Probe



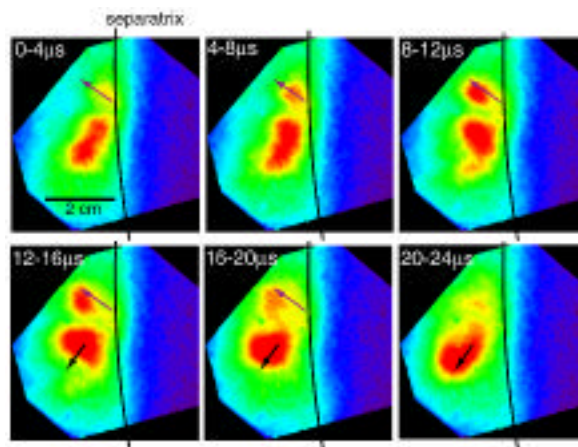
Frequency spectra of GPI and Langmuir probe similar (as at ASDEX and Caltech)



Relative fluctuation level of GPI and Langmuir probe are similar (contains both n_e/n_e and T_e/T_e)

Motion of Edge Turbulence

- Made with Princeton Scientific Instruments PSI-3 camera, with 12 frames of 64 x 64 pixels at 4 $\mu\text{sec}/\text{frame}$

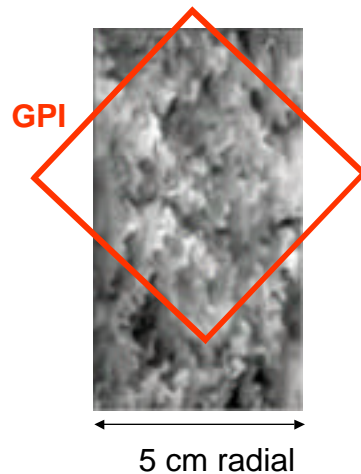


Simulations of Edge Turbulence

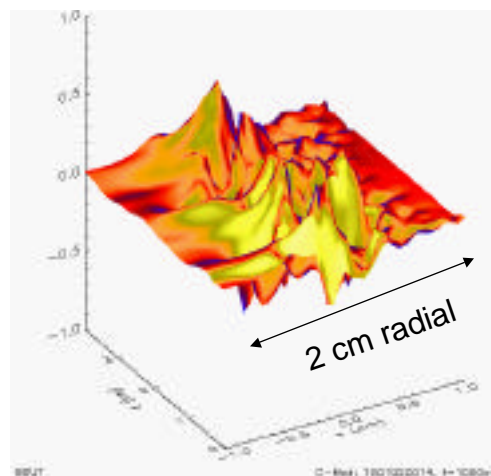
- Use 3-D EM fluid (Braginskii) codes for edge simulation, which are good approximations when:
 - gyroradius small: $k_{\perp} \rho_s \approx 0.026$
 - collisionality large: $\nu_d \approx 0.35$ ($\nu_{ie}/L_{\parallel} \approx 0.1$)
 - beta small: $\beta_{mhd} \approx 0.03$ ($\beta \approx 3 \times 10^{-5}$)
- Dominant linear instability is ES resistive ballooning mode
 - typical RBM length scale: $\rho_o \approx 2.2$ mm
 - typical RBM time scale: $t_o \approx 1.2$ μ sec
- Local DBM-type code has circular closed field line geometry but may be OK at high collisionality outside separatrix
- BOUT code uses profiles and realistic magnetic geometry

Simulation Results - the videos

Local "DBM"
(Hallatschek, Rogers)



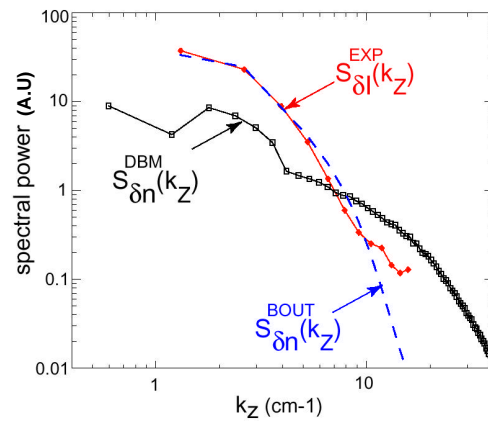
BOUT
(Xu, Nevins)



Comparison of Experiment and Simulations

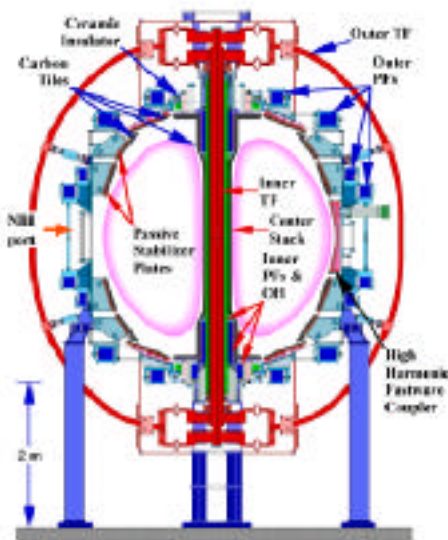
Comparison with Local Simulation:

- L_{pol} 0.6 cm in simulation
 L_{pol} 0.8 ± 0.2 cm in GPI
- \tilde{n}/n 18% in simulation
 T_e/T_e 13% in simulation
 I/I 25-35% in LP/Diode
- τ_{auto} 5-6 μ sec in simulation
 τ_{auto} 10-20 μ sec in LP/Diode
- D ~ 0.2 m²/sec in simulation
 D ~ 0.2 m²/sec in experiment
 (LaBombard - KP1.023)



=> Initial comparisons are encouraging

NSTX - National Spherical Torus Exp't



$R = 85$ cm

$a = 68$ cm

$A = 1.25$

$I = 1.5$ MA

$B = 5$ T

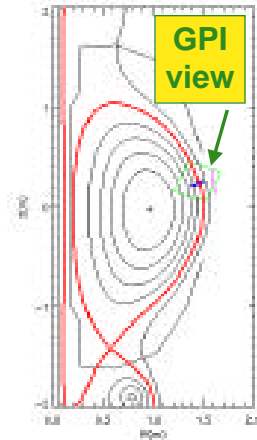
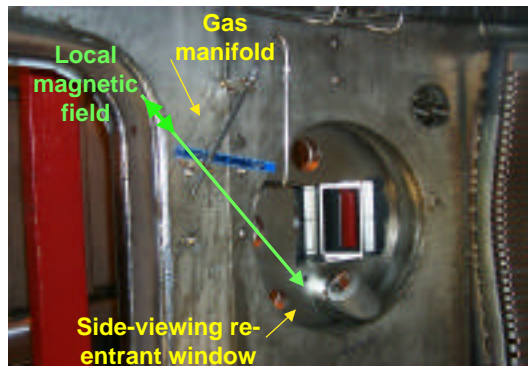
5 MW NBI

6 MW ICRH

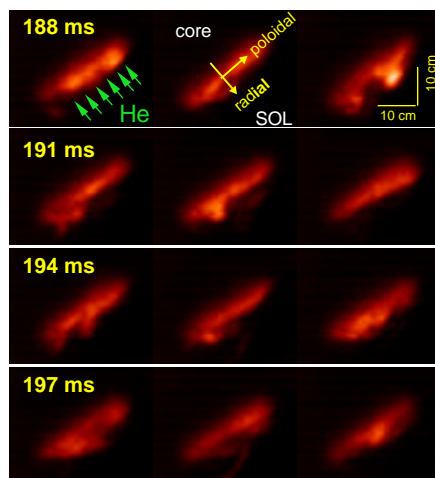
$\tau_T = 32\%$

GPI Diagnostic Set-up in NSTX

- Similar to C-Mod system but using re-entrant port and linear gas manifold instead of single-point gas nozzle
- Use He puffs into D plasmas or D puffs into He plasmas (with similar results)



Typical Edge Turbulence Images in NSTX



Shot 105711

- $B=4.5$ kG, $I=1.0$ MA Ohmic
- He puff into D_2 discharge
- GPI with HeI filter (587.6 nm)
- $10 \mu\text{s}$ exposure each frame

GPI Videos for NSTX

- For LANL camera at 1000 frames/sec, images are uncorrelated from frame-to-frame

#105711
(1.0 MA, 4.5 kG)
no-H mode

#105710
(1.0 MA, 4.5 kG)
H-L transition

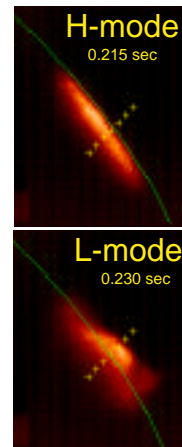
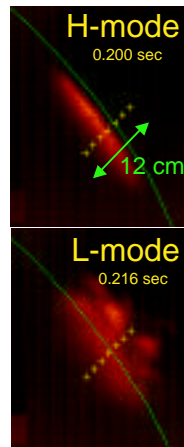
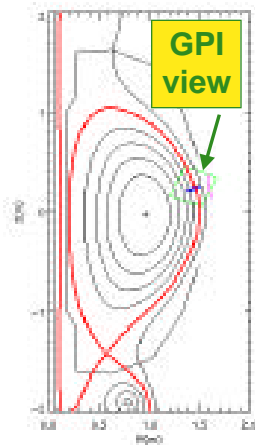
#105627
(0.7 MA, 3 kG)
Ohmic

- For PSI-4 camera, captures motion at 100,000 frames/sec

108164
(0.8 MA, 3.0 T)
Ohmic

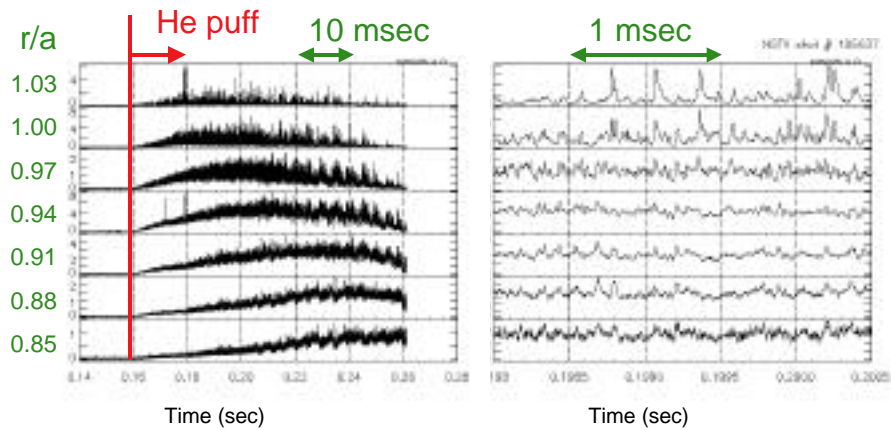
Radial Array of Fast GPI Chords

- Fast chords view 2 cm diameter “spots” in 7 channel radial array with PM tube (bandwidth 100 kHz)

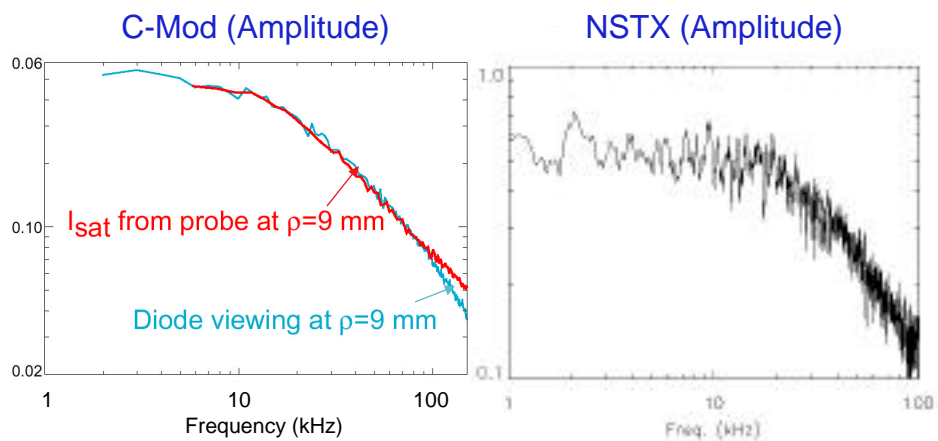


Typical Signals from Fast GPI Chords

- Signals from 7 chords digitized for 0.128 msec @ 500 kHz
- Near outer wall see “intermittant” 100% fluctuations
- Nearer center, see “Gaussian” 20% fluctuations



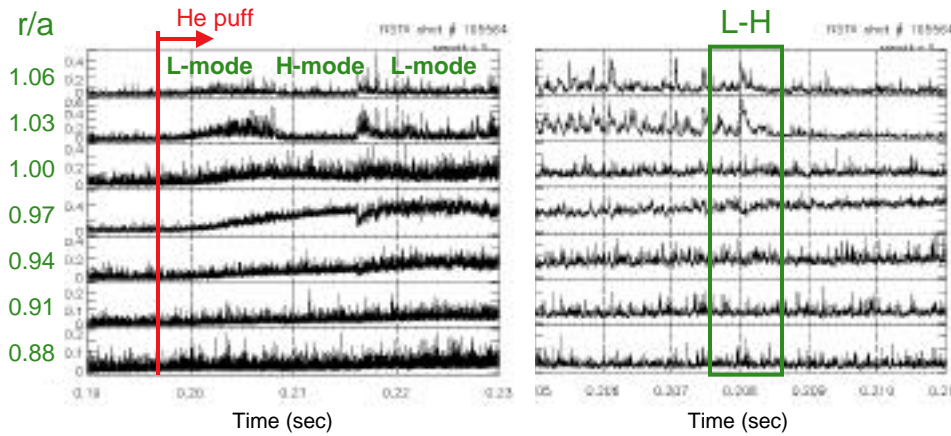
Comparison of Frequency Spectra



- Frequency spectra are similar, as seen previously for other tokamaks and stellarators (e.g. Pedrosa et al PRL '99)

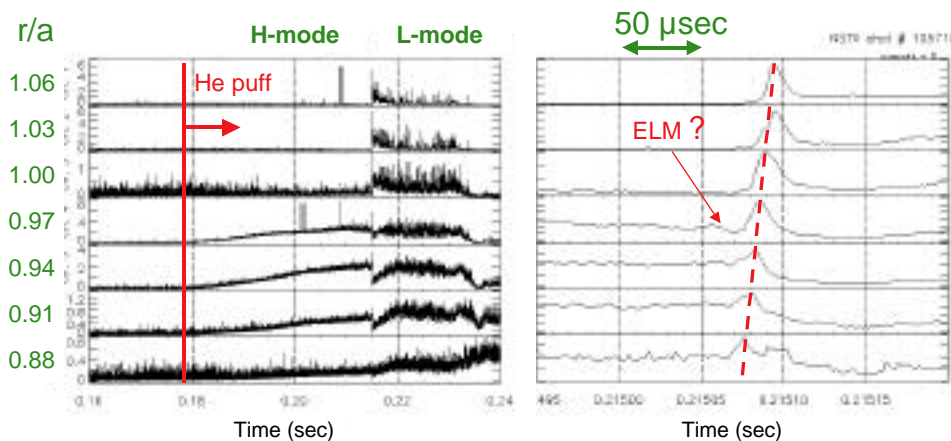
Edge Turbulence Signals at L-H Transition

- L-H transition occurs within 1 msec on most channels
- Perhaps preceded by large “bursts” in outermost channels



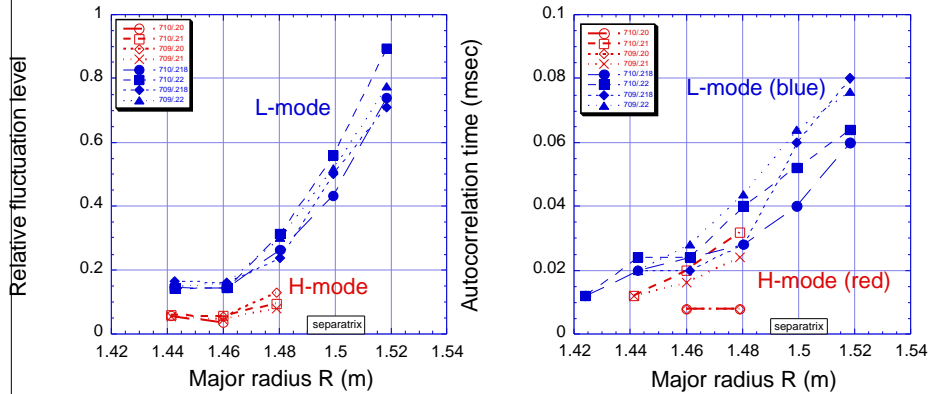
Edge Turbulence Signals at H-L Transition

- H-L transitions occurs within 20 μ sec in most channels
- Transition seems to propagate outward at 10 cm / 20 μ sec
- Transition seems to start with coherent mode (ELM ?)

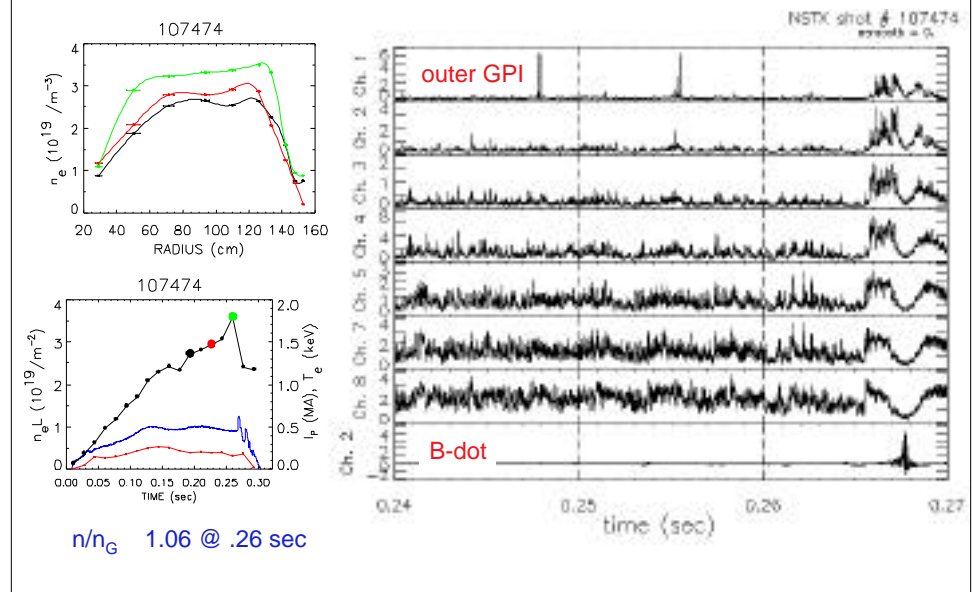


Fluctuation Profiles in H-mode vs. L-mode

- GPI fluctuation level lower in H-mode (but not-zero)
- **BUT** Need to consider profile changes and Hel $f(n_e, T_e)$

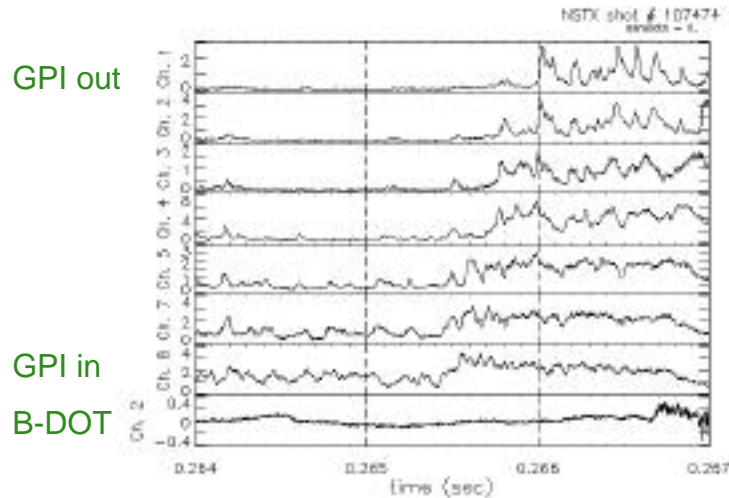


GPI Signals Near Density Limit



Outward Propagation Near Density Limit

- Increase in GPI signal seems to propagate from inside to outside at a speed $10 \text{ cm} / 0.5 \text{ msec} = 200 \text{ m/sec}$



Initial Comparison with Theory

- BOUT simulations have been run, but only for estimated edge profiles in NSTX (not yet measured well)
- Results from BOUT are qualitatively similar to GPI results
 - autocorrelation time $15\text{-}30 \mu\text{sec}$
 - fluctuation level $n/n = 10\text{-}20\%$ near separatrix
 - T_e/T_e (0.2-0.5) n/n due to low collisionality
 - 5 cm , $\parallel 10 \text{ m}$ near outer midplane
- Degas-2 modeling of HeI (587.6) has been started
 - $S \propto n_e^{0.7} T_e^{0.5}$ near center of GPI cloud for NSTX
 - H-mode emission profile is narrower mainly due to the narrower time-averaged edge profiles

Tentative Conclusions

- Edge turbulence structure looks like a combination of “blobs” and “waves”, similar to that seen with other diagnostics
- Initial comparisons with simulation / theory are encouraging
- Definitive comparisons with theory will have to address what are the appropriate “inputs” needed for the simulations:
 - use time averaged profiles or fluxes through edge ?
 - need limiter configuration and/or divertor geometry ?
 - are atomic physics and neutral effects significant ?
 - can the edge really be decoupled from the core ?

Plans

- 2-color and 2-time GPI imaging in Alcator C-Mod using two Xybion intensified gated cameras (60 Hz each)
- Make more ultra-fast videos using 28 frame PSI-4 camera, and later the new 312 frame PSI-5 camera
- Ask theorists to run simulations for more cases and to “benchmark” their codes against each other
- Feed 2-D patterns of density and temperature fluctuations into DEGAS 2 model and compare with GPI data