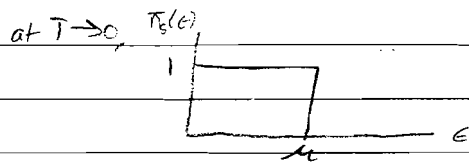


May 2003 #2 (SM)

a. degenerate electrons: Fermi-Dirac Distribution: $\bar{n}_s = \frac{1}{e^{\beta(\epsilon_s - \mu)} + 1}$ # of particles in a state with a given energy ϵ



all states filled up to Fermi energy μ

corresponds to some Fermi momentum $p_F = \hbar k_F$, so $\mu = \frac{\hbar^2 k_F^2}{2m}$

Volume of sphere in k -space: $\frac{4}{3}\pi k_F^3$

Density of states: $\rho(k)d^3k = \frac{V}{(2\pi)^3}$ # of translational states available between \vec{k} and $\vec{k} + d\vec{k}$ (independent of spin)

$$2 \cdot \frac{V}{(2\pi)^3} \cdot \left(\frac{4}{3}\pi k_F^3\right) = N$$

$$\left[\text{explicitly } 2 \int_0^\infty f_{FD}(\epsilon) \rho(\epsilon) d\epsilon = N \right]$$

$$k_F = \left(\frac{3\pi^2 N}{V} \right)^{1/3}$$

$$\mu = \frac{\hbar^2 k_F^2}{2m} = \frac{\hbar^2}{2m} \left(\frac{3\pi^2 N}{V} \right)^{2/3}$$

Total energy: $\sum_s n_s \epsilon_s \rightarrow 2 \int f_{FD}(\epsilon) \rho(\epsilon) \epsilon d\epsilon$

$$= 2 \int_0^\mu \epsilon \rho(\epsilon) d\epsilon$$

$$\rho(\epsilon) d\epsilon = \rho(k(\epsilon)) \left| \frac{dk}{d\epsilon} \right| d\epsilon$$

$$\rho(|\vec{k}|) dk = \frac{V}{(2\pi)^3} 4\pi k^2 dk = \frac{V}{2\pi^2} k^2 dk$$

$$k = \sqrt{\frac{2m}{\hbar^2}} \epsilon^{1/2} \quad \frac{dk}{d\epsilon} = \sqrt{\frac{m}{2\hbar^2}} \epsilon^{-1/2}$$

$$\rho(\epsilon) = \frac{V}{2\pi^2} \frac{2m\epsilon}{\hbar^2} \cdot \sqrt{\frac{m}{2\hbar^2}} \epsilon^{-1/2} = \frac{V m^{3/2}}{\sqrt{2}\pi^2 \hbar^3} \epsilon^{1/2}$$

$$\bar{E} = \frac{2 \cdot V m^{3/2}}{\sqrt{2}\pi^2 \hbar^3} \int_0^\mu \epsilon^{3/2} d\epsilon = \frac{\sqrt{2} V m^{3/2}}{\pi^2 \hbar^3} \mu^{5/2} \frac{2}{5} = \frac{2\sqrt{2} V m^{3/2}}{5\pi^2 \hbar^3} \frac{\hbar^5}{4 \cdot \sqrt{2} \cdot m^{5/2}} \left(\frac{3\pi^2 N}{V} \right)^{5/3}$$

$$\text{total } E = \frac{3N\hbar^2\pi^2}{10m_e} \left(\frac{3N}{\pi V} \right)^{2/3}$$

b. $U_{\text{grav}} = -\frac{3GM^2}{5R}$ binding energy for uniform density sphere

$$U_K \sim \frac{1}{2}v^2 \sim \frac{1}{R^2}$$

$$E = \frac{\alpha}{R^2} - \frac{\beta}{R} \quad \frac{dE}{dR} = \frac{-2\alpha}{R^3} + \frac{\beta}{R^2} = 0 \quad -2\alpha + \beta R = 0$$

$$\beta R = 2\alpha$$

$$R = \frac{2\alpha}{\beta}$$

$$E = \frac{\alpha \beta^2}{4\alpha^2} - \frac{\beta R}{2\alpha} = \frac{\beta^2}{4\alpha} - \frac{\beta^2}{2\alpha} = \frac{-\beta^2}{4\alpha} \quad \text{minimum energy}$$

minimizing E : eqb $R = \frac{2\alpha}{\beta}$

$$\alpha = \frac{3N(\hbar^2 \pi^2)}{10m_e} \cdot \frac{3^{2/3} N^{2/3}}{(\pi^{4/3} \pi)^{2/3}}$$

$$\beta = \frac{3GM^2}{5}$$

$$M = 2Nm_p \quad N = \frac{M}{2m_p}$$

$$R_{\text{eq}} = \frac{2 \cdot 3\pi^2 \hbar^2 (9)^{2/3}}{10m_e (4\pi^2)^{2/3}} \frac{M^{5/3}}{(2m_p)^{5/2}} = \frac{\pi^2 \hbar^2}{Gm_e (2m_p)^{5/2}} \frac{(9)^{2/3}}{(4\pi^2)^{2/3}} \frac{1}{M^{1/3}}$$

$$\frac{3GM^2}{5}$$

$$R \sim M^{-1/3}$$

c. highly relativistic: $E = cp \Rightarrow$ same k_f

$$\mu = ck k_f = c\hbar \left(\frac{3\pi^2 N}{V} \right)^{1/3}$$

$$\text{total } E = \int_0^\mu \epsilon p(\epsilon) d\epsilon \quad p(\epsilon) d\epsilon = p(k(\epsilon)) \left| \frac{dk}{d\epsilon} \right| d\epsilon$$

$$p(k) = \frac{V}{2\pi^2} k^2 \quad k = \frac{\epsilon}{c\hbar} \quad \frac{dk}{d\epsilon} = \frac{1}{c\hbar}$$

$$p(\epsilon) = \frac{V}{2\pi^2} \frac{\epsilon^2}{c^2 \hbar^2} \cdot \frac{1}{c\hbar}$$

$$E = 2 \cdot \frac{V}{2\pi^2} \frac{1}{c^3 \hbar^3} \int_0^\mu \epsilon^3 d\epsilon = \frac{V}{4\pi^2 c^3 \hbar^3} \mu^4 = \frac{V c \hbar}{4\pi^2} \frac{(3\pi^2)^{4/3} N^{4/3}}{V^{4/3}}$$

$$E = \frac{3N \hbar \pi c}{4} \left(\frac{3N}{\pi V} \right)^{1/3}$$

$$d. E = \frac{3N\hbar c}{4} \left(\frac{3N}{\pi \cdot \frac{4}{3}\pi} \right)^{1/3} \frac{1}{R}$$

$$U_g = -\frac{3GM^2}{5R}$$

If $|U_g| > E$, then collapse occurs $M = 2Nm_p$ $N = \frac{M}{2m_p}$

$$\cdot \frac{3\pi\hbar c}{4} \left(\frac{9}{4\pi^2} \right)^{1/3} \frac{M^{4/3}}{(2m_p)^{4/3}} < \frac{3GM^2}{5}$$

$$M^{2/3} > \frac{\hbar c}{G} \frac{5}{16} \frac{(9\pi)^{1/3}}{m_p^{4/3}}$$

$$M > \left(\frac{\hbar c}{G} \right)^{3/2} \left(\frac{5}{16} \right)^{3/2} \frac{3\pi^{1/2}}{m_p^2} = 3.4 \cdot 10^{30} \text{ kg}$$

$$M > 1.72 M_\odot$$