

Reaching the Nonlinear Regime of Raman Amplification of Ultrashort Laser Pulses

W. Cheng, Y. Avitzour, Y. Ping, and S. Suckewer*

Department of MAE, School of Engineering and Applied Science, Princeton University, Princeton, New Jersey 08544, USA

N. J. Fisch

Princeton Plasma Physics Laboratory & Astrophysics Department, Princeton University, Princeton, New Jersey 08544, USA

M. S. Hur and J. S. Wurtele

Lawrence Berkeley National Laboratory & Physics Department, University of California, Berkeley, California 94720, USA

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The intensity of a subpicosecond laser pulse was amplified by a factor of up to 1000 using the Raman backscatter interaction in a 2 mm long gas jet plasma. The process of Raman amplification reached the nonlinear regime, with the intensity of the amplified pulse exceeding that of the pump pulse by more than an order of magnitude. Features unique to the nonlinear regime such as gain saturation, bandwidth broadening, and pulse shortening were observed. Simulation and theory are in qualitative agreement with the measurements.

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The invention of chirped pulse amplification (CPA) [1,2] led to a tremendous increase of ultrashort laser pulse intensities to above 10^{20} W/cm² [3,4]. However, such an ultrahigh intensity laser system was achieved using very large (on the order of ~ 1 m²) and expensive compressor gratings [4]. A further increase of the pulse intensity using the CPA technique would require even larger gratings in order not to exceed the material damage threshold. Such a system would be very difficult to implement in university-scale laboratories.

In order to overcome the CPA material limit at ultrahigh intensities, different backscattering coupling techniques were proposed in plasma, including Compton scattering [5], resonant Raman backscattering [6], and Raman backscattering at an ionization front [7]. The experiments reported here utilize the resonant Raman mechanism [6], where a short seed laser pulse is amplified by a counterpropagating long pump pulse, with their frequencies satisfying the resonance relation, $\omega_{\text{pump}} = \omega_{\text{seed}} + \omega_{pe}$ where ω_{pump} , ω_{seed} , and ω_{pe} are frequencies of pump, seed, and plasma, respectively; $\omega_{pe} = \sqrt{4\pi e^2 n_e / m_e}$, n_e is the plasma electron density, and m_e and e are mass and charge of an electron. The energy transfer from pump to seed is in proportion to their frequencies, so for $\omega_{\text{pump}} \geq 10\omega_{pe}$, the efficiency can be as high as 90%. What makes the resonant Raman backscatter regime attractive is that it is a simple resonant interaction, with the seed amplification strong enough to outrun other deleterious competing instabilities (such as modulational instability that can lead to the filamentation of the laser beam) [6] or to avoid superluminous precursor solutions [8], and with realizable highly compressed ultrashort pulse solutions [9].

The Raman backscattering (RBS) amplification can be divided into linear and nonlinear regimes. In the linear

regime the pump depletion is negligible and the gain is independent of the seed intensity. The seed pulse is amplified and increased in duration due to the narrow bandwidth of the linear amplification. The nonlinear regime, the so-called π -pulse regime, is characterized by pump depletion and the simultaneous temporal compression of the amplified pulse. In this regime, the Raman amplification and compression of ultrashort pulses in a plasma allow intensities to reach 10^{20} – 10^{21} W/cm² in a compact university-scale device, and unprecedented high intensity on the order of 10^{25} W/cm² in a larger system [9]. Such intensities open new frontiers in physics and should also lead to practical applications such as fast ignition for inertial fusion [10], x-ray lasers [11], or laser wake field accelerators [12].

Previous experiments demonstrated low amplifications in a short microcapillary [13] and gas jet plasmas [14]. Recently, energy amplification of close to 100 was demonstrated in a plasma created in a 2 mm long gas jet with effective (i.e., with proper density) plasma length ~ 1 – 1.5 mm. In that experiment [15], a 550 fs broadband seed (from an optical parametric generation and amplification system) was pumped by a counterpropagating 10 ps Ti:sapphire laser pulse at a wavelength of 800 nm. Only 10% of the seed spectrum (0.8 μ J) that satisfied the RBS resonance condition was amplified. While some temporal compression of seed pulse was observed at maximum amplification, the analysis of data did not provide clear evidence that the nonlinear regime was reached. In similar experiments, energy amplification of ~ 100 [16] were demonstrated just recently. Energy amplification of ~ 10 was also reported [17], with the result ascribed in part to Compton scattering [5], but the far greater part of the amplification, if not all of it, appears to be in the resonant regime associated with the other early experiments.

What this Letter reports is the first demonstration of reaching the important nonlinear regime of RBS amplification, where we can expect the onset of pump depletion and significant pulse shortening. This regime is reached here with pulse intensity amplification ~ 1000 in a 2 mm long gas jet plasma. The intensity of the amplified exceeded that of the pump by more than 1 order of magnitude. Three major changes in the present experiment led to the new results: (1) Development of a narrow-band seed pulse with a Gaussian profile in space and time. (2) Improvement of the uniformity of the plasma with an effective length up to 2 mm. (3) Installation of diagnostics for precise temporal and spatial alignment of the laser pulses in plasma.

The experimental setup is shown schematically in Fig. 1. The electron density n_e was $1.1 \times 10^{19} \text{ cm}^{-3}$ ($\omega_{pe} = 2 \times 10^{14} \text{ s}^{-1}$), which matched the difference between frequencies of the pump ($\omega_{\text{pump}} = 2.4 \times 10^{15} \text{ s}^{-1}$) and the seed ($\omega_{\text{seed}} = 2.2 \times 10^{15} \text{ s}^{-1}$). The plasma was created in a propane (C_3H_8) gas jet through optical breakdown induced by a low-power Nd:YAG laser pulse (the ‘‘prepulse’’: wavelength $1.064 \mu\text{m}$, pulse duration 6 ns, and energy $\sim 500 \text{ mJ}$), which was focused to a $\sim 50 \mu\text{m}$ spot just above and parallel to a gas jet nozzle. The Rayleigh length of the prepulse was $\sim 2 \text{ mm}$. Propane was used in order to have a large number of electrons from a single molecule so that a sufficiently high electron density could be achieved. A stretched Ti:sapphire laser beam (central wavelength 800 nm with a bandwidth of 11 nm and a pulse duration of 170 ps) was split into two beams. The first beam, which served as a pump pulse, was compressed by compressor C1 to 10 ps and maintained the original bandwidth. The second beam was used as a pump pulse for a Raman shifter [18] (a 7 cm long barium nitrate crystal). The first Stokes beam, with the wavelength center at 873 nm and the same bandwidth as that of the pump, was generated in the crystal and compressed by the second compressor C2 to about

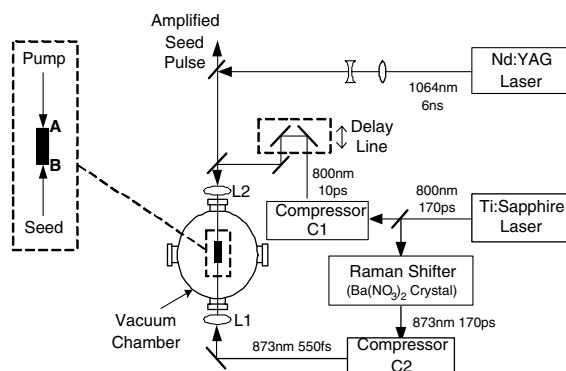


FIG. 1. Schematic of the experimental setup. Plasma is formed by the Nd:YAG laser. Part of the output from the Ti:sapphire laser is compressed by C1 to 10 ps and serves as the pump. The other part passes through a Raman shifter, is compressed by C2 to 550 fs, and serves as the seed pulse. The two pulses are focused to the plasma in the vacuum chamber.

550 fs. The pump entered the plasma at point A and the seed at point B. The output seed passed through a long wavelength ($>850 \text{ nm}$) pass filter so that any reflected pump beam was filtered out. A small part of the amplified seed pulse was directed to a spectrometer for spectral analysis. The rest of the beam was split by a 50% beam splitter. One part was directed to a power meter and the other part was directed to an autocorrelator.

Figure 2(a) shows the output energy versus the input seed energy when the pump energy was maintained at 40 mJ. The thermal Raman was the output signal when the seed pulse was blocked. Its energy was measured to be $\sim 80 \mu\text{J}$. While the thermal Raman output was fairly stable, we observed fluctuations of the energy of the amplified pulse, possibly because the amplification process has a narrower bandwidth than the noise, and thus is more sensitive to the fluctuations of the plasma density. Only the largest five outputs (which represented the optimal conditions for amplification) from an average of 15 shots were shown in each data point. When the seed pulse was let into the plasma (the solid line), the output grew very rapidly if the seed pulse energy was less than $1.5 \mu\text{J}$. As the input seed energy was increased beyond $1.5 \mu\text{J}$, the growth rate decreased and the gain was saturated. For the minimum initial seed energy of $0.05 \mu\text{J}$, the output was about $350 \mu\text{J}$. Subtracting the thermal Raman energy (about $80 \mu\text{J}$) from the total output, we had an energy amplification factor of ~ 5000 . As the gain was saturated, the amplification dropped to ~ 60 when the input seed was at its maximum ($7.5 \mu\text{J}$). The output energy versus the pump energy is shown in Fig. 2(b) (solid line) with the energy of the input seed pulse maintained at $\sim 7.5 \mu\text{J}$. The dotted line shows the energy of the thermal Raman.

The pulse compression in the nonlinear regime is accompanied by bandwidth broadening. Typical spectra of the pump, the input seed and the amplified seed, are shown

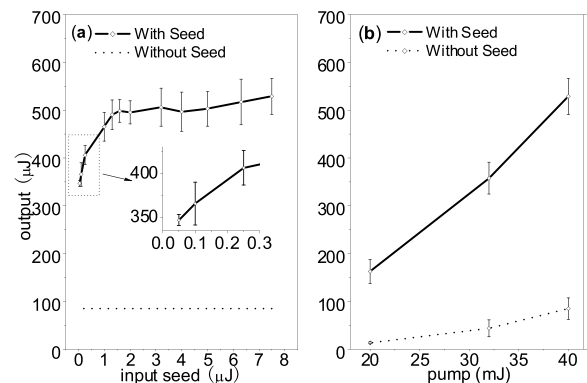


FIG. 2. Output energy versus input seed energy and pump energy. (a) Output energy versus input seed energy when the pump energy was maintained at 40 mJ. Inset shows the output when the input was low and has the same scale as the main figure. (b) Output energy versus pump energy when the seed energy was maintained at $7.5 \mu\text{J}$. In both cases, the dotted line denotes the thermal Raman.

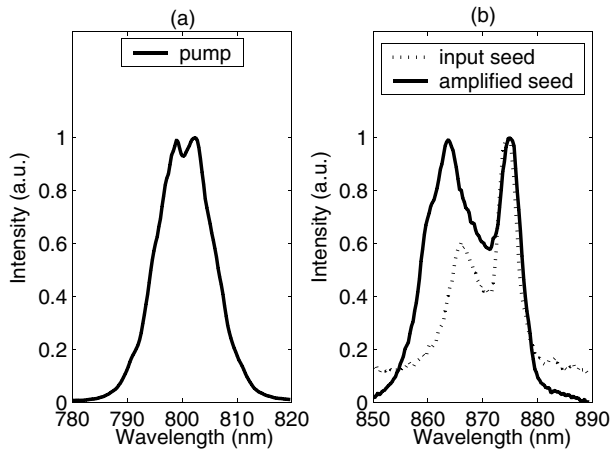


FIG. 3. Normalized (to 1) spectra of (a) pump pulse and (b) input seed pulse and amplified seed pulse for amplification of 40.

in Fig. 3. The pump (for pumping of both the barium nitrate crystal and the plasma) and the input seed had a bandwidth of about 12 nm. The dip in the center of the spectrum of the pump beam may have contributed to the big dip in the spectrum of the input seed pulse because of the nonlinear process of the Raman shifter [18]. The spectrum of an amplified seed pulse was measured when the energy amplification was ~ 40 for an initial seed energy of $7.5 \mu\text{J}$. The bandwidth of the amplified seed was ~ 19 nm.

In order to observe the spatial profile evolution of the seed beam, it was imaged onto a CCD camera at the exit of the plasma (point A in Fig. 1). In the experiment, both the pump and the seed were focused to point B and had a beam size of $\sim 30 \mu\text{m}$ FWHM. The seed beam profile at point A in the absence of the plasma is shown in Fig. 4(a). The size of the beam was measured to be $55 \mu\text{m}$ FWHM. When the plasma was present and the seed pulse was amplified by the counterpropagating pump pulse, it preserved the Gaussian profile and showed significant beam narrowing. The beam size shown in Fig. 4(b) has a FWHM of $13 \mu\text{m}$. (A neutral density filter was used to reduce the amplified seed energy to avoid saturation of the CCD camera). The beam narrowing effect may be explained by the nonuniform amplification across the beam profile: the central part of the seed was amplified more due to the stronger pump in the center. The self-focusing of the amplified seed pulse in plasma may also have contributed to the beam narrowing, although we did not observe significant seed beam narrowing after it passed through the plasma without the pump pulse.

With the high amplification of the seed pulse, we were able to perform a pulse duration measurement of the amplified seed using a standard autocorrelator (Positive Light Model SSA) at various amplification stages. With the pump and seed energy maintained at 40 mJ and $7.5 \mu\text{J}$ respectively, the energy amplification spanned a range from 10 to 60 due to the fluctuations of the plasma parameters. The pulse was observed to lengthen to a maximum of ~ 800 fs when the amplification was close to 10.

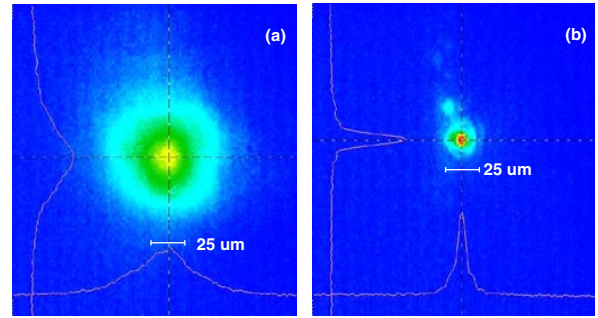


FIG. 4 (color online). Output seed pulse profiles (a) without amplification (no plasma) and (b) after amplification (image obtained through a 2.5% transmission filter).

The pulse duration began to decrease when the amplification approached 30. While the plasma fluctuations caused large standard deviations of the output pulse duration, the trend that the pulse became shorter as the amplification became higher can be seen clearly in Fig. 5. The shortest pulse we have observed so far had a FWHM of ~ 150 fs when the energy amplification was ~ 60 with a $7.5 \mu\text{J}$ input seed. At this point, the power of the seed pulse increased by more than 200 times with the intensity amplification reaching ~ 1000 times. The intensity of the amplified seed pulse at point A in Fig. 1 was $1.7 \times 10^{15} \text{ W/cm}^2$. The pump pulse diameter was about $70 \mu\text{m}$ FWHM at point A and had an intensity of $\sim 1 \times 10^{14} \text{ W/cm}^2$. The intensity of the amplified seed exceeded the pump intensity at this point by more than 1 order of magnitude.

The physical parameters of the experiment are such that kinetic plasma effects influence the seed pulse amplification process. We have investigated these effects using the averaged particle-in-cell simulation code APIC [19]. The code models the electromagnetic waves in the slowly varying amplitude and phase approximation, and uses standard plasma simulation techniques to model the Langmuir wave and plasma dynamics. The seed and pump pulses are assumed to be initially Gaussian with FWHM given by the experimental values of 550 fs and

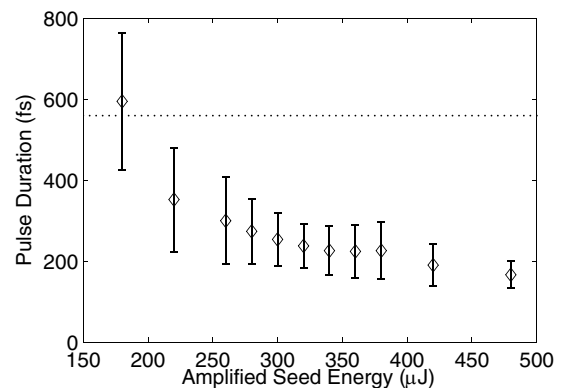


FIG. 5. The pulse duration versus energy of the amplified seed pulse. Dotted line denotes the initial pulse duration.

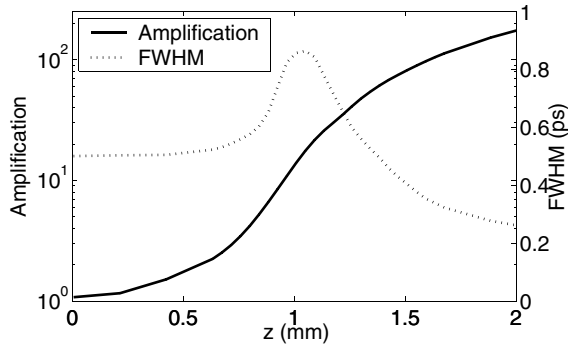


FIG. 6. The 1D, time dependent computer simulation of energy amplification (solid line) and amplified pulse duration at its FWHM (dotted line) versus plasma length for 10 ps pump.

10 ps, respectively. The initial seed intensity was 1.6×10^{12} W/cm², pump intensity was 1×10^{14} W/cm², and both were close to experimental parameters. The plasma density was chosen so that the plasma wave (including a thermal shift) was resonant with the beat frequency of the lasers. The electrons were assumed to be initially homogeneous with a Maxwellian velocity distribution with $T_e = 50$ eV, and ions were assumed to be motionless over the short time scale of the simulation. Typical simulation results are shown in Fig. 6. The seed energy amplification and the FWHM of the seed pulse duration are plotted as a function of the interaction length. The simulation clearly shows pulse stretching of the seed during the initial amplification (until roughly 1 mm, corresponding to an energy amplification of ~ 20), followed by pulse shortening to ~ 240 fs at the end of 2 mm plasma. While the pulse duration evolution agreed quite well with the experimental measurements, the energy gain factor did not agree so well. The simulation results shown here are one-dimensional and therefore do not include the effects of diffraction and transverse variation of the laser pulses. Kinetic effects are significant when the initial value chosen for the temperature is 100 eV. At temperatures of 50 eV and below, with Gaussian pump pulses, the qualitative features of the experiment are consistent with the experimental results.

In summary, we demonstrated energy amplification of up to ~ 5000 for a seed with a low input energy (~ 0.05 μ J), and the amplification was reduced, by saturation effects, to ~ 60 for seeds with significantly higher input energy (7.5 μ J). The beam transverse profile of the amplified seed was narrowed by the amplification process. The measurement of the pulse duration and its spatial

profile allowed us to determine a maximum amplified pulse intensity to be $\sim 1.7 \times 10^{15}$ W/cm², which was higher than pump intensity by more than 1 order of magnitude. This is a strong indication of reaching the nonlinear regime. The pulse was lengthened from approximately 0.5 ps to up to 0.8 ps when the energy amplification was ~ 10 , and was compressed to below ~ 200 fs when the energy amplification reached ~ 60 for maximum input seed energy. This is additional evidence that the nonlinear regime was reached. We noticed, however, only a relatively small fraction of the total energy of the pump pulse was transferred to the seed pulse. Increasing the effective plasma length and pump pulse duration, together with the guiding of the pump and seed pulses through a plasma channel [20] should dramatically increase the energy transfer between the pump and the seed.

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*Electronic address: suckewer@princeton.edu

- [1] D. Strickland and G. Mourou, *Opt. Commun.* **56**, 219 (1985).
- [2] M. Pessot *et al.*, *Opt. Lett.* **14**, 797 (1989).
- [3] M. D. Perry and G. Mourou, *Science* **264**, 917 (1994).
- [4] M. D. Perry *et al.*, *Opt. Lett.* **24**, 160 (1999).
- [5] G. Shvets *et al.*, *Phys. Rev. Lett.* **81**, 4879 (1998).
- [6] V. M. Malkin, G. Shvets, and N. J. Fisch, *Phys. Rev. Lett.* **82**, 4448 (1999).
- [7] V. M. Malkin and N. J. Fisch, *Phys. Plasmas* **8**, 4698 (2001).
- [8] Y. A. Tsidulko *et al.*, *Phys. Rev. Lett.* **88**, 235004 (2002).
- [9] N. J. Fisch and V. M. Malkin, *Phys. Plasmas* **10**, 2056 (2003).
- [10] M. Tabak *et al.*, *Phys. Plasmas* **1**, 1626 (1994).
- [11] J. Dunn *et al.*, *Phys. Rev. Lett.* **84**, 4834 (2000).
- [12] P. Sprangle *et al.*, *Phys. Rev. Lett.* **69**, 2200 (1992).
- [13] Y. Ping *et al.*, *Phys. Rev. E* **66**, 046401 (2002).
- [14] Y. Ping, I. Geltner, and S. Suckewer, *Phys. Rev. E* **67**, 016401 (2003).
- [15] Y. Ping *et al.*, *Phys. Rev. Lett.* **92**, 175007 (2004).
- [16] A. Balakin *et al.*, *JETP Lett.* **80**, 15 (2004).
- [17] M. Dreher *et al.*, *Phys. Rev. Lett.* **93**, 095001 (2004).
- [18] N. Zhavoronkov *et al.*, *Opt. Lett.* **26**, 47 (2001).
- [19] M. S. Hur *et al.*, *Phys. Plasmas* **11**, 5204 (2004).
- [20] P. Mardahl *et al.*, *Phys. Lett. A* **296**, 109 (2002).