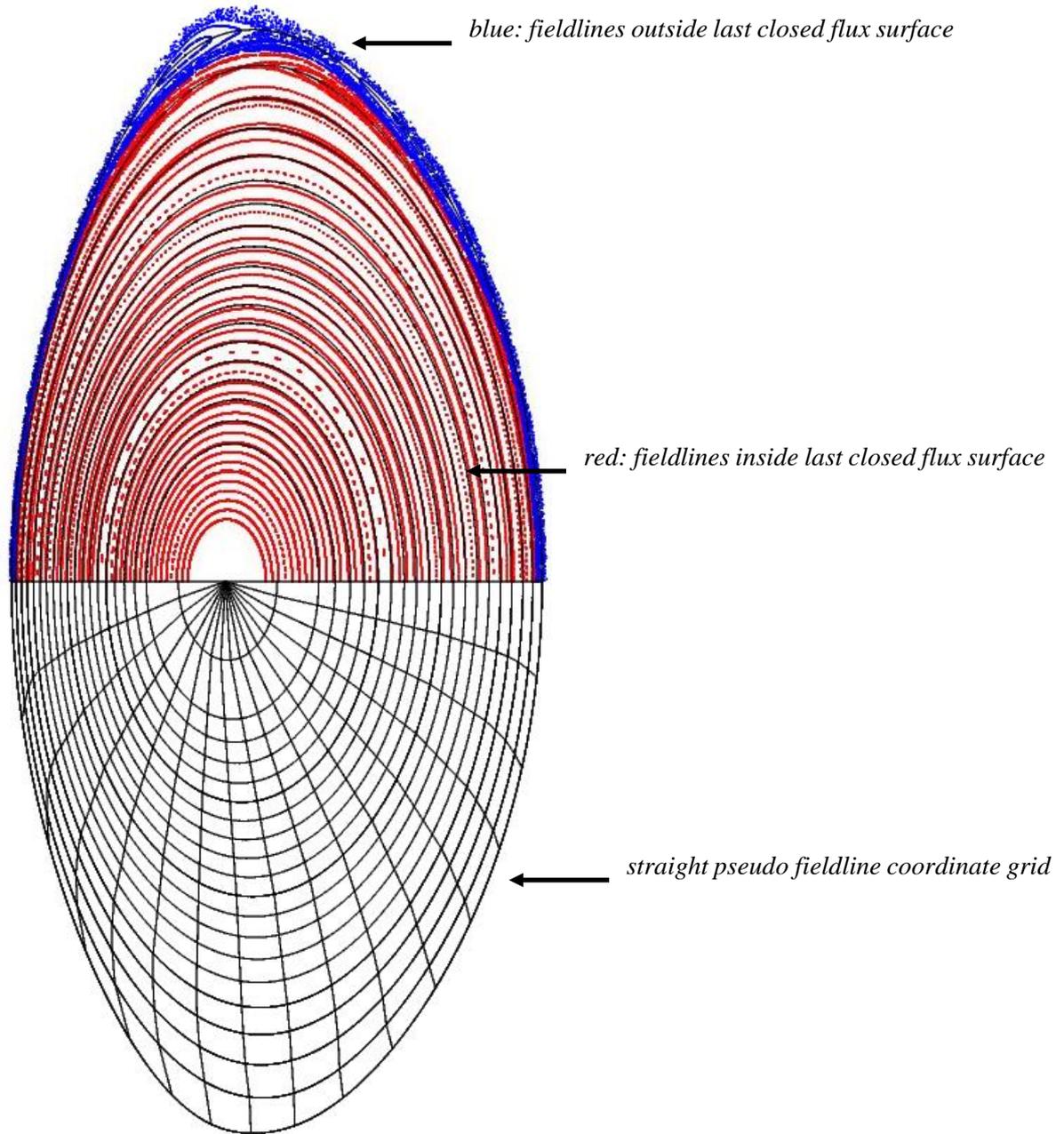
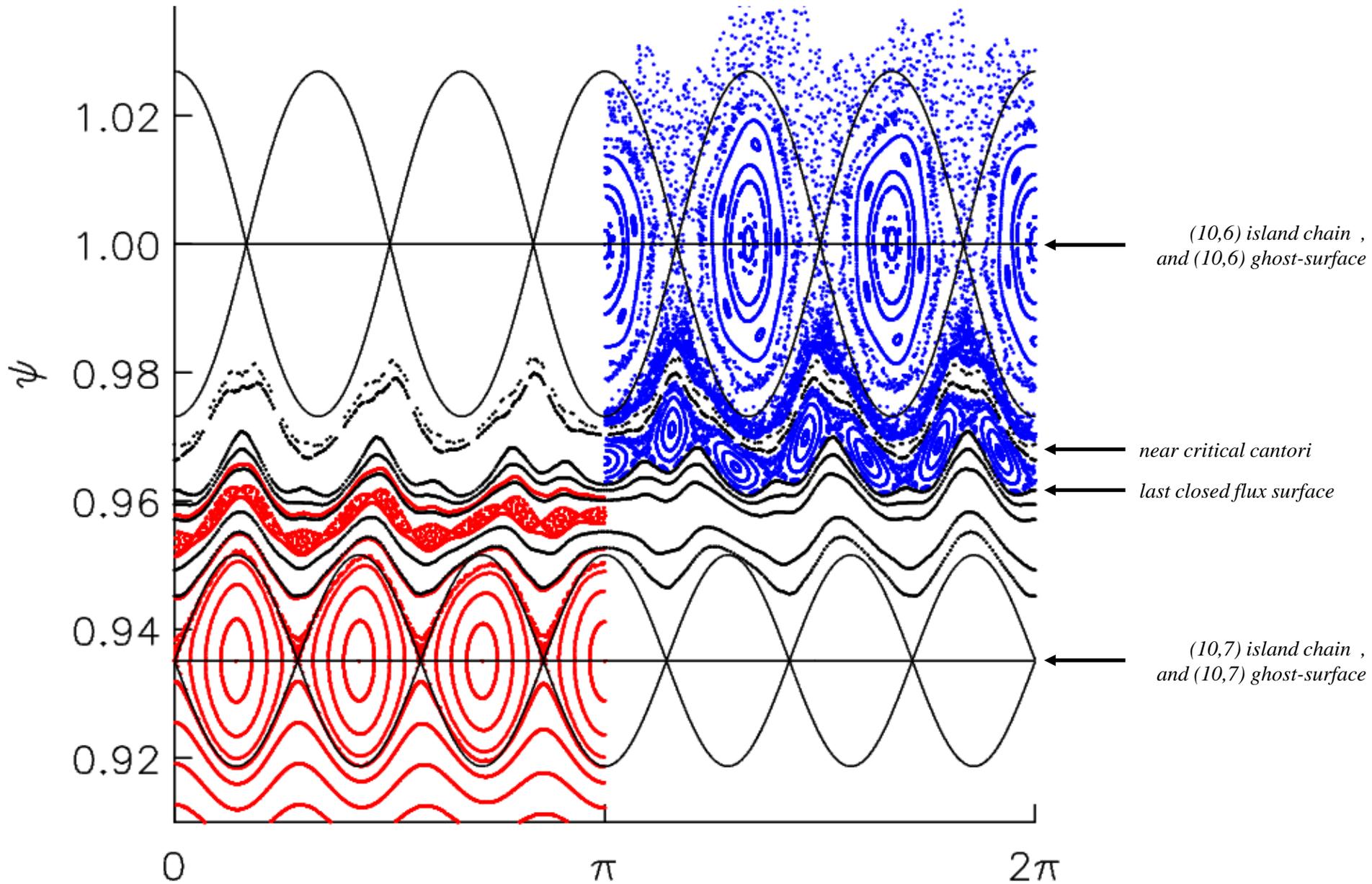


An examination of the chaotic edge of LHD



An examination of the chaotic edge of LHD

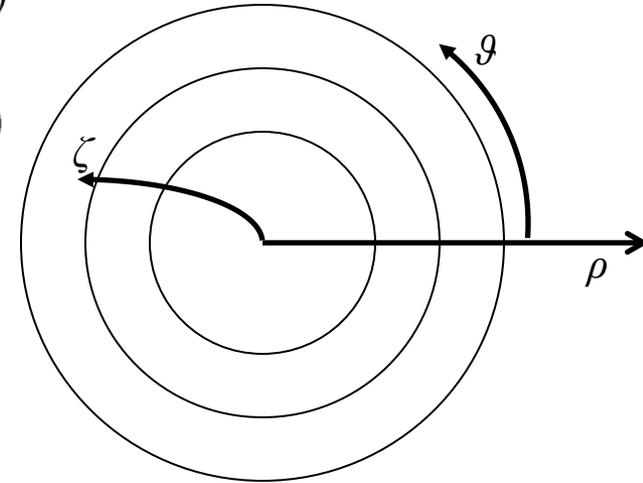


The magnetic field is given in cylindrical coordinates, and arbitrary, toroidal coordinates are introduced.

$$\begin{aligned} R &= R(\rho, \theta, \zeta) = \sum_{m,n} R_{m,n}(\rho) \cos(m\theta - n\zeta) \\ \phi &= \zeta \\ Z &= Z(\rho, \theta, \zeta) = \sum_{m,n} Z_{m,n}(\rho) \sin(m\theta - n\zeta) \end{aligned}$$

$$\mathbf{B} = B^R \mathbf{e}_R + B^\phi \mathbf{e}_\phi + B^Z \mathbf{e}_Z = B^\rho \mathbf{e}_\rho + B^\theta \mathbf{e}_\theta + B^\zeta \mathbf{e}_\zeta$$

$$\begin{pmatrix} B^R \\ B^\phi \\ B^Z \end{pmatrix} = \begin{pmatrix} R_\rho & R_\theta & R_\zeta \\ \phi_\rho & \phi_\theta & \phi_\zeta \\ Z_\rho & Z_\theta & Z_\zeta \end{pmatrix} \begin{pmatrix} B^\rho \\ B^\theta \\ B^\zeta \end{pmatrix}$$



Begin with circular cross section coordinates, centered on the magnetic axis.

In practice, we will have a discrete set of toroidal surfaces that will be used as “coordinate surfaces”.

The Fourier harmonics, $R_{m,n}$ & $Z_{m,n}$, of a discrete set of toroidal surfaces are interpolated using piecewise cubic polynomials.

If the surfaces are smooth and well separated, this “simple-minded” interpolation works.

A regularization factor is introduced, e.g. $R_{m,n}(\rho) = \rho^{m/2} \bar{X}_{m,n}(\rho) + R_{m,n}(0)$ to ensure that the interpolated surfaces do not overlap near the coordinate origin=magnetic axis.

A magnetic vector potential, in a suitable gauge, is quickly determined by radial integration.

$$\mathbf{A} = A_\theta(\rho, \theta, \zeta)\nabla\theta + A_\zeta(\rho, \theta, \zeta)\nabla\zeta$$

$$\sqrt{g}B^\rho = \partial_\theta A_\zeta - \partial_\zeta A_\theta$$

$$\sqrt{g}B^\theta = -\partial_\rho A_\zeta$$

$$\sqrt{g}B^\zeta = \partial_\rho A_\theta$$

$$\partial_\rho A_{\theta,m,n} = (\sqrt{g}B^\zeta)_{m,n}$$

$$\partial_\rho A_{\zeta,m,n} = -(\sqrt{g}B^\theta)_{m,n}$$

hereafter, we will use the commonly used notation

$$\mathbf{A} \equiv \psi\nabla\theta - \chi\nabla\zeta$$

ψ is the toroidal flux, and χ is called the magnetic field-line Hamiltonian

The magnetic field-line **action** is the

$$S \equiv \int_C \mathbf{A} \cdot d\mathbf{l}$$

line integral of the vector potential

C is an arbitrary “trial” curve

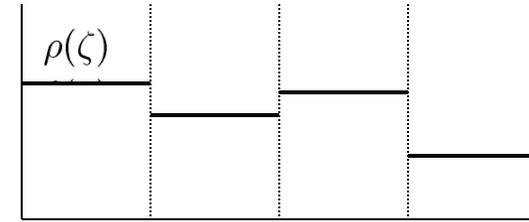
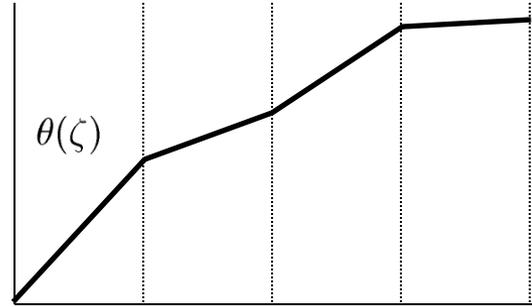
piecewise-constant, piecewise-linear

For $\zeta \in (\zeta_{i-1}, \zeta_i)$

$$\rho(\zeta) = \rho_i$$

$$\theta(\zeta) = \theta_{i-1} + \dot{\theta} (\zeta - \zeta_{i-1})$$

where $\dot{\theta} \equiv (\theta_i - \theta_{i-1})/\Delta\zeta$ is constant,



$$S \equiv \sum_{i=1}^N \int_{\zeta_{i-1}}^{\zeta_i} d\zeta \mathbf{A} \cdot d\mathbf{l} \equiv \sum_{i=1}^N \sum_{m,n} \left[\psi_{mn}(\rho_i) \dot{\theta} - \chi_{mn}(\rho_i) \right] \int_{\zeta_{i-1}}^{\zeta_i} d\zeta \cos(m\theta - n\zeta)$$

$$\int_{\zeta_{i-1}}^{\zeta_i} d\zeta \cos(m\theta - n\zeta) = \frac{\sin(m\theta_i - n\zeta_i) - \sin(m\theta_{i-1} - n\zeta_{i-1})}{m\dot{\theta} - n}$$

the piecewise-linear approximation allows
the cosine integral to be evaluated analytically,
i.e. method is FAST

To find extremizing curves, use Newton method to set $\partial_\rho S=0$, $\partial_\theta S=0$

$$\frac{\partial S}{\partial \rho_i} = 0 \quad \text{reduces to} \quad \frac{\partial S_i}{\partial \rho_i} = 0, \quad \text{which can be solved locally, } \rho_i = \rho_i(\theta_{i-1}, \theta_i)$$

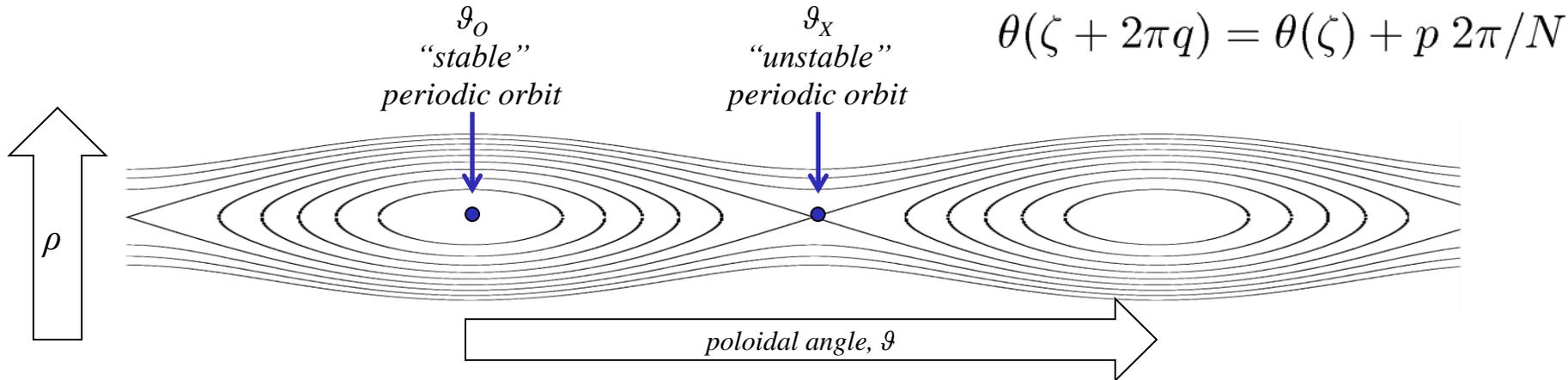
$$\frac{\partial S}{\partial \theta_i} = \partial_2 S_i(\theta_{i-1}, \theta_i) + \partial_1 S_{i+1}(\theta_i, \theta_{i+1})$$

tridiagonal Hessian, inverted in $O(N)$ operations, *i.e. method is FAST*

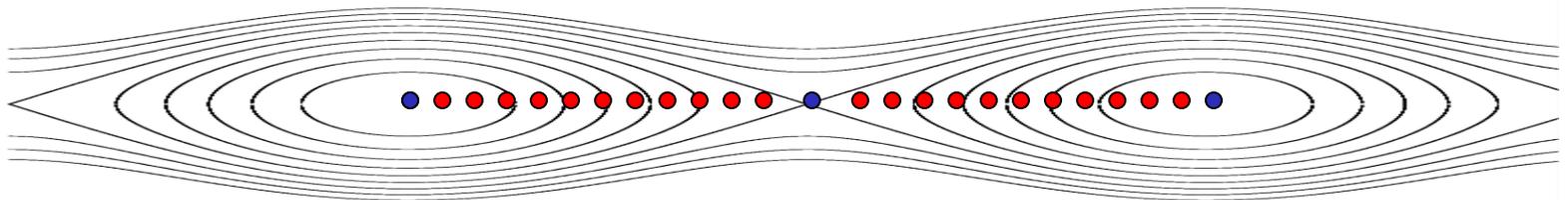
**Not required to follow magnetic field lines, and
does not depend on coordinate transformation.**

The trial-curve is constrained to be periodic, and a family of periodic curves is constructed.

Usually, there are only the “stable” periodic field-line and the “unstable” periodic field line,



However, we can “artificially” constrain the poloidal angle, i.e. $\vartheta(0) = \text{given constant}$, and search for extremizing periodic curve of the *constrained* action-integral $S \equiv \int_C \mathbf{A} \cdot d\mathbf{l} - \nu [\theta(0) - \theta_0]$



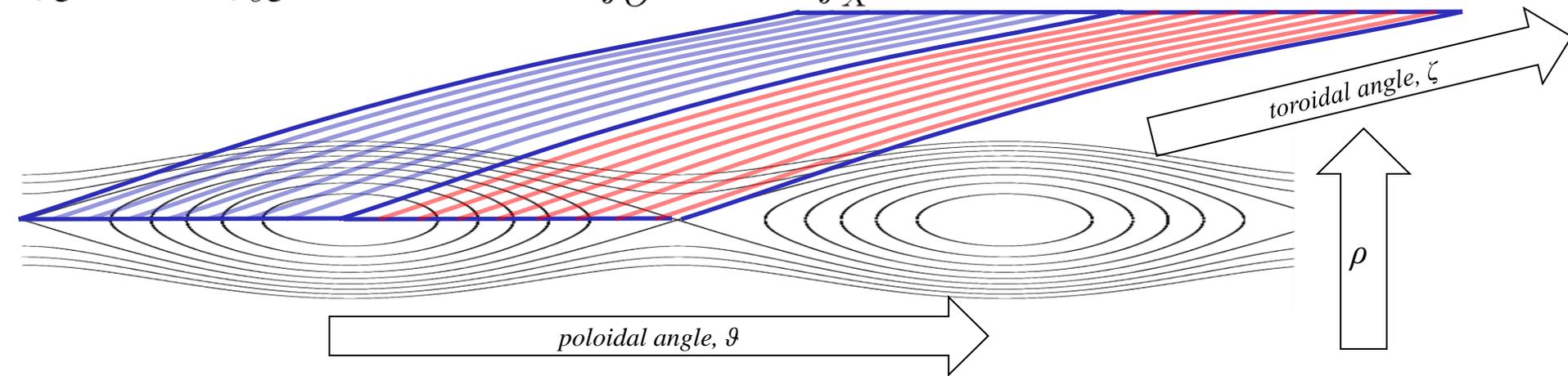
A rational, quadratic-flux minimizing surface
 is a family of periodic, extremal curves of the constrained action integral, and
is closely related to the rational ghost-surface,
 which is defined by an action-gradient flow between the minimax periodic orbit and the minimizing orbit.

$$\varphi_2 \equiv \frac{1}{2} \int_{\Gamma} w |B_n|^2 dS,$$

The “upward” flux = “downward” flux across a toroidal surface passing through an island chain can be computed.

$$\int_{\partial V} \mathbf{B} \cdot d\mathbf{S} \equiv \int_V \nabla \cdot \mathbf{B} = 0 \quad \text{the total flux across any closed surface of a divergence free field is zero.}$$

$$\int_S \mathbf{B} \cdot d\mathbf{s} \equiv \int_{\partial S} \mathbf{A} \cdot d\mathbf{l} \quad \Psi_{p/q} \equiv \int_O \mathbf{A} \cdot d\mathbf{l} - \int_X \mathbf{A} \cdot d\mathbf{l}$$



consider a sequence of rationals, p/q , that approach an irrational,

If $\Psi_{p/q} \rightarrow 0$ as $p/q \rightarrow t$, then KAM surface exists

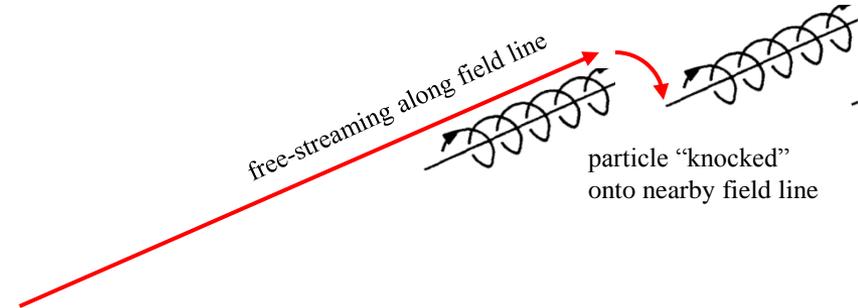
If $\Psi_{p/q} \rightarrow \Delta$, where $\Delta \neq 0$, then the KAM surface is “broken”, and $\Psi_{p/q}$ is the upward-flux across the cantorus

The diagnostics include:

1. **Greene's residue criterion:** the existence of an irrational surface can be determined by calculating the stability of nearby periodic orbits.
2. **Chirikov island overlap:** flux surfaces are destroyed when magnetic islands overlap.
3. **Cantori:** can present effective, partial barriers to fieldline transport, and cantori can be approximated by high-order periodic orbits.

The construction of chaotic coordinates simplifies anisotropic diffusion

$$\frac{\partial T}{\partial t} = \nabla \cdot (\kappa_{\parallel} \nabla_{\parallel} T + \kappa_{\perp} \nabla_{\perp} T) + Q,$$



In chaotic coordinates, the temperature becomes a surface function, $T=T(s)$, where s labels invariant (flux) surfaces or almost-invariant surfaces.

If $T=T(s)$, the anisotropic diffusion equation can be solved analytically, $\frac{dT}{ds} = \frac{c}{\kappa_{\parallel} \varphi + \kappa_{\perp} G}$,

where c is a constant, and

$\varphi = \int \int d\theta d\phi \sqrt{g} B_n^2$, is related to the quadratic-flux across an invariant or almost-invariant surface,

$G = \int \int d\theta d\phi \sqrt{g} g^{ss}$, is a geometric coefficient.

An expression for the temperature gradient in chaotic fields

S.R. Hudson, Physics of Plasmas, 16:010701, 2009

Temperature contours and ghost-surfaces for chaotic magnetic fields

S.R.Hudson and J.Breslau

Physical Review Letters, 100:095001, 2008

When the upward-flux is sufficiently small, so that the parallel diffusion across an almost-invariant surface is comparable to the perpendicular diffusion, the plasma cannot distinguish between a perfect invariant surface and an almost invariant surface

List of publications, <http://w3.pppl.gov/~shudson/>

Generalized action-angle coordinates defined on island chains

R.L.Dewar, S.R.Hudson and A.M.Gibson

Plasma Physics and Controlled Fusion, 55:014004, 2013

Unified theory of Ghost and Quadratic-Flux-Minimizing Surfaces

Robert L.Dewar, Stuart R.Hudson and Ashley M.Gibson

Journal of Plasma and Fusion Research SERIES, 9:487, 2010

Are ghost surfaces quadratic-flux-minimizing?

S.R.Hudson and R.L.Dewar

Physics Letters A, 373(48):4409, 2009

An expression for the temperature gradient in chaotic fields

S.R.Hudson

Physics of Plasmas, 16:010701, 2009

Temperature contours and ghost-surfaces for chaotic magnetic fields

S.R.Hudson and J.Breslau

Physical Review Letters, 100:095001, 2008

Calculation of cantori for Hamiltonian flows

S.R.Hudson

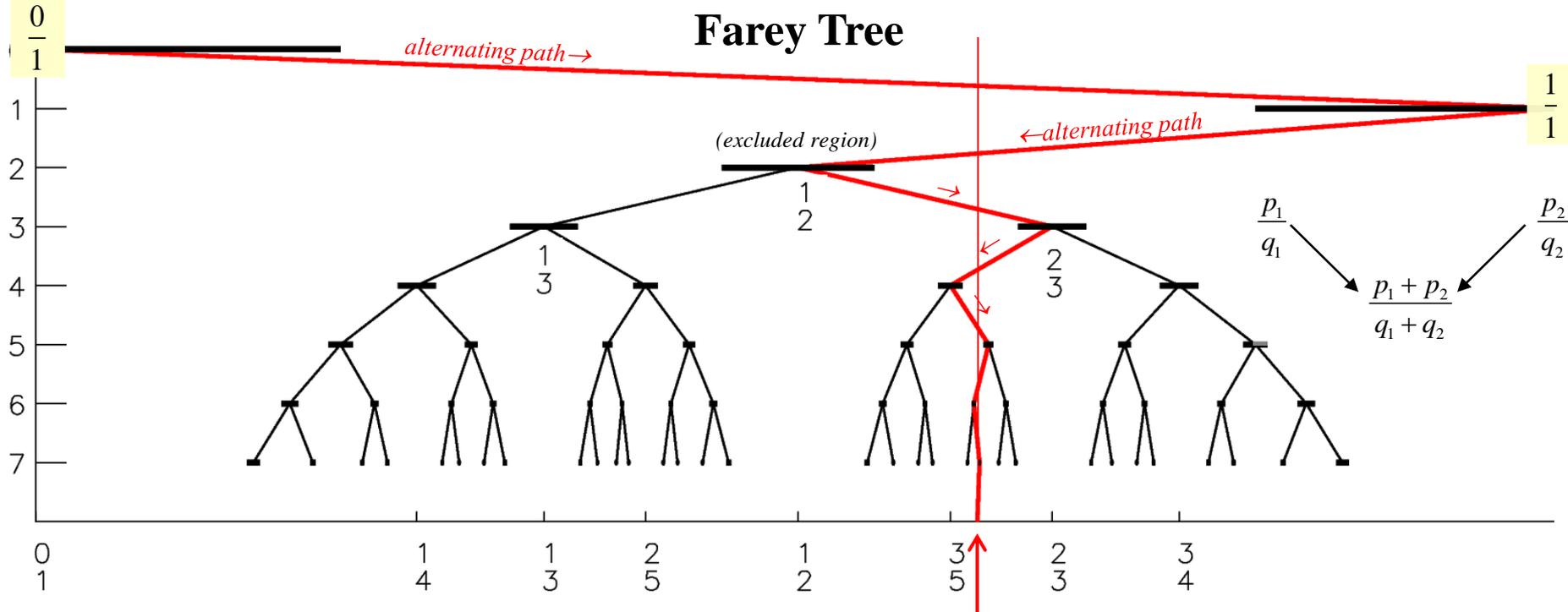
Physical Review E, 74:056203, 2006

Almost invariant manifolds for divergence free fields

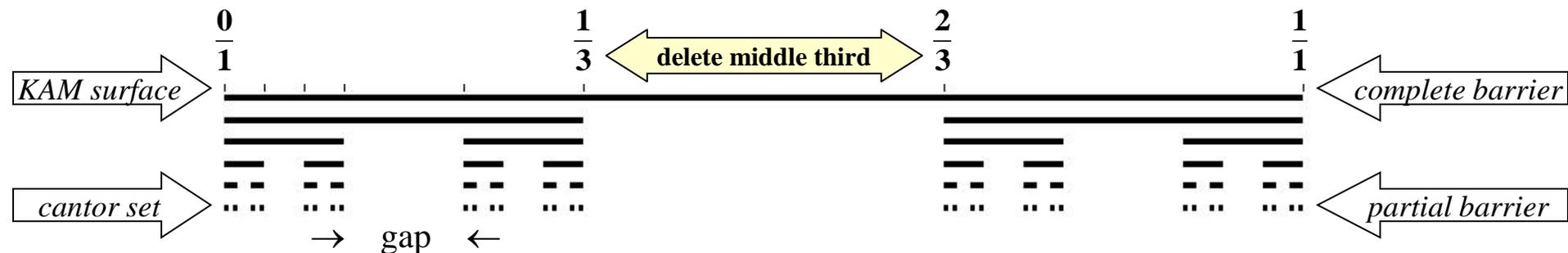
R.L.Dewar, S.R.Hudson and P.Price

Physics Letters A, 194(1-2):49, 1994

The fractal structure of chaos is related to the structure of numbers

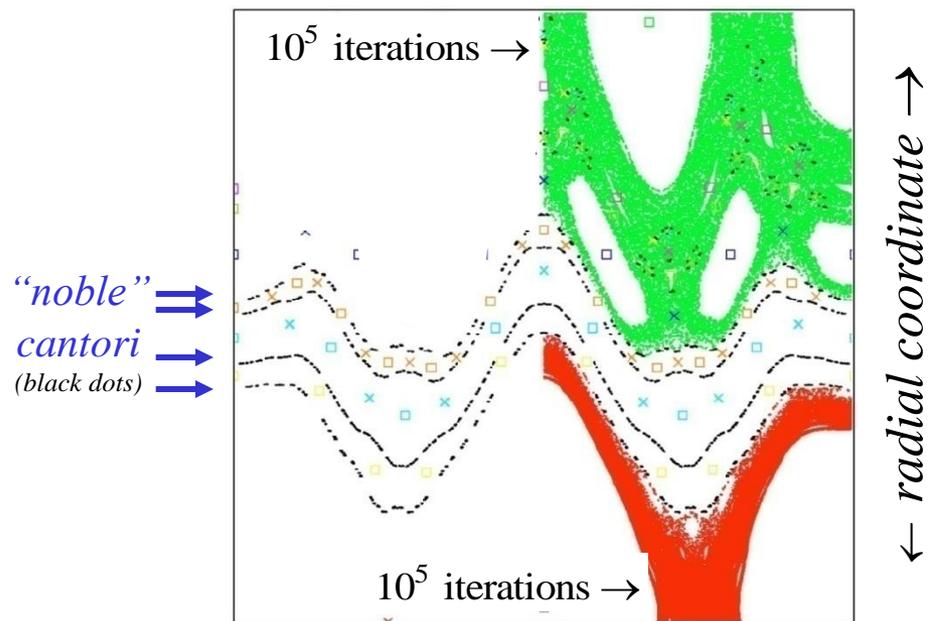


For non-integrable fields, field line transport is restricted by KAM surfaces and cantori



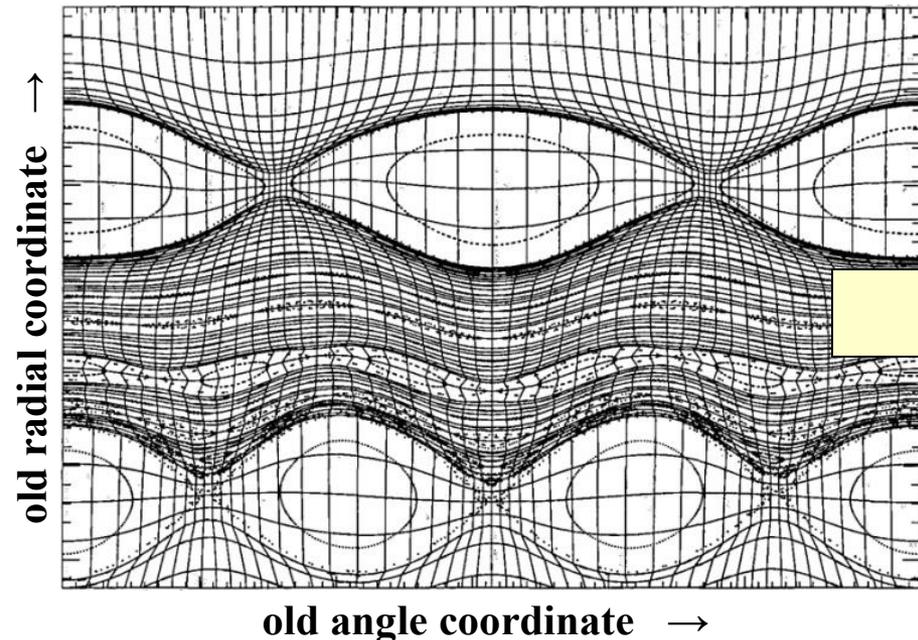
Calculation of cantori for Hamiltonian flows
S.R. Hudson, Physical Review E 74:056203, 2006

- *KAM surfaces are closed, toroidal surfaces that **stop** radial field line transport*
- *Cantori have “holes” or “gaps”; but cantori can severely “slow down” radial field line transport*
- *Example: all flux surfaces destroyed by chaos, but even after **100 000 transits** around torus the field lines cannot get past cantori*

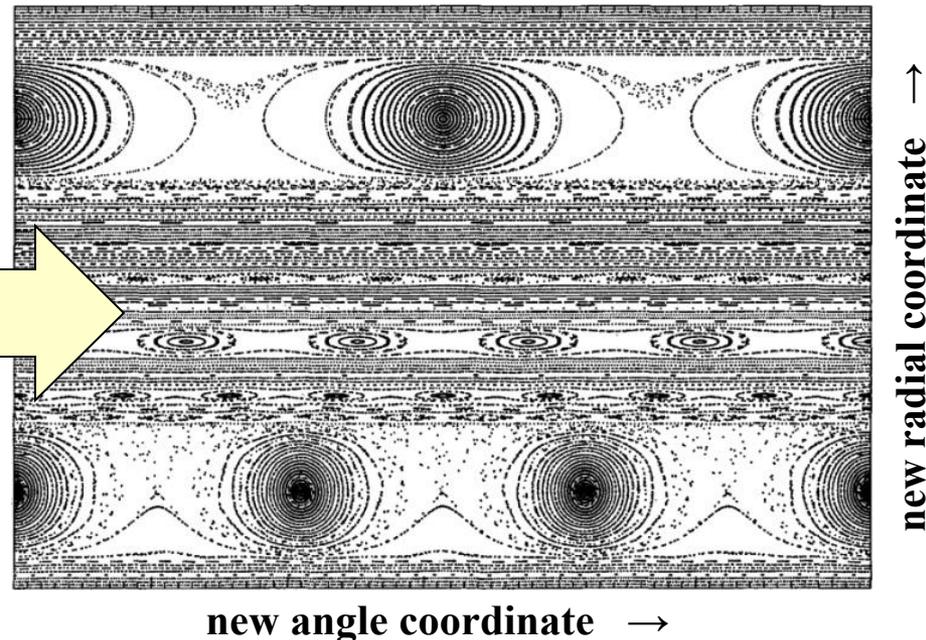


Chaotic coordinates “straighten out” chaos

Poincaré plot of chaotic field
(in **action-angle** coordinates of **unperturbed** field)



Poincaré plot of chaotic field
in **chaotic** coordinates



phase-space is partitioned into (1) **regular (“irrational”) regions**
and (2) **irregular (“rational”) regions**

with “good flux surfaces”, temperature gradients
with islands and chaos, flat profiles

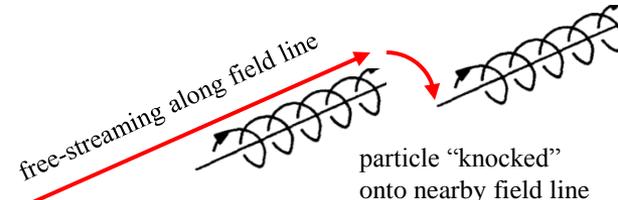
Generalized magnetic coordinates for toroidal magnetic fields

S.R. Hudson, Doctoral Thesis, The Australian National University, 1996

Chaotic coordinates simplify anisotropic transport

The temperature is constant on ghost surfaces, $T=T(s)$

1. Transport *along* the magnetic field is *unrestricted*
 → consider parallel random walk, with **long** steps \approx collisional mean free path
2. Transport *across* the magnetic field is *very small*
 → consider perpendicular random walk with **short** steps \approx Larmor radius



3. Anisotropic diffusion balance $\kappa_{\parallel} \nabla_{\parallel}^2 T + \kappa_{\perp} \nabla_{\perp}^2 T = 0$, $\kappa_{\parallel} \gg \kappa_{\perp}$, $\kappa_{\perp} / \kappa_{\parallel} \sim 10^{-10}$

$2^{12} \times 2^{12} = 4096 \times 4096$ grid points
(to resolve small structures)

4. Compare solution of numerical calculation to ghost-surfaces

5. The temperature adapts to KAM surfaces, cantori, and ghost-surfaces!

i.e. $T=T(s)$, where $s=const.$ is a ghost-surface

from $T=T(s, \theta, \phi)$ to $T=T(s)$ is a fantastic simplification, allows analytic solution

$$\frac{dT}{ds} \propto \frac{1}{\kappa_{\parallel} \phi_2 + \kappa_{\perp} G}$$

Temperature contours and ghost-surfaces for chaotic magnetic fields
 S.R. Hudson et al., Physical Review Letters, 100:095001, 2008
 Invited talk 22nd IAEA Fusion Energy Conference, 2008
 Invited talk 17th International Stellarator, Heliotron Workshop, 2009

An expression for the temperature gradient in chaotic fields
 S.R. Hudson, Physics of Plasmas, 16:100701, 2009

