# Present Status of M3D 

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## Capsule History of M3D

- Original MH3D (W.P., early 1980s) was a serial Fortran code in a single source file solving resistive MHD using finite differences on a radial mesh with spectral treatment of $\theta$ and $\phi$.
- Over more than a decade, gradual refinements and enhancements of the physical model (hybrid [W.P.] and two-fluid [L.S.] models) and numerical scheme (finite elements [H.S.]) were accreted onto this program, forming the Multilevel 3D Code (M3D). This was eventually parallelized using OpenMP.
- Around 1999, X.T. set out to create an MPI version of the code. Doing a complete rewrite, he created a C code distributed over many files within two layers of directories, using linear triangular finite elements on a domain decomposed both poloidally and toroidally to solve MHD only, using the PETSc software library to handle communications and linear solves. This was ParM3D.
- In order to retain much of the physics and flexibility of the original version, H.S. undertook to couple the two codes together, using ParM3D for mesh generation, I/O, and linear solvers with the original Fortran "m1.F" as the physics driver. Data would be passed between the C and Fortran parts of the new code using a new set of Fortran and C interface routines. Much of the now-unused part of ParM3D was left in the distribution in vestigial form. This is M3DP (still referred to as M3D).
- A CVS repository for the modern M3D was started in 2001. Changes made since then are archived in $/ \mathrm{p} / \mathrm{m} 3 \mathrm{~d} /$ README on the PPPL Unix cluster. Highlights include refinement of the two-fluid options; improvement and parallelization of the hot particle treatment; addition of $2^{\text {nd }}$ - and $3^{\text {rd }}$-order element options; and addition of vacuum region/resistive wall capability. The current version number is 3.5.12.


## DIAtfornns

M3D has been ported to the following computers at NERSC, NCCS, Princeton, and ANL:

|  | OpenMP | MPI |
| :--- | :---: | :---: |
| IBM SP (Seaborg) | Y | Y |
| Opteron cluster (Jacquard) |  | Y |
| IBM Power 5 (Bassi) | $?$ | Y |
| Cray X1E (Phoenix) |  | $\mathrm{Y}^{*}$ |
| Cray XT3, XT4 (Jaguar) |  | Y |
| SGI Origin 2000 (Hecate) | Y |  |
| SGI Altix (MHD) | Y | Y |
| BlueGene/L, Argonne |  | Y |

## Statistics

- Source code is divided into four directories (m3d, mhd, mesh, utility) with 34 subdirectories.
- There are approximately 264 C source files, 216 C header files, 33 Fortran source files, 16 Fortran header files, and 35 Makefiles.
- There are approximately 52,000 lines of $C$ and 97,000 lines of Fortran source code.
- This includes a lot of code that is no longer executed (or, in many cases, compiled), but excludes standalone post-processing utilities and many trial routines that have not yet been committed to the repository.
- Libraries required include PETSc, parallel HDF5, and sometimes FFTW.
- Three standard input files (plus batch script), others optional; recently consolidated to a single Python script.
- Performance record: 240 Gflops on 10,240 XT3 cores (VN mode) during a 1D weak scaling test.


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## Extended MHD Equations

$$
\begin{aligned}
& \frac{\partial \rho}{\partial t}+\nabla \cdot\left(\rho \mathbf{v}_{i}\right)=0 \\
& \rho\left[\frac{\partial \mathbf{v}}{\partial t}+\mathbf{v} \cdot \nabla \mathbf{v}+\left(\mathbf{v}_{i}^{*} \cdot \nabla\right) \mathbf{v}_{\perp}\right]=-\nabla \mathbf{p}+\mathbf{J} \times \mathbf{B}+\mu \nabla^{2} \mathbf{v} \\
& \mathbf{E}+\mathbf{v} \times \mathbf{B}=\eta \mathbf{J}-\frac{\nabla_{\|} p_{e}}{n e} \\
& \frac{\partial \mathbf{B}}{\partial t}=-\nabla \times \mathbf{E} \\
& \mathbf{J}=\nabla \times \mathbf{B}
\end{aligned}
$$

$$
\frac{\partial p}{\partial t}+\mathbf{v} \cdot \nabla p=-\gamma p \nabla \cdot \mathbf{v}+\nabla \cdot n \chi_{\perp} \nabla\left(\frac{p}{\rho}\right)-\mathbf{v}_{i}^{*} \cdot \nabla p-\gamma p \nabla \cdot \mathbf{v}_{i}^{*}+\frac{\mathbf{J} \cdot \nabla p_{e}}{n e}+\gamma p_{e} \mathbf{J} \cdot \nabla\left(\frac{1}{n e}\right)
$$

$$
\frac{\partial p_{e}}{\partial t}+\mathbf{v} \cdot \nabla p_{e}=-\gamma p_{e} \nabla \cdot \mathbf{v}+\nabla \cdot n \chi_{\perp e} \nabla\left(\frac{p_{e}}{\rho}\right)+\frac{\mathbf{J}_{\|} \cdot \nabla p_{e}}{n e}-\gamma p_{e} \nabla \cdot\left(\mathbf{v}_{e}^{*}-\frac{\mathbf{J}_{\|}}{n e}\right)
$$

where
$\mathbf{v}_{e}^{*} \equiv-\frac{\mathbf{B} \times \nabla p_{e}}{n e B^{2}}, \quad \mathbf{v}_{i}^{*} \equiv \mathbf{v}_{e}^{*}+\frac{\mathbf{J}_{\perp}}{n e}$,
$\mathbf{v} \equiv \mathbf{v}_{i}-\mathbf{v}_{i}^{*}=\mathbf{v}_{e}-\mathbf{v}_{e}^{*}+\frac{\mathbf{J}_{\|}}{n e}$

## M3D Scalar Variables

## Field Variables

Write

$$
\vec{B}=\nabla \psi \times \nabla \phi+\frac{1}{R} \nabla_{\perp} F+\left(R_{0}+\tilde{I}\right) \nabla \phi
$$

where

$$
\nabla_{\perp}^{2} F=-\frac{1}{R} \frac{\partial \tilde{I}}{\partial \phi}
$$

so that

$$
\vec{J}=\left(\nabla \tilde{I}-\frac{1}{R} \nabla_{\perp} F^{\prime}\right) \times \nabla \phi+\frac{1}{R^{2}} \nabla_{\perp} \psi^{\prime}-C \nabla \phi
$$

where primes denote derivatives with respect to $\phi$ and

$$
C \equiv-R J_{\phi}=\Delta^{*} \psi+\frac{1}{R} \frac{\partial F}{\partial z}
$$

## Velocity Variables

Write

$$
\vec{V}=\frac{R^{2}}{R_{0}} \nabla U \times \nabla \phi+\nabla_{\perp} \chi+V_{\phi} \hat{\phi}
$$

Others

$$
\rho, p_{(e, i)} \text { or } T_{(e, i)}
$$

Note that

$$
\nabla_{\perp}^{2} \psi \equiv \frac{\partial^{2} \psi}{\partial R^{2}}+\frac{\partial^{2} \psi}{\partial z^{2}},
$$

$\Delta^{*} \psi \equiv \nabla_{\perp}^{2} \psi-\frac{1}{R} \frac{\partial \psi}{\partial R}=\frac{\partial^{2} \psi}{\partial R^{2}}-\frac{1}{R} \frac{\partial \psi}{\partial R}+\frac{\partial^{2} \psi}{\partial z^{2}}$,
and

$$
\Delta^{\dagger} \psi \equiv \nabla_{\perp}^{2} \psi+\frac{1}{R} \frac{\partial \psi}{\partial R}=\frac{\partial^{2} \psi}{\partial R^{2}}+\frac{1}{R} \frac{\partial \psi}{\partial R}+\frac{\partial^{2} \psi}{\partial z^{2}} .
$$

## M3D Form of the Resistive MHD Equations

Define Poisson Bracket $[A, B] \equiv \nabla_{\perp} A \times \nabla_{\perp} B \cdot \hat{\phi}=\frac{\partial A}{\partial R} \frac{\partial B}{\partial z}-\frac{\partial A}{\partial z} \frac{\partial B}{\partial R}$
and $(A, B) \equiv \nabla_{\perp} A \cdot \nabla_{\perp} B=\frac{\partial A}{\partial R} \frac{\partial B}{\partial R}+\frac{\partial A}{\partial z} \frac{\partial B}{\partial z}$
Continuity:

$$
\begin{equation*}
\frac{\partial \rho}{\partial t}=-\nabla \cdot(\rho \vec{V})=-\rho\left(\Delta^{\dagger} \chi+\frac{2}{R_{0}} \frac{\partial U}{\partial z}+\frac{1}{R} \frac{\partial V_{\phi}}{\partial \phi}\right)-\frac{R}{R_{0}}[\rho, U]-(\rho, \chi)-\frac{V_{\phi}}{R} \frac{\partial \rho}{\partial \phi} \tag{1}
\end{equation*}
$$

Operate on the momentum equation with $-R_{0} \hat{\phi} \cdot \nabla \times$ to get an equation for the evolution of $\Delta^{\dagger} U \equiv \nabla_{\perp}^{2} U+\frac{1}{R} \frac{\partial U}{\partial R}$ (called " $W$ " in the code):

$$
\begin{align*}
& \frac{\partial}{\partial t} \Delta^{\dagger} U=\frac{R}{R_{0}}\left[U, \Delta^{\dagger} U\right]-\left(\chi, \Delta^{\dagger} U\right)-\Delta^{\dagger} U\left(\Delta^{\dagger} \chi+\frac{2}{R_{0}} \frac{\partial U}{\partial z}\right)-\frac{V_{\phi}}{R} \frac{\partial}{\partial \phi} \Delta^{\dagger} U-\left(\frac{V_{\phi}}{R}, \frac{\partial U}{\partial \phi}\right) \\
& +2 R_{0} \frac{V_{\phi}}{R} \frac{\partial}{\partial z} \frac{V_{\phi}}{R}+\frac{R_{0}}{R}\left[\frac{V_{\phi}}{R}, \frac{\partial \chi}{\partial \phi}\right]+R_{0}\left\{\vec{B} \cdot \nabla\left(\frac{C}{R^{2} \rho}\right)+\vec{J} \cdot \nabla\left(\frac{1+\tilde{I} / R_{0}}{R^{2} \rho}\right)\right\}+\frac{2}{R^{2} \rho} \frac{\partial p}{\partial z}  \tag{2a}\\
& +R\left[\frac{1}{R^{2} \rho}, p\right]-R_{0} \nabla \phi \cdot \nabla \times\left(\frac{\mu \nabla^{2} \vec{V}}{\rho}\right)
\end{align*}
$$

## Evolution of the Compressible Velocity

From the definition of the velocity, it is clear that

$$
\frac{\partial \chi}{\partial R}=\hat{R} \cdot \vec{V}-\frac{R}{R_{0}} \frac{\partial U}{\partial z} \quad \text { and } \quad \frac{\partial \chi}{\partial z}=\hat{z} \cdot \vec{V}+\frac{R}{R_{0}} \frac{\partial U}{\partial R}
$$

so that, again using the momentum equation,

$$
\begin{align*}
& \frac{\partial}{\partial t}\left(\frac{\partial \chi}{\partial R}\right)=-\frac{R}{R_{0}} \frac{\partial}{\partial z}\left(\frac{\partial U}{\partial t}\right)-\vec{V}_{\perp} \cdot \nabla_{\perp}\left(\frac{\partial \chi}{\partial R}+\frac{R}{R_{0}} \frac{\partial U}{\partial z}\right)-\frac{V_{\phi}}{R_{0}} \frac{\partial U^{\prime}}{\partial z}-\frac{V_{\phi}}{R} \frac{\partial \chi^{\prime}}{\partial R}+\frac{V_{\phi}^{2}}{R}-\frac{1}{\rho} \frac{\partial p}{\partial R} \\
& +\frac{1}{R^{2} \rho}\left(R_{0}+\tilde{I}\right)\left[\frac{1}{R}\left(\frac{\partial F^{\prime}}{\partial R}+\frac{\partial \psi^{\prime}}{\partial z}\right)-\frac{\partial \tilde{I}}{\partial R}\right]+\frac{C}{R^{2} \rho}\left(\frac{\partial F}{\partial z}-\frac{\partial \psi}{\partial R}\right)+\frac{\mu}{\rho} \hat{R} \cdot \nabla^{2} \vec{V} \tag{2b}
\end{align*}
$$

and

$$
\begin{align*}
& \frac{\partial}{\partial t}\left(\frac{\partial \chi}{\partial z}\right)=\frac{R}{R_{0}} \frac{\partial}{\partial R}\left(\frac{\partial U}{\partial t}\right)-\vec{V}_{\perp} \cdot \nabla_{\perp}\left(\frac{\partial \chi}{\partial z}-\frac{R}{R_{0}} \frac{\partial U}{\partial R}\right)+\frac{V_{\phi}}{R_{0}} \frac{\partial U^{\prime}}{\partial R}-\frac{V_{\phi}}{R} \frac{\partial \chi^{\prime}}{\partial z}-\frac{1}{\rho} \frac{\partial p}{\partial z} \\
& +\frac{1}{R^{2} \rho}\left(R_{0}+\tilde{I}\right)\left[\frac{1}{R}\left(\frac{\partial F^{\prime}}{\partial z}-\frac{\partial \psi^{\prime}}{\partial R}\right)-\frac{\partial \tilde{I}}{\partial z}\right]-\frac{C}{R^{2} \rho}\left(\frac{\partial F}{\partial R}+\frac{\partial \psi}{\partial z}\right)+\frac{\mu}{\rho} \hat{z} \cdot \nabla^{2} \vec{V} \tag{2c}
\end{align*}
$$

## Evolution of the Toroidal Velocity

Dot the momentum equation with $\hat{\phi}$ to find

$$
\begin{align*}
& \frac{\partial V_{\phi}}{\partial t}=\frac{R}{R_{0}}\left[U, V_{\phi}\right]-\left(\chi, V_{\phi}\right)-\frac{V_{\phi}}{R}\left(V_{\phi}^{\prime}+\frac{\partial \chi}{\partial R}\right)-\frac{V_{\phi}}{R_{0}} \frac{\partial U}{\partial z}-\frac{1}{R \rho} \frac{\partial p}{\partial \phi} \\
& +\frac{1}{R^{2} \rho}[\tilde{I}, \psi]+\frac{1}{R^{2} \rho}(\tilde{I}, F)+\frac{1}{R^{3} \rho} \frac{\partial}{\partial \phi}[\psi, F]-\frac{1}{2 R^{3} \rho} \frac{\partial}{\partial \phi}\left(\left|\nabla_{\perp} \psi\right|^{2}+\left|\nabla_{\perp} F\right|^{2}\right)  \tag{2d}\\
& +\frac{\mu}{\rho}\left[\nabla^{2} V_{\phi}-\frac{V_{\phi}}{R^{2}}+\frac{2}{R^{2}} \frac{\partial}{\partial \phi}\left(\frac{R}{R_{0}} \frac{\partial U}{\partial z}+\frac{\partial \chi}{\partial R}\right)\right]
\end{align*}
$$

## Electrostatic Potential

If $\vec{B}=\nabla \times \vec{A}$ and $\frac{\partial \vec{B}}{\partial t}=-\nabla \times \vec{E}$ then $\frac{\partial \vec{A}}{\partial t}=-\vec{E}+\nabla \Phi \quad$ where, if we choose the gauge $\nabla_{\perp} \cdot \vec{A}=0$, we find $\nabla_{\perp}^{2} \Phi=\nabla_{\perp} \cdot \vec{E}$.

For the resistive MHD Ohm's law, that means

$$
\begin{align*}
& \nabla_{\perp}^{2} \Phi=\frac{1}{R_{0}}(\tilde{I}, U)+\left(1+\frac{\tilde{I}}{R_{0}}\right) \nabla_{\perp}^{2} U-\frac{V_{\phi}}{R} \Delta^{*} \psi+\frac{R_{0}}{R^{2}} \frac{\partial \chi}{\partial z}+\left[\chi, \frac{\tilde{I}}{R}\right]-\left[F, \frac{V_{\phi}}{R}\right]-\frac{1}{R}\left(V_{\phi}, \psi\right) \\
& +\frac{\eta}{R^{2}}\left[\frac{1}{R}\left(\frac{\partial F^{\prime}}{\partial z}-\frac{\partial \psi^{\prime}}{\partial R}\right)-\frac{\partial \tilde{I}}{\partial z}+\frac{\partial C}{\partial \phi}\right]+\frac{1}{R}[\eta, \tilde{I}]-\frac{1}{R^{2}}\left[\eta, F^{\prime}\right]+\frac{1}{R^{2}}\left(\eta, \psi^{\prime}\right) \tag{3}
\end{align*}
$$

## Evolution of the Poloidal Field

The time derivative of $\psi$ (called "a" in the code) is simply $R \hat{\phi} \cdot \frac{\partial \vec{A}}{\partial t}$,

$$
\begin{equation*}
\frac{\partial \psi}{\partial t}=\frac{R}{R_{0}}[U, \psi]+\frac{R}{R_{0}}(U, F)-(\chi, \psi)+[\chi, F]+\eta C+\frac{\partial \Phi}{\partial \phi} . \tag{4}
\end{equation*}
$$

but for numerical stability, the quantity we generally choose to evolve is instead $C_{a} \equiv \Delta^{*} \psi$ :

$$
\begin{align*}
& \frac{\partial C_{a}}{\partial t}=\frac{R}{R_{0}}\left\{\left[U, C_{a}\right]+\left[\Delta^{\dagger} U, \psi\right]+2\left[\frac{\partial U}{\partial R}, \frac{\partial \psi}{\partial R}\right]+2\left[\frac{\partial U}{\partial z}, \frac{\partial \psi}{\partial z}\right]\right\}+\frac{2}{R_{0}}\left[U, \frac{\partial \psi}{\partial R}\right]+\frac{2}{R_{0} R} \frac{\partial U}{\partial z} \frac{\partial \psi}{\partial R} \\
& +\frac{R}{R_{0}}\left\{\left(U, \nabla_{\perp}^{2} F\right)+\left(\Delta^{\dagger} U, F\right)+2\left(\frac{\partial U}{\partial R}, \frac{\partial F}{\partial R}\right)+2\left(\frac{\partial U}{\partial z}, \frac{\partial F}{\partial z}\right)\right\}+\frac{1}{R_{0}}\left(\frac{\partial F}{\partial R}, U\right)-\frac{1}{R_{0} R} \frac{\partial F}{\partial z} \frac{\partial U}{\partial z} \\
& -\left\{\left(\psi, \nabla_{\perp}^{2} \chi\right)+\left(C_{a}, \chi\right)+2\left(\frac{\partial \psi}{\partial R}, \frac{\partial \chi}{\partial R}\right)+2\left(\frac{\partial \psi}{\partial z}, \frac{\partial \chi}{\partial z}\right)\right\}+\frac{1}{R}\left(\frac{\partial \chi}{\partial R}, \psi\right)+\frac{1}{R^{2}} \frac{\partial \psi}{\partial R} \frac{\partial \chi}{\partial R} \\
& +\left\{\left[\nabla_{\perp}^{2} \chi, F\right]+\left[\chi, \nabla_{\perp}^{2} F\right]+2\left[\frac{\partial \chi}{\partial R}, \frac{\partial F}{\partial R}\right]+2\left[\frac{\partial \chi}{\partial z}, \frac{\partial F}{\partial z}\right]\right\}-\frac{1}{R}\left\{\left[\frac{\partial \chi}{\partial R}, F\right]+\left[\chi, \frac{\partial F}{\partial R}\right]\right\} \\
& +\frac{\partial}{\partial \phi}\left(\nabla_{\perp}^{2} \Phi\right)-\frac{1}{R} \frac{\partial^{2} \Phi}{\partial \phi \partial R}
\end{align*}
$$

## Evolution of the Toroidal Field

The magnetic field is completely specified by two scalar functions; the auxiliary variable $F$ is related to the non-vacuum toroidal field $\tilde{I} / R$ by the elliptic equation given earlier. The evolution of $\tilde{I}$ can be found from the toroidal component of the field equation:

$$
\begin{align*}
& \frac{\partial \tilde{I}}{\partial t}=\frac{R}{R_{0}}[U, \tilde{I}]-(\chi, \tilde{I})+R\left[\frac{V_{\phi}}{R}, \psi\right]+R\left(\frac{V_{\phi}}{R}, F\right)-\left(R_{0}+\tilde{I}\right) \Delta^{*} \chi-\frac{V_{\phi}}{R} \frac{\partial \tilde{I}}{\partial \phi}  \tag{5}\\
& +\eta\left[\Delta^{*} \tilde{I}-\frac{1}{R} \nabla_{\perp}^{2} F^{\prime}+\frac{2}{R^{2}}\left(\frac{\partial \psi^{\prime}}{\partial z}+\frac{\partial F^{\prime}}{\partial R}\right)\right]-\frac{1}{R}\left[\eta, \psi^{\prime}\right]+(\eta, \tilde{I})-\frac{1}{R}\left(\eta, F^{\prime}\right)
\end{align*}
$$

## The Energy Equation

The energy equation in the resistive MHD version M3D is normally solved in terms of the plasma pressure; simple substitution of the code variables into the pressure equation gives

$$
\begin{equation*}
\frac{\partial p}{\partial t}=\frac{R}{R_{0}}[U, p]-(\chi, p)-\frac{V_{\phi}}{R} \frac{\partial p}{\partial \phi}-\gamma p\left(\frac{2}{R_{0}} \frac{\partial U}{\partial z}+\Delta^{\dagger} \chi+\frac{1}{R} \frac{\partial V_{\phi}}{\partial \phi}\right)+\rho \nabla \cdot\left[\kappa_{\perp} \nabla\left(\frac{p}{\rho}\right)\right] \tag{6}
\end{equation*}
$$

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## The M3D Mesh

- Uses linear basis functions on unstructured triangular finite element mesh in each constant- $\phi$ plane.
- 3 parameters control mesh resolution: \# of planes, \# of radial grids, \# of theta sections.
- Mesh has same topology in all planes. In the tokamak case, it has the same geometry in all planes as well.
- Mesh is aligned with equilibrium flux surfaces (from VMEC-generated input files) but does not follow field lines.
- Uses either $4^{\text {th }}$-order finite differences or pseudo-spectral derivatives between
 planes.


## Packing the Mesh at a Flux Surface

In order to resolve fine structures at a particular surface, the option exists to concentrate zones of the M3D mesh about a given minor radius (1D packing).

## Command line options:

- packingFactor <pf> Ratio of packed to unpacked mesh density.
- packingRadius $\left\langle x_{0}\right\rangle$ Relative position of packing surface (from 0 to 1).
-packingWidth $\langle w\rangle$ Relative width of peak packing area (on 0 to 1 scale).
Example: $p f=4.0 ; x_{0}=0.5 ; w=0.12$ :

Before packing
After packing

## Linear Finite Elements

Linear basis functions on a triangle:


$$
\lambda_{\alpha}\left(\vec{r}_{\alpha}\right)=1 ; \lambda_{\alpha}\left(\vec{r}_{\beta \neq \alpha}\right)=0
$$

Galerkin method: integrate equations over each basis function to get "weak form" $\rightarrow$ linear algebraic equation.

$$
f(R, z)=\sum_{j} f_{j} \lambda_{j}(R, z)
$$

Mass matrix: $\iint \lambda_{i} f(R, z) d^{2} x=\sum_{j} f_{j} \iint \lambda_{i} \lambda_{j} d^{2} x \equiv \sum_{j} M_{i, j} f_{j}$
Stiffness matrix: $\iint \lambda_{\mathrm{i}} \nabla_{\perp}^{2} f(R, z) d^{2} x=\sum_{j} f_{j} \iint \lambda_{i} \nabla_{\perp}^{2} \lambda_{j} d^{2} x=\sum_{j} f_{j}\left\{\iint \nabla_{\dot{f}} \cdot\left(\lambda_{i} \nabla \lambda_{j}\right) d^{2} x-\iint \nabla_{\perp} \lambda_{i} \cdot \nabla_{\perp} \lambda_{j} d^{2} x\right\} \equiv \sum_{j} S_{i, j} f_{j}$
"dRoverR" matrix: $\iint \frac{\lambda_{i}}{R} \frac{\partial}{\partial R} f(R, z) d^{2} x=\sum_{j} f_{j} \iint \frac{\lambda_{i}}{R} \frac{\partial \lambda_{j}}{\partial R} d^{2} x \equiv \sum_{j} R_{i, j} f_{j}$
Handy identity: $\iint_{\Delta} \lambda_{1}^{\ell} \lambda_{2}^{m} \lambda_{3}^{n} d^{2} x=2 \Delta \frac{\ell!m!n!}{(\ell+m+n+2)!}$
Lumped mass matrix (diagonal): $\overline{\mathbf{M}}_{i, j} \equiv \delta_{i, j} \sum_{j} M_{i, j}$

## Boundary Conditions

- All calculations use a fixed boundary.
- Standard cases use perfectly conducting wall, with or without a "slot".
- Slip or no-slip conditions may be imposed.
- Most of these are realized as Dirichlet b.c.s in linear solves. Exceptions: F, $\chi$ use Neumann.


## Domain Decomposition

3 parameters control domain decomposition: \# of toroidal PEs, \# of radial PEs, \# of theta PEs.

## Toroidal <br> (overhead view)

Poloidal
(cross-section view)

$B=16$
Linear solves are independent on each processor


D = 1
$F=5$


D = 3
$F=3$

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## Time Discretization, Overview

- Equations (1-6) are advanced explicitly, except for parabolic and fast wave terms.
- Time discretization is typically $1^{\text {st }}$ order, forward-in-time. $2^{\text {nd }}$-order predictor-corrector is also an option.
- Artifical sound term, if selected, is advanced in subcycles of the main time step.
- Code execution time is dominated by $\sim 13$ linear solves per time step, each of size $N$, where $N$ is the number of vertices in a single plane.
- Elliptic solves are more expensive than Helmholtz.
- Neumann b.c.s are more expensive than Dirichlet.


## Schematic of Equation Solve

Generic mixed hyperbolic/parabolic equation:

$$
\frac{\partial f}{\partial t}+u \frac{\partial f}{\partial x}=D \frac{\partial^{2} f}{\partial x^{2}}
$$

## Galerkin F.E. method

1. Explicit solve: $f^{*}=f^{n}-(\delta t) u^{n}\left(\frac{\partial f}{\partial x}\right)^{n}$
2. Implicit solve:

$$
\left[\frac{\partial^{2}}{\partial x^{2}}-\frac{1}{D \delta t}\right]\left(f^{n+1}-f_{\text {source }}\right)=-\frac{\left(f^{*}-f_{\text {source }}\right)}{D \delta t}
$$

## Order of Operations in Main Loop

1. Recompute dt based on CFL condition for shear Alfvén wave.
2. Adjust resistivity profile to track temperature.
3. Compute $I=\varepsilon+\tilde{I}, B^{2}$
4. Advance particles if hybrid option is on.
5. Solve (2a) for vorticity $w=\Delta^{\dagger} U$; ideal terms explicitly, followed by implicit solve for viscous term and elliptic solve for $U$.
6. Simultaneously solve (5) for toroidal field, (2b-c) for $\nabla_{\perp} \chi$, and ideal part of (6) for pressure or temperature implicitly (in-plane) to step over fast wave time scale. Integrate to solve for $\chi$. Many terms are still explicit; resistivity, viscosity and heat diffusion are still implicit, perpendicular to $\nabla \varphi$.
7. Apply perpendicular (or isothermal) heat conduction.
8. Advance (1) for density $\rho$.
9. Advance artificial sound wave.
10. Advance (2d) for toroidal velocity.
11. Solve elliptic equation (3) for electrostatic potential.
12. Solve (4) for $\psi$ or $\Delta^{\star}(4)$ for $C_{a}$ followed by an elliptic solve for $\psi$.
13. Solve elliptic equation for $F$.
14. Diagnostics, output, checkpointing.

## Artificial Sound Wave Substep

$$
\frac{\partial T}{\partial t}=s \frac{\mathbf{B} \cdot \nabla u}{\rho}
$$

$$
\frac{\partial u}{\partial t}=s \mathbf{B} \cdot \nabla T+v \nabla^{2} u
$$

Repeat napmax times:

- Solve $T$ equation with reduced time step rdtdp explicitly.
- Solve hyperbolic part of $u$ equation explicitly.
- Solve parabolic part of $u$ equation implicitly.
- Check stability.


## Linear vs. Nonlinear

By default, the time advance is fully nonlinear. However an option exists to search for linear toroidal eigenmodes.

- Begin by adding a perturbation with toroidal mode $\# n$ to velocity variable $U$ in equilibrium.
- With pseudospectral method, only three planes are needed to resolve the mode. (Use number of field periods $=n$ ).
- After each nonlinear advance of a variable, find the mode $n$ component, add to the $n=0$ component from the original equilibrium to get advanced-time value.
- Fastest-growing mode will eventually dominate over others; growth rate determined from rate of change of total kinetic energy.
- Rescale perturbed quantities periodically to keep total kinetic energy below nonlinear level but above noise.


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## PETSc

- Portable, Extensible Toolkit for Scientific Computation.
- MPI-based suite of data structures \& routines for parallel solution of PDEs.
- Maintained by PETSc group, Mathematics and Computer Science Division, Argonne National Lab.
- Latest version is 2.3.2.
- The MPI version of M3D is highly dependent on PETSc.
- Uses versions 2.1.6, 2.3.0.
- Parallel data structures, ghost exchanges
- Vectors (variable fields)
- Matrices (linear operators)
- Linear solves - great flexibility in solver choices
- Asymmetric operators: GMRES
- Symmetric operators: CG
- Direct solves (SuperLU), Multigrid
- Most of M3D computation occurs in PETSc solves, so we rely on PETSc optimization for performance, scalability.


## HDF5

- Hierarchical Data Format
- Widely adopted and supported portable binary format
- Allows self-describing data organized in file-systemlike hierarchies.
- Random access
- M3D uses HDF5 as its primary output option.
- A subset of the fields in the checkpoint (12 scalar, 1 vector) is written every several time steps, in single precision.
- Mesh is described as a set of triangular prisms.
- Data values are given at vertices.
- Checkpoint files can also be converted between native binary and HDF5 for intersystem portability.
- UCD (text) output is another option; the OpenMP version can also produce NCAR graphics.


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## Stellarator

## (H. Strauss)

- By generating a mesh from a 3D equilibrium file, M3D can run stellarator cases.
- Planes can be made to span just one field period.
- Toroidal derivatives require extra terms for toroidal mesh variation, impacting speed and accuracy.


$\varphi=0$

$\varphi=\pi / 6$

$\varphi=\pi / 3$


## Two Fluid

(L. Sugiyama)

- A hierarchy of extended MHD models exists in M3D.
- The simplest uses the drift ordering to approximate the ion gyroviscous stress tensor term in the momentum equation ( $-\nabla \cdot \Pi_{i}^{g v}$ ) using the diamagnetic drift velocity:

$$
\begin{gathered}
\rho\left[\frac{\partial \mathbf{v}}{\partial t}+\mathbf{v} \cdot \nabla \mathbf{v}+\left(\mathbf{v}_{i}^{*} \cdot \nabla\right) \mathbf{v}_{\perp}\right]=-\nabla\left(\mathbf{p}_{e}+\mathbf{p}_{i}\right)+\mathbf{J} \times \mathbf{B}+\mu \nabla^{2} \mathbf{v} \\
\mathbf{v}_{i}^{*} \equiv \mathbf{v}_{e}^{*}+\frac{\mathbf{J}_{\perp}}{n e}, \quad \mathbf{v}_{e}^{*} \equiv-\frac{\mathbf{B} \times \nabla p_{e}}{n e B^{2}}, \quad \mathbf{v} \equiv \mathbf{v}_{i}-\mathbf{v}_{i}^{*}=\mathbf{v}_{e}-\mathbf{v}_{e}^{*}+\frac{\mathbf{J}_{\|}}{n e}
\end{gathered}
$$

- The Hall term can also be added to Ohm's law, introducing the dispersive whistler wave, which is very difficult to stabilize.


## Hybrid (Kinetic Hot Ions)

(G. Fu)

- Gyrokinetic particle push based on GTC group's formulation.
- Large ensemble of ions substepped through interpolated M3D $\boldsymbol{B}$ field.
- Hot ions couple back to fluid model through pressure tensor:

$$
\rho \frac{d \mathbf{v}}{d t}+\rho\left(\mathbf{v}_{i}^{*} \cdot \nabla\right) \mathbf{v}_{\perp}=-\nabla P-\nabla \cdot \mathbf{P}_{h}+\mathbf{J} \times \mathbf{B}-\mathbf{b} \mathbf{b} \cdot \nabla \cdot \Pi_{i}
$$

where $\mathbf{P}_{h}=P_{\perp} \mathbf{I}+\left(P_{\|}-P_{\perp}\right) \mathbf{b b}$
based on moments taken over the particle distribution function

$$
f=\sum_{i} \delta\left(\mathbf{R}-\mathbf{R}_{i}\right) \delta\left(v_{\|}-v_{\|, i}\right) \delta\left(\mu-\mu_{i}\right)
$$

- MPI Parallelization follows domain decomposition of M3D mesh; particles can move between processors.
- Typical particle push time is comparable to fluid advance time.
- Fully kinetic ion model (with fluid electrons) also exists.


## Higher-Order Elements

(H. Strauss, J. Chen)

- $2^{\text {nd }}$ and $3^{\text {rdd }}$ order polynomial elements are available.
- Formed by adding nodes to existing mesh triangles.
- In "lumped" elements, nodes are placed at quadrature points of integral, resulting in a diagonal mass matrix for much faster evaluation, at a cost of more vertices.


$2^{\text {nd }}$-order lumped



## Resistive Wall

## (H. Strauss, J. Breslau)

- OpenMP code uses external package to generate vacuum-region mesh extending M3D mesh out to wall.
- Mesh may exclude axis region (not shown) with internal boundary condition.
- MPI version can initialize from mesh+data file generated by OMP version.
- Vacuum region treated as low density, low temperature (high $\eta$ ) plasma.
- Boundary conditions on fields at wall are applied using Green's functions precomputed by GRIN for each toroidal mode based on boundary geometry.



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## Concluding Thoughts

- The proliferation of new physics modules, options, and numerical techniques has made M3D very versatile and flexible but also very complex and challenging to maintain.
- A set of thorough standard tests for validation is badly needed.
- The code has been very productive on present machines, producing results few other MHD codes are capable of.
- But it could be a lot more efficient, and scaling up usefully to petascale runs remains a formidable hurdle.
- Need more implicitness.
- Need higher order elements.
- Need efficient, scalable solvers.

