

Potential for Astrophysics Experiments in the Laboratory Using High- Brightness Particle Beams

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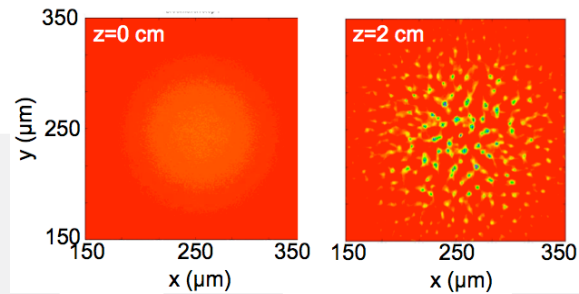
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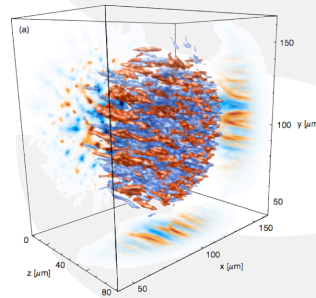
POSSIBLE EXPERIMENTS

1. Current Filamentation Instability (CFI) of the BNL-ATF e-beam



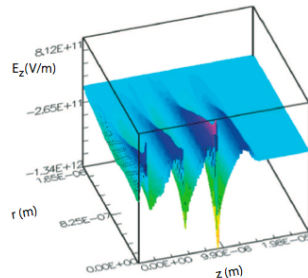
Experiment under way

2. CFI instability of the SLAC e-/e⁺ “fireball beam”



To be proposed

3. SLAC-LCLS fs e-beam for TV/m beam/plasma interaction



To be proposed

MOTIVATION

➤ Why is Current filamentation instability (CFI) important?

- Relativistic outflows are ubiquitous in astrophysics
- CFI likely plays an important role in flow/ambient plasmas interaction
- Relatively few experiments to test CFI
- CFI also has important laboratory applications (ICF)

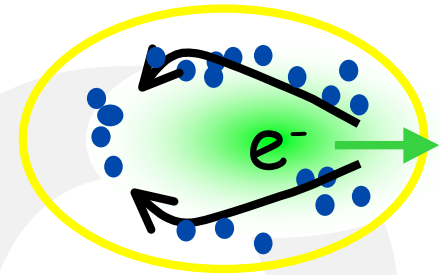
➤ Plan:

- Take advantage of PWFA experiments/experience
- Develop new and appropriate diagnostics
- Measure CFI-quantities (σ_r , Δ , B , δn_e , γ , Γ , ...)
- Set up strong experimental/simulation synergy

TWO REGIMES OF BEAM-PLASMA INTERACTION

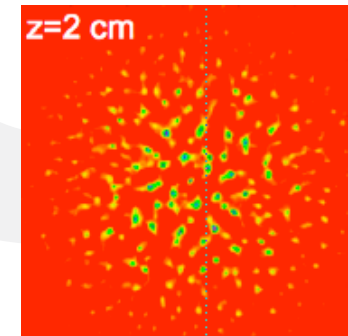
1) $\sigma_r \ll c / \omega_{pe}$

- Return current outside of the bunch
- Plasma wakefield regime $\sigma_z \cong c / \omega_{pe}$
- Blow-out regime $n_b \geq n_{pe}$
- Large decelerating, accelerating, focusing fields
- Hose instability

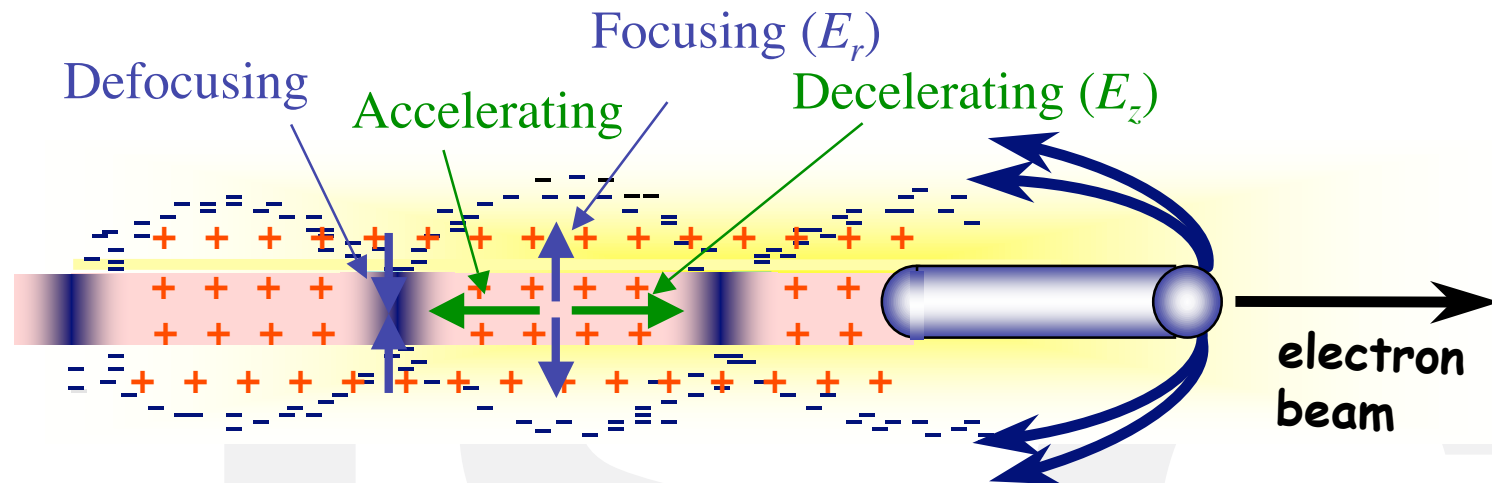


2) $\sigma_r \gg c / \omega_{pe}$

- Return current outside of the bunch
- CFI, two-stream, oblique instabilities



PLASMA WAKEFIELD (e^-)

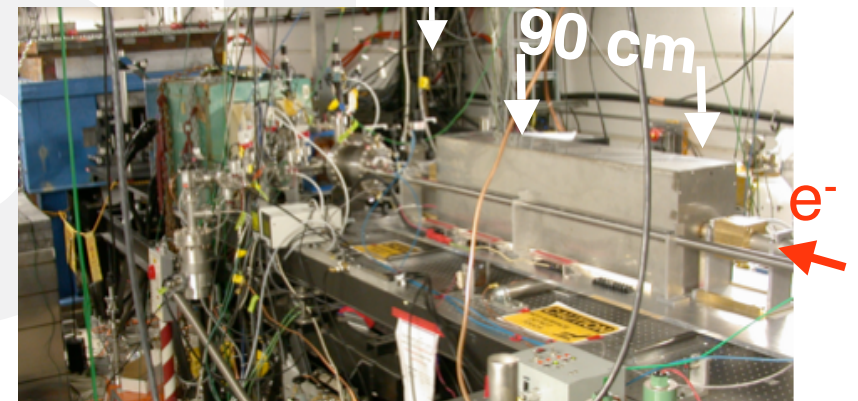
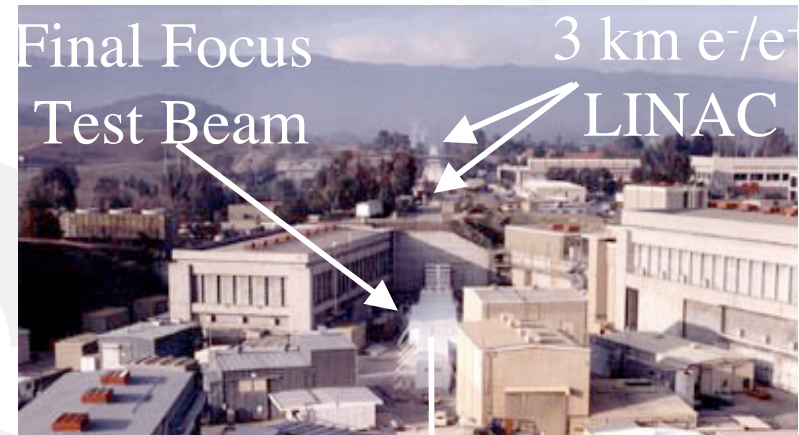
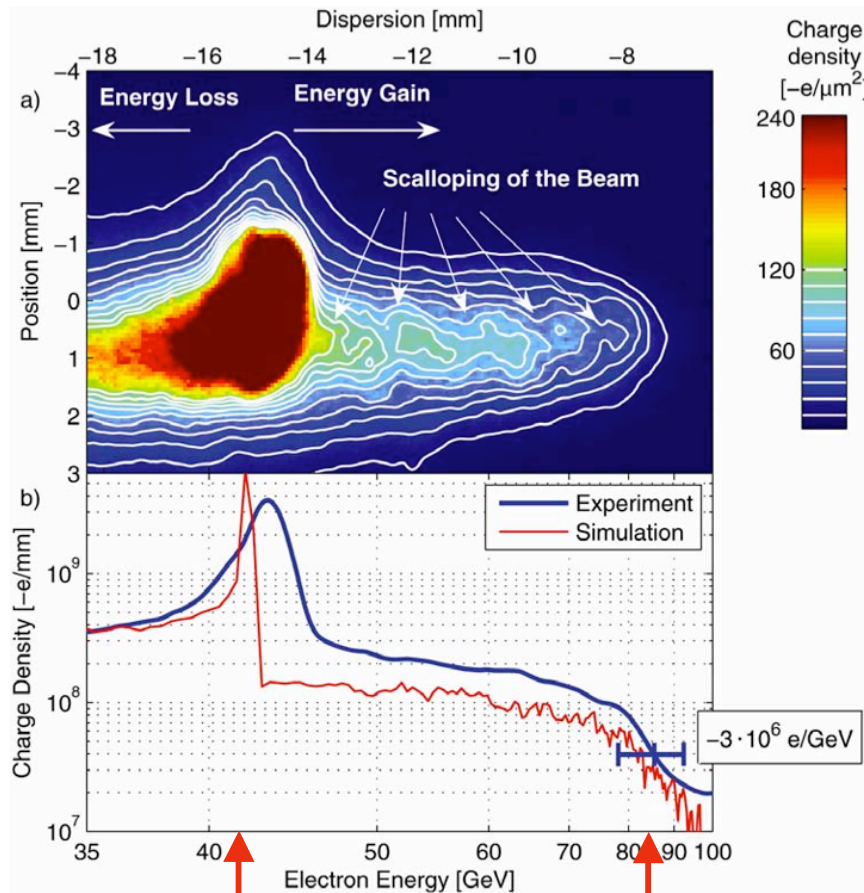


- ➔ Plasma wave/wake excited by a relativistic particle bunch
- ➔ Plasma e^- expelled by space charge forces => energy loss + focusing
- ➔ Plasma e^- rush back on axis => energy gain

e⁻ ENERGY DOUBLING

I. Blumenfeld *et al.*, Nature 445, 2007

$E_0 = 42 \text{ GeV}$



➡ Energy doubling of e⁻ over $L_p \approx 85 \text{ cm}$, $2.7 \times 10^{17} \text{ cm}^{-3}$ plasma

➡ Unloaded gradient $\approx 52 \text{ GV/m}$ ($\approx 150 \text{ pC accel.}$)

CFI @BNL-ATF

- Regime 2: $\sigma_r \gg c/\omega_{pe}$ ($k_p \sigma_r \gg 1$), Current Filamentation Instability
 - Plasma return current flows inside beam
 - Dominant instability dependent on relativistic beam factor (γ_0)

$\gamma_0 \gg 1$ - CFI (transverse instability)

- Filament size and spacing $\sim c/\omega_{pe}$
- Growth rate⁽¹⁾:

$$\Gamma = \beta_0 \sqrt{\frac{\alpha}{\gamma_0}} \omega_{pe} \quad \text{or} \quad \Gamma = \beta_0 \omega_{pb} / \sqrt{\gamma_0} \sim n_b \sim Q / (\sigma_r^2 \sigma_z)$$

Ratio of beam to plasma density: $\alpha = n_b / n_e$

BNL-ATF Beam

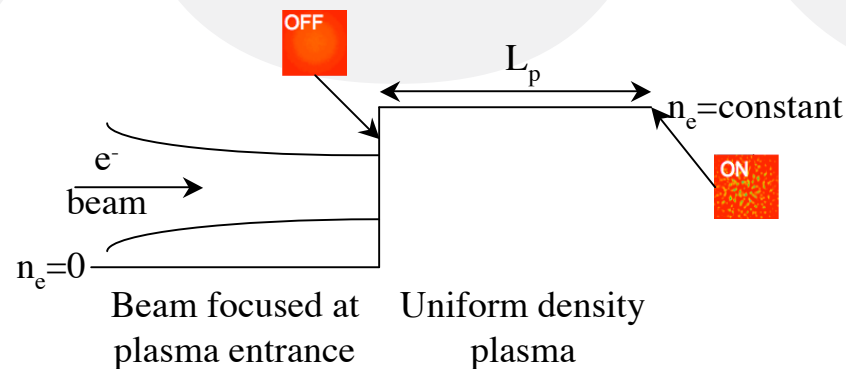
ATF Over-compressed Beam Parameters	
Parameter	Value
Charge (pC)	150 to 200
Beam Transverse Waist Size (μm)	100 to 200
Bunch Length (fs)	100
Beam Density (cm^{-3})	10^{14}
Energy (MeV)	59
Emittance (mm-mrad)	1 to 2

- $\gamma_0 = 117$, CFI regime
- Growth rate estimate:
 $\Gamma = 8.6 \times 10^{10} \text{ s}^{-1}$ or 3.5 mm at c

➤ CFI should be observable on a cm-length plasma scale

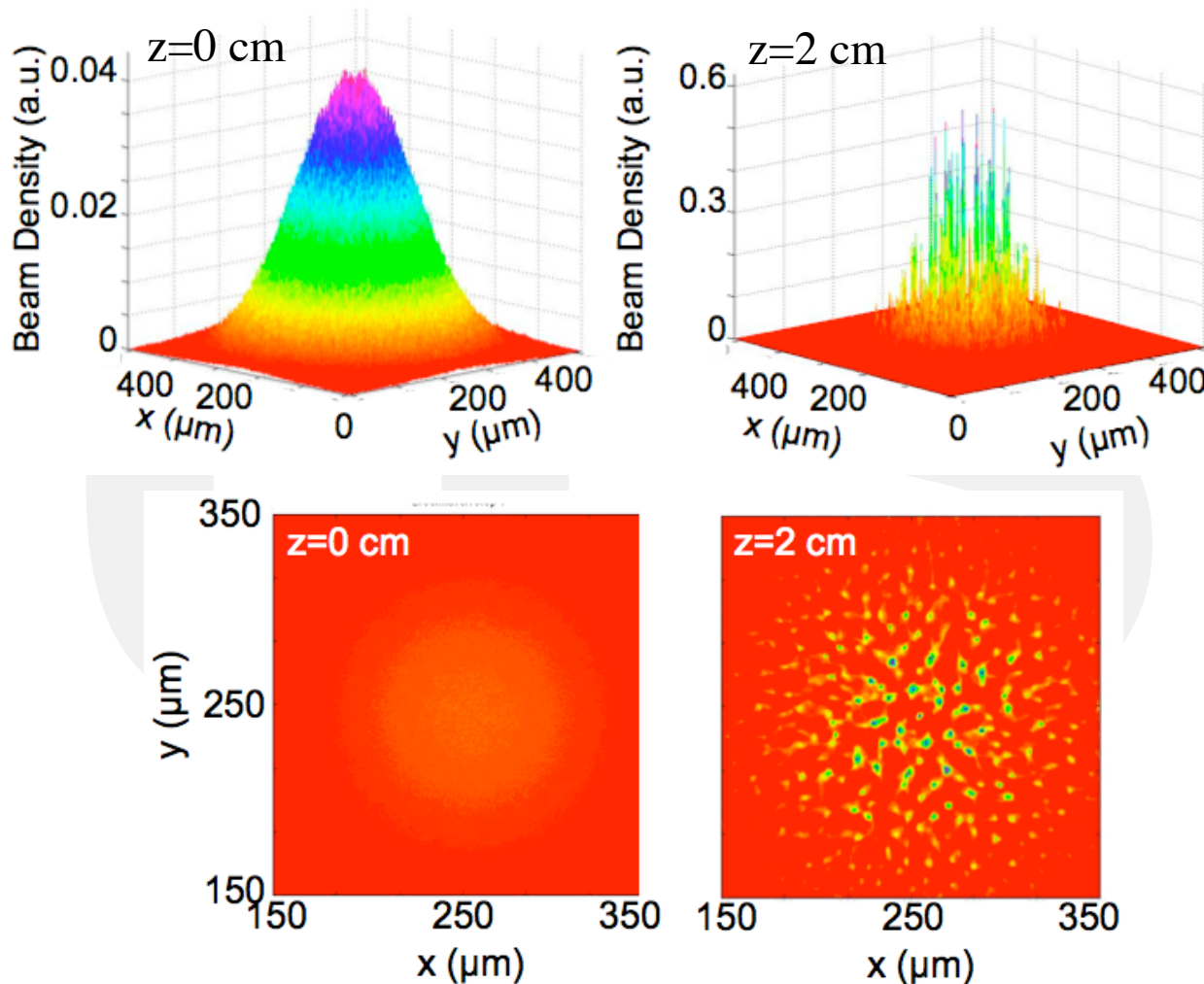
SIMULATIONS TOOLS

- Simulation Tools: Particle-in-cell codes
 - OSIRIS
 - QuickPIC
 - PWFA, ICF
 - Both codes have been benchmarked against other codes and with experimental results
- Simulation Setup



FILAMENTATION SIMULATION

Beam focused at the entrance of a uniform density plasma



ATF Beam and Plasma Simulation Parameters		
Parameter	Value	Value
Simulation Box - X	548 μm	512 cells
Simulation Box - Y	548 μm	512 cells
Simulation Box - Z	136 μm	128 cells
Plasma Particles/Cell	4.0	
3-D Time step (μm)	100	
Number Beam Particles -X	128	
Number Beam Particles -Y	128	
Number Beam Particles -Z	512	
Relativistic Factor	100	
Beam Transverse Waist	100	
Size (μm)		
Bunch Length (μm)	15	
Charge (pC)	200	
Plasma Density (cm^{-3})	5×10^{17}	
Capillary Length (cm)	2	
Skin depth (c/ω_{pe}) (μm)	7.5	

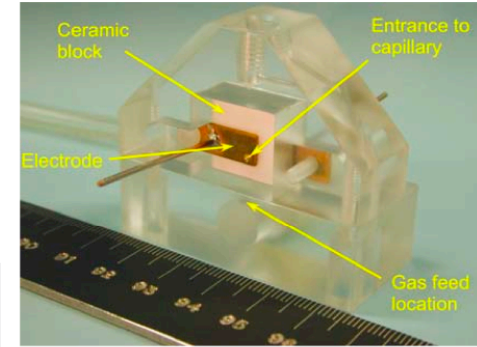
Filament size $\approx 4 \mu\text{m}$
 Filament spacing $\approx 20 \mu\text{m}$
 Both $\approx c/\omega_p$

These images are similar to those we expect to measure in the experiments

EXPERIMENTAL SETUP

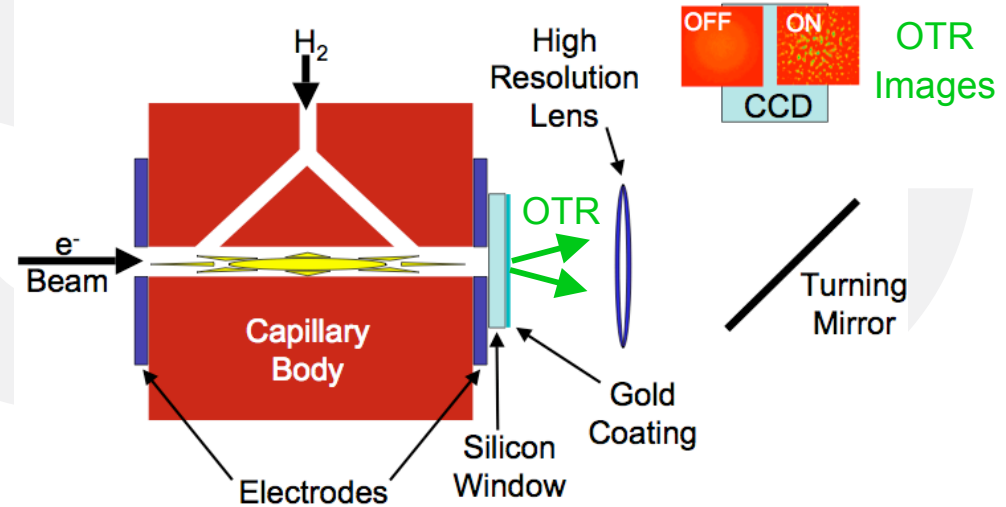
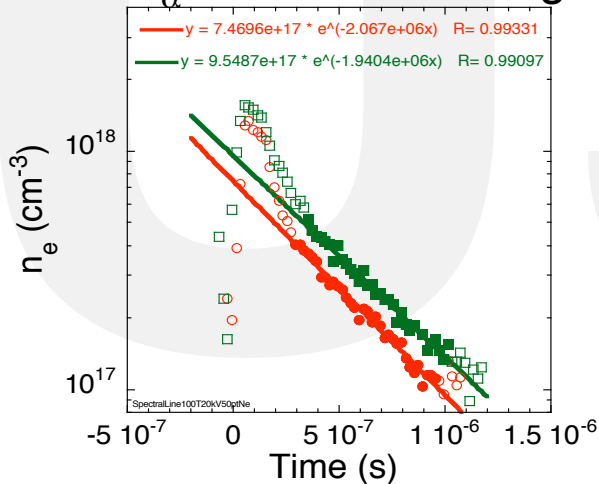
• Plasma Source

- H₂ capillary discharge, radius of ~ 500 μm
- Variable plasma parameters:
 - Capillary length (1cm and 2cm >L_G=3.5mm)
 - Plasma density: H₂ pressure and beam arrival time
- Variable beam parameters:
 - Charge, length, radius (n_b)



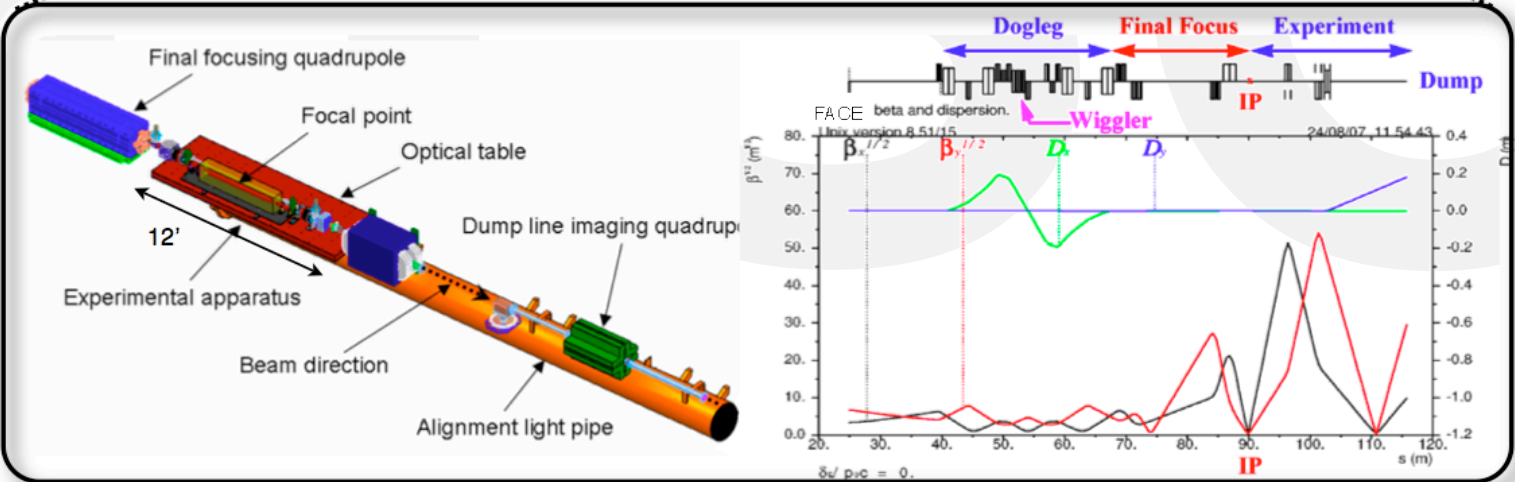
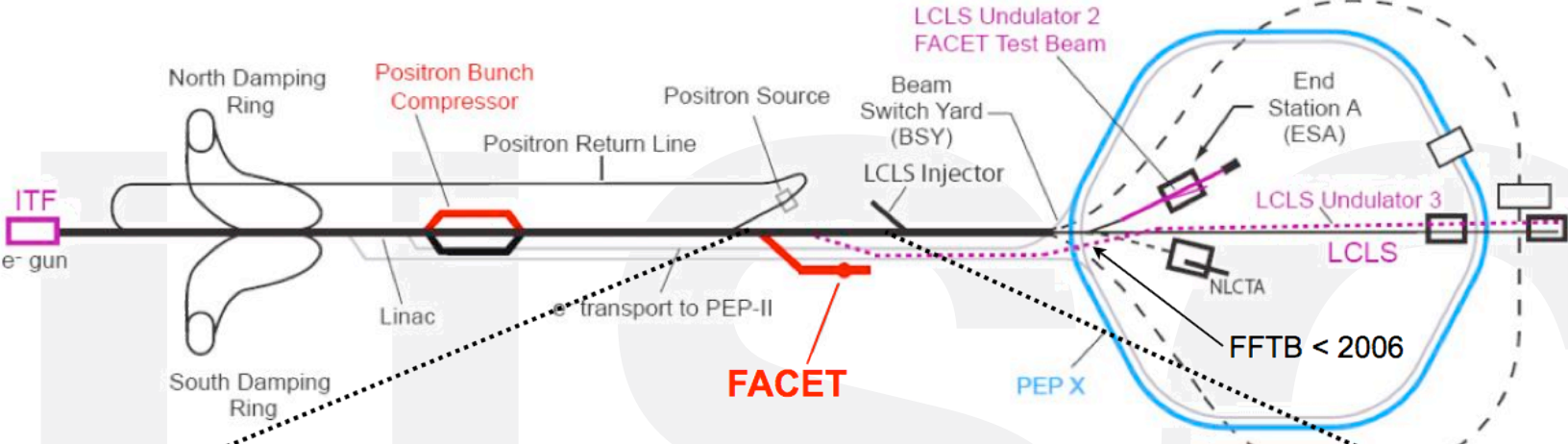
W.D. Kimura et. Al, AIP Conference Proceedings Volume 877, 534

H_α Stark Broadening



- ➡ Diagnostics: OTR for filaments, radiation
- ➡ Measure filaments size, spacing, growth rate, etc.
- ➡ Experiment approved at BNL-ATF, program funded by NSF

FACET is a new facility to provide high-energy, high peak current e^- & e^+ beams for PWFA experiments



NEUTRAL RELATIVISTIC BEAM

- Extract e- & e+ from damping rings on same linac pulse
- Accelerate bunches to sector 20 5cm apart
- Use 'Sailboat Chicane' to put them within 100 μ m at entrance to plasma

Energy	24 GeV
Charge	3 nC
Sigma z	17 μ m
Sigma r	< 10 μ m
Peak Current	22 kAmps
Species	e- & e+

$$\gamma \approx 5 \times 10^4$$

$$E_0 = 80 \text{ J}$$

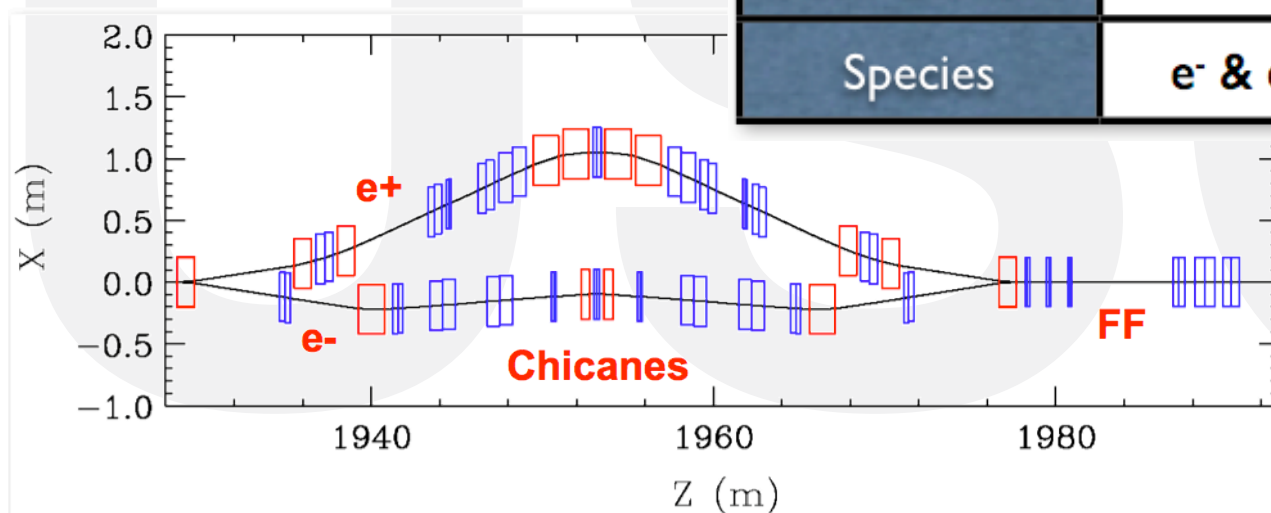
$$P = 1 \text{ PW}$$

$$I = 4.4 \times 10^{20} \text{ W/cm}^2$$

$$E_{r \text{ max}} = 61 \text{ GV/m}$$

$$B_{\theta \text{ max}} = 2 \text{ MG}$$

$$n_b = 7 \times 10^{17} \text{ cm}^{-3}$$

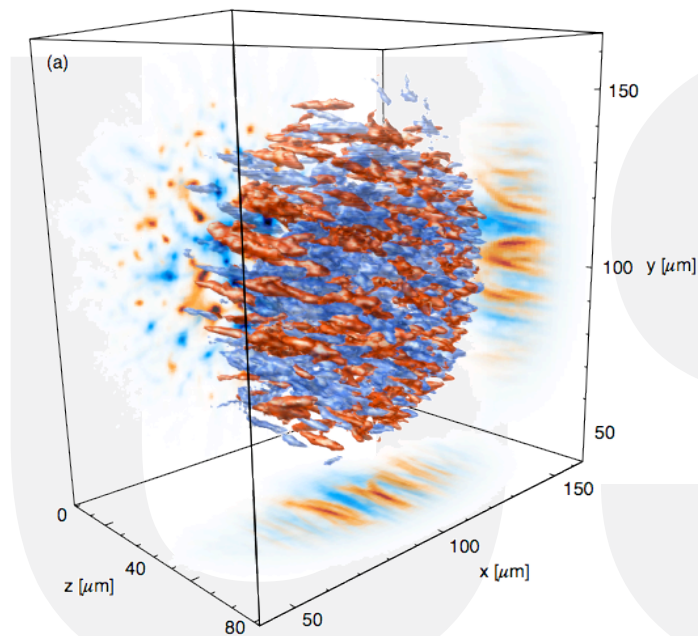


“Fireball Beam”:
Neutral
Ultra-relativistic
e-/e+ beam

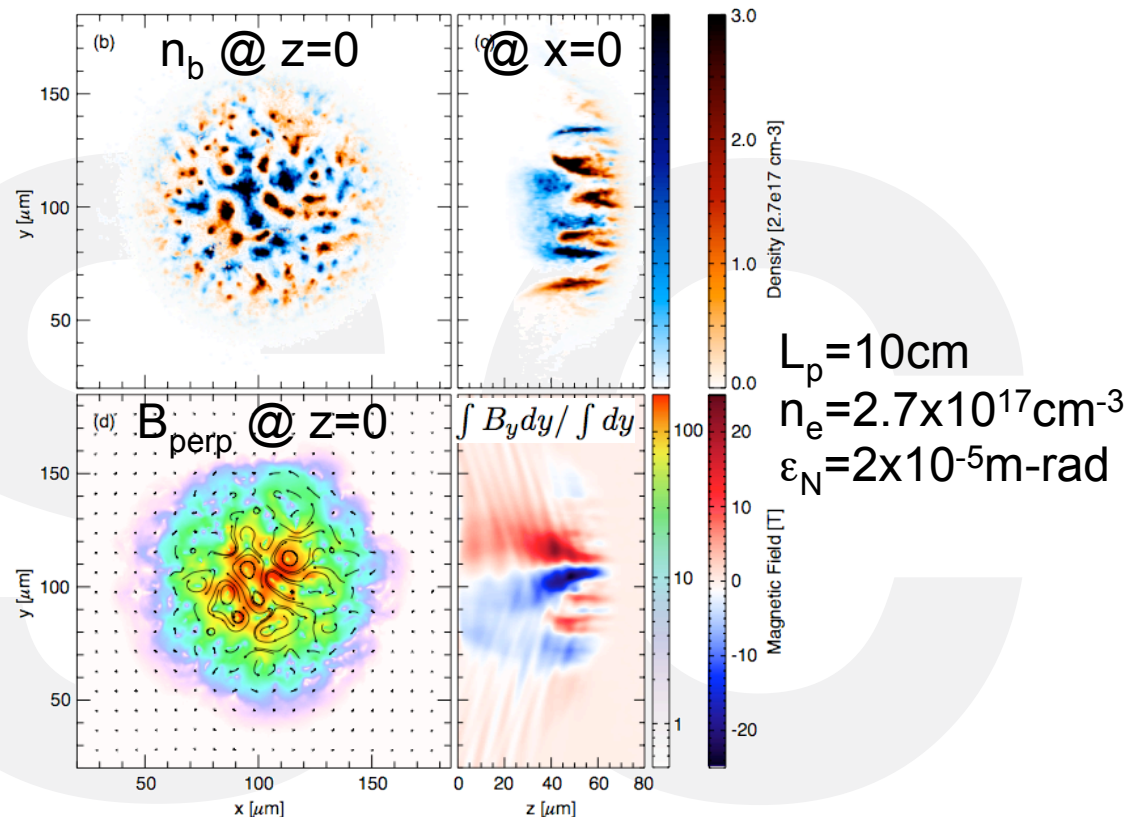
- ➔ Space-time overlap of equal e-/e+ for “fireball beam”
- ➔ Beam at 10Hz, very stable, Li plasma(?)
- ➔ Collision of relativistic “fireball beam” with lab plasma

SIMULATION: FILAMENTATION

Density plot: e^- & e^+



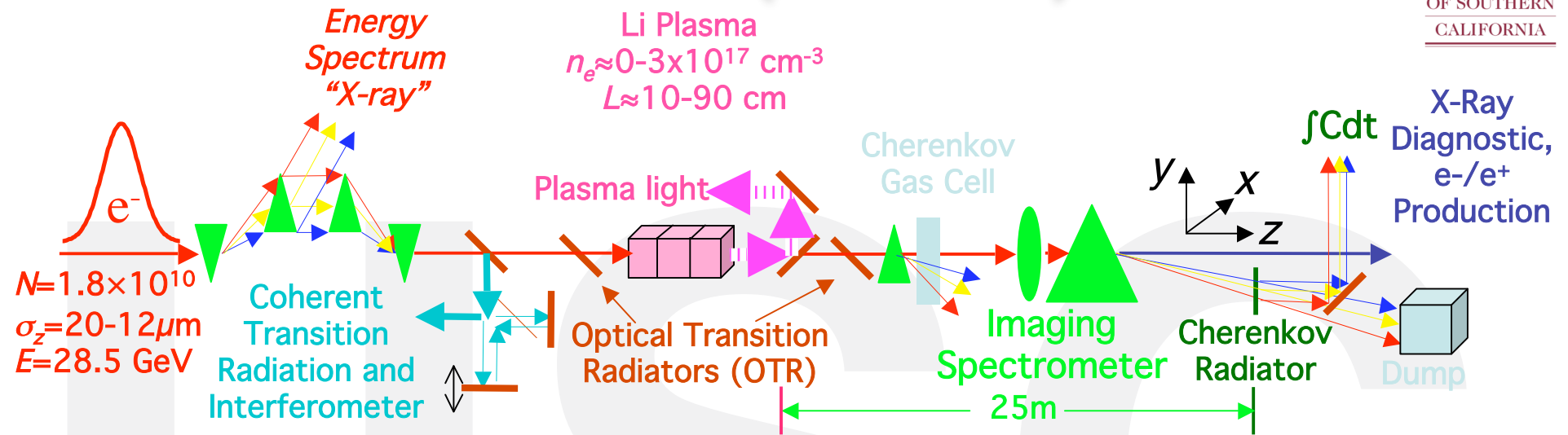
Simulations: S. Martins, OSIRIS



➡ “Fireball beam” CFI after 10cm of plasma!

➡ Same beam parameters as for PWFA, very different result with neutral beam

EXPERIMENTAL SET UP (GENERIC)



➔ Li-plasma source similar to PWFA experiments

➔ Diagnostics:

δn_e : shadowgraphy, Schlieren

B: Faraday rotation

Radiation

Energy loss

...

} as a fct of plasma length

SUMMARY

- Accelerators produce high quality particle beams that could be used for astrophysics-relevant experiments: opportunities!
- CFI of BNL-ATF e^- beam and SLAC e^-/e^+ ultra-relativistic “fireball beam”
- Experiments benefit from PWFA experience/setup
- Other experiment possible? Beam-driven relativistic shocks?
- Input welcome!

Thank you!