Adiabatic nonlinear waves with trapped particles. II. Wave dispersion

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A general nonlinear dispersion relation is derived in a nondifferential form for an adiabatic sinusoidal Langmuir wave in collisionless plasma, allowing for an arbitrary distribution of trapped electrons. The linear dielectric function is generalized, and the nonlinear kinetic frequency shift $\omega_{\rm NL}$ is found analytically as a function of the wave amplitude *a*. Smooth distributions yield $\omega_{\rm NL} \propto \sqrt{a}$, as usual. However, beam-like distributions of trapped electrons result in different power laws, or even a logarithmic nonlinearity, which are derived as asymptotic limits of the same dispersion relation. Such beams are formed whenever the phase velocity changes, because the trapped distribution is in autoresonance and thus evolves differently from the passing distribution. Hence, even adiabatic $\omega_{\rm NL}(a)$ is generally nonlocal. © 2012 American Institute of Physics. [doi:10.1063/1.3662115]

I. INTRODUCTION

Within the geometrical-optics approximation, adiabatic waves in collisionless plasma are described conveniently within the average-Lagrangian formalism originally proposed by Whitham.^{1,2} In Ref. 3, further called Paper I, this formalism was restated to also accommodate the effects of particles trapped in wave troughs. Specifically, the wave Lagrangian density Ω was derived explicitly under the assumption that the number of trapped particles within each wavelength is a fixed independent parameter of the problem. Hence, the wave nonlinear dispersion relation (NDR) can be expressed in a nondifferential form which is more easily tractable than solutions obtained straightforwardly from integrating the Vlasov-Maxwell system (see, e.g., Refs. 4–9) or those inferred from the "multiwater bag" formulation.¹⁰ Namely, the general NDR for an adiabatic wave reads as (Paper I)

$$\partial_a \mathfrak{L} = 0, \tag{1}$$

with *a* being the wave amplitude; thus, analyzing the function $\partial_a \Omega$ would be sufficient to infer the wave dispersion.

Here, we perform a systematic study of kinetic nonlinearities that Eq. (1) yields for one-dimensional sinusoidal electrostatic waves with trapped particles, extending our brief letter¹¹ and providing important technical details missing in that work. Specifically, a general NDR is derived in a nondifferential form for an adiabatic sinusoidal Langmuir wave in collisionless plasma, allowing for an arbitrary distribution of trapped electrons [Eq. (11)]. The linear dielectric function is generalized, and the nonlinear kinetic frequency shift $\omega_{\rm NL}$ is found analytically as a function of the wave amplitude a. Smooth distributions yield $\omega_{\rm NL} \propto \sqrt{a}$, as usual. However, beam-like distributions of trapped electrons result in different power laws, or even a logarithmic nonlinearity, which are derived as asymptotic limits of the same dispersion relation. Such beams are formed whenever the phase velocity changes, because the trapped distribution is in autoresonance and thus evolves differently from the passing distribution; hence, even adiabatic $\omega_{NL}(a)$ is generally nonlocal. These results are also used in Paper III,¹² where evolution of waves with trapped particles is studied.

The work is organized as follows. In Sec. II, the general NDR is derived, connecting the wave frequency ω with the wave number *k* for a given *a* and particle distribution; also the properties of the NDR kernel are explored. In Sec. III, the general NDR is given an asymptotic representation for distributions smooth near the resonance; in particular, the linear dielectric function is generalized. In Sec. IV, the effect of particle beams on NDR is studied. In Sec. V, implications of these results are discussed; specifically, $\omega_{NL}(a)$ is predicted to be a nonlocal function, and the friction drag on trapped particles is predicted to affect the wave frequency sweeping in collisional plasmas. In Sec. VI, we summarize our main results. Some auxiliary formulas are also presented in the Appendix.

II. GENERAL DISPERSION RELATION

A. Wave Lagrangian

From Paper I, it follows that the Lagrangian density for a one-dimensional adiabatic nonlinear electrostatic electron mode in collisionless plasma reads as

$$\mathfrak{L} = \bar{\mathfrak{L}}_{\rm em} + \frac{\langle \tilde{E}^2 \rangle}{8\pi} - n^{(p)} \langle \mathcal{H}^{(p)} \rangle_{f_p} - n^{(t)} \langle \mathcal{H}^{(t)} \rangle_{f_t}, \qquad (2)$$

assuming that the ion contribution is negligible. Here, $\bar{\mathfrak{L}}_{em}$ describes quasistatic fields (if any), the second term is the timeaveraged energy density of the wave electric field \tilde{E} , $n^{(p)}$ and $n^{(t)}$ are the densities of passing and trapped electrons, respectively, and $\mathcal{H}^{(p)}$ and $\mathcal{H}^{(t)}$ are the electron oscillation-center (OC) energies averaged over the corresponding distributions f_p and f_t of the particle canonical momenta. Specifically,

$$\mathcal{H}^{(p)} = \mathcal{E} + Pu - mu^2/2, \tag{3}$$

$$\mathcal{H}^{(t)} = \mathcal{E} - mu^2/2,\tag{4}$$

where *m* is the electron mass and $\mathcal{E}(J, a)$ is the electron energy in the frame \hat{K} moving with respect to the laboratory

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frame *K* at the wave phase velocity $u = \omega/k$. Also, *J* is the trapped-particle canonical momentum, or the canonical action in \hat{K} , and

$$P = mu \pm kJ \tag{5}$$

is the passing-particle canonical momentum, where the plus sign is chosen for particles traveling with velocities larger than u and the minus sign is chosen otherwise.

Combining Eqs. (2)–(4), one gets

$$\mathfrak{L} = \bar{\mathfrak{L}}_{\rm em} + \frac{nmu^2}{2} - n_p^{(p)} \langle P \rangle_{f_p} u + \frac{\langle \vec{E}^2 \rangle}{8\pi} - n \langle \mathcal{E} \rangle_f,$$

where we introduced the electron total density $n = n^{(p)} + n^{(t)}$, and the energy \mathcal{E} is now being averaged over all particles. By definition (Paper I), the particle densities ω and k are the arguments of \mathfrak{L} that are independent of a. [It is only through *solutions* of Eq. (1) that they will become interconnected.] Hence, Eq. (1) yields the general NDR in the form

$$\partial_a \langle \tilde{E}^2 \rangle / (8\pi) - n \,\partial_a \langle \mathcal{E} \rangle_f = 0, \tag{6}$$

where ∂_a is taken at constant (ω , k) and also at fixed canonical momenta of individual particles. Specifically, fixed at differentiation can be J, not only for trapped particles but also for passing ones, due to Eq. (5).

B. Sinusoidal-wave approximation

Equation (6) describes the dispersion of onedimensional electron Langmuir waves completely, accounting simultaneously for linear and nonlinear dispersive effects, both fluid and kinetic. In principle, each of those effects can then be assessed from the general NDR by substituting the appropriate $\mathcal{E}(J, a)$.¹³ Yet, we will address only one specific regime here, as an illustration of how our Lagrangian formalism can be useful for finding NDRs explicitly; namely, we will assume that the wave can be considered monochromatic. The conditions under which this assumption is justified as a precise asymptotic limit at small amplitudes are discussed in detail in Refs. 14 and 15 (see also Refs. 16 and 17) and will not be repeated here. Instead, our purpose will be to show that, once the sinusoidal-wave approximation has been adopted (like in traditional calculations), our Lagrangian formalism predicts new effects and also yields simultaneously various types of kinetic nonlinearities that were previously known from different contexts.

In a sinusoidal wave, the oscillations of both passing and trapped electrons in \hat{K} satisfy

$$p^2/(2m) + e\tilde{\varphi}_0 \cos(kx) = \mathcal{E},\tag{7}$$

where *x* and *p* are the particle coordinate and momentum and *e* is the electron charge. (For clarity, we will assume $e\tilde{\phi}_0 > 0$ and k > 0. Also, the large-scale electrostatic potential $\bar{\phi}$ is omitted here because it does not affect the wave local properties, albeit it may affect the dynamics of the wave (Paper III) and of the particle distribution.¹⁸) The canonical action *J* is hence introduced as follows. For passing particles, we will



FIG. 1. (Color online) Schematic of single-particle trajectories in phase space (x, p), illustrating the definition of $2\pi J$ (shaded area): (a) for a passing particle and (b) for a trapped particle. For *J* to be continuous at the separatrix (dashed line), with the passing-particle action defined as $2\pi J = \oint p dx$, for trapped particles, one must use the definition $2\pi J = (1/2) \oint p dx$. This area is encircled by a particle within half of the bounce period; thus, the corresponding canonical frequency Ω equals twice the bounce frequency.

use the definition $J = (2\pi)^{-1} \oint p dx$, and thus, J = |p|/k for large p. For trapped particles, however, we will use $J = (4\pi)^{-1} \oint p dx$, so that J is continuous across the separatrix (Fig. 1). From Eq. (7), one can write then

$$J = \sqrt{aJj}.$$
 (8)

Here, we introduced the wave amplitude as $a = e\tilde{\varphi}_0 k^2/(m\omega^2) > 0$; also $\hat{J} = m\omega/k^2$, and the dimensionless action *j* is a continuous function of the normalized energy $r = (\mathcal{E} + e\tilde{\varphi}_0)/(2e\tilde{\varphi}_0)$,

$$j = j(r). \tag{9}$$

Specifically, j = 0 for a particle resting at the bottom of the wave trough (r = 0), with the corresponding value at the separatrix (r = 1) being $j_* = 4/\pi$ (Appendix).

Hence, one can express \mathcal{E} as

$$\mathcal{E}(J,a) = [2ar(j) - a]\hat{J}\omega, \qquad (10)$$

with j(J, a) taken from Eq. (8). Then Eq. (6) reads as

$$\omega^2 = \omega_p^2 \frac{2}{a} \int_0^\infty GF(J) \, dJ,\tag{11}$$

where $\omega_p^2 = 4\pi ne^2/m$ is the electron plasma frequency squared, F(J) is the action distribution normalized such that $\int_0^\infty F(J) dJ = 1$, and

$$G = \frac{\partial_a \mathcal{E}(J, a)}{mu^2}.$$
 (12)

Equation (11) is a master equation, which accounts for all dispersive effects simultaneously, both linear and nonlinear (to the extent that the sinusoidal-wave approximation applies¹⁹). In particular, it shows that the contribution to ω^2 of particles with given *J* is determined solely by the kernel *G*, thus acting as the weight function. We study the properties of this function below.

C. Weight function G

From Eq. (12), one gets

$$G = [\partial_a (2ar - a)]_J = 2r - 1 + 2a (\partial_a r)_J, \qquad (13)$$

where the partial derivative is taken at fixed J (as denoted by the subindex). The latter term is found from

$$(\partial_a r)_J = -\frac{(\partial_a J)_r}{(\partial_r J)_a} = -\frac{j}{2aj'},\tag{14}$$

where we substituted Eq. (8). Thus, G = 2r - 1 - j/j', where the right-hand side here is a function of *r* only or a function of *j* only [due to Eq. (9)]. We will introduce these two representations separately, with

$$g(r) = 2r - 1 - j(r)/j'(r)$$
(15)

(Appendix), serving as an auxiliary function for

$$G(j) = g(r(j)). \tag{16}$$

Hence, the properties of G(j) are understood as follows.

First, let us consider asymptotics of G(j), starting with $j \ll 1$. As seen from Eq. (15), G(0) = -1, which also flows from the fact that $\mathcal{E} \approx -e\tilde{\varphi}_0$ for particles close to the bottom of the wave trough. The next-order term can be inferred from $\mathcal{E} \approx -e\tilde{\varphi}_0 + J\Omega$, which is due to oscillations of deeply trapped particles being approximately linear. In particular, one can take $\Omega \approx 2\omega_E$ (Fig. 1), where $\omega_E = (ek^2\tilde{\varphi}_0/m)^{1/2} = \omega\sqrt{a}$ is the linear bounce frequency.²⁰ Hence, Eq. (12) yields $G \approx -1 + j$.

The asymptotic behavior of G(j) at $j \gg 1$ is understood by taking $\mathcal{E} \approx p^2/(2m) + \Phi$, where $\Phi = e^2 m \tilde{\varphi}_0^2/(4p^2)$ is the ponderomotive energy in the dipole approximation (Paper I) for a zero-frequency wave [Eq. (7)]. Remember that $|p| \approx Jk$ in this case, so $(\partial_a \mathcal{E})_J = 2\Phi/a$. Thus,

$$G(j \gg 1) \approx 1/(2j^2),\tag{17}$$

and higher-order corrections can be found similarly.

While negative at small *j*, G(j) is positive at large *j*. (This means that deeply trapped particles reduce ω^2 , whereas those near the separatrix and untrapped ones increase ω^2 ; cf., Refs. 21 and 22.) Now let us consider its behavior near the separatrix, which is where the canonical frequency Ω approaches zero. Since \mathcal{E} equals the OC energy in \hat{K} (Paper I), one has

$$\Omega = \partial_J \mathcal{E} = \omega \,\partial_j (2ar - a) = 2a\omega/j'. \tag{18}$$

Therefore, j'(r=1) is infinite, and Eq. (15) yields that *G* be continuous (albeit nonanalytic), with $G(j_*) = 1$. Hence, we summarize our results as follows:

$$G(j) = \begin{cases} -1+j+\dots, & j \ll 1, \\ 1, & j=1, \\ \frac{1}{2j^2} + \frac{5}{16j^6} + \dots, & j \gg 1. \end{cases}$$
(19)

These are matched (Fig. 2) by direct calculations in which an explicit expression for j(r) is substituted into Eq. (15). Also, manipulating that and related formulas given in the Appendix shows that the NDR arising from the waterbag formulation (Ref. 10, with zero driving field) can, in principle,²³ be reduced to Eq. (11).



FIG. 2. Weight function G(j). Dashed lines are approximate solutions given by Eq. (19), the asymptote, and $j = j_*$.

D. Small-amplitude cold-plasma limit

It is instructive to consider, as an example, the coldplasma case in the limit of small amplitudes. Then, all particles are passing, and J = |p|/k, so (Paper I)

$$v^{\pm} = u \pm kJ/m, \tag{20}$$

where the sign index denotes sgn (v - u). Yet v^+ are not present then, and $v^- \ll u$, in which case Eq. (20) gives $J \approx \hat{J}$. [In other words, one may assume $F(J) \approx \delta(J - \hat{J})$.] From Eq. (17), one thereby obtains G = a/2, and thus, Eq. (11) predicts $\omega^2 = \omega_n^2$, as expected.

Below we will calculate ω^2 also for other representative cases, by formally constructing various asymptotic expansions of the integral in Eq. (11). Remarkably, Eq. (19) will be sufficient for that, i.e., knowing the exact form of G(j) is not needed, except for evaluating insignificant numerical coefficients (Appendix).

III. SMOOTH DISTRIBUTIONS

First, let us consider the case when the distribution function F(J) is smooth, e.g., thermal. Of course, the wave *dynamics* could not be described then within the present approach, on the score of particle unavoidable trapping and detrapping that would be associated with variations of *a*, rendering the wave nonadiabatic. Yet, assuming that the wave is quasihomogeneous and quasistationary (which are our necessary requirements in any case), its local dispersion must be approximately the same as that of a truly homogeneous and stationary wave.²⁴ Since the latter has $n^{(t)}$ fixed and is clearly adiabatic, our formalism is hence applicable for deriving the local NDR and can be built on as follows.

Suppose that F(J) remains finite at small J or, at least, diverges less rapidly than J^{-1} . Then, one can take the integral in Eq. (11) by parts and obtain

$$1 - \frac{2\omega_p^2}{a\omega^2} \int_0^\infty \Upsilon(J, a) F'(J) \, dJ = 0, \tag{21}$$

where $\Upsilon(J, a) = \Psi(j)\hat{J}\sqrt{a}$, with the dimensionless function $\Psi(j)$ [Fig. 3(a)] introduced as



FIG. 3. (a) Function $\Psi(j)$, Eq. (22). (b) Function Q(j), Eq. (26). The dashed lines show $j = j_*$ and also the asymptote of Q(j). For details, see the Appendix.

$$\Psi(j) = -\int_0^j G(j) \, dj. \tag{22}$$

At $a \ll 1$, Υ changes rapidly with *J* compared to *F*(*J*). Then, without using the explicit form of $\Upsilon(J, a)$ but rather drawing on the leading terms in

$$\Upsilon(J,a) = \begin{cases} J - \frac{J^2}{2\hat{J}a^{1/2}} + \dots, & J \ll \hat{J}a^{1/2}, \\ \frac{a\hat{J}^2}{2J} + \frac{a^3\hat{J}^6}{16J^5} + \dots, & J \gg \hat{J}a^{1/2} \end{cases}$$
(23)

[obtained from Eq. (19)], one can put Eq. (21) in an asymptotic form. Specifically, this is done as follows.

A. Asymptotic representation

To approximate the integral

$$U \equiv \frac{2}{a} \int_0^\infty \Upsilon(J, a) F'(J) \, dJ, \tag{24}$$

let us replace $\Upsilon F'$ with

$$\Upsilon F' \approx (\Upsilon F')_1 + (\Upsilon F')_2 - (\Upsilon F')_3. \tag{25}$$

Here, the individual terms on the right-hand side correspond to asymptotics of $\Upsilon F'$ at small, large, and intermediate *J*, correspondingly, so that $\Upsilon F'$ is approximated accurately at all *J*. Specifically, we take

$$(\Upsilon F')_i = \begin{cases} \Upsilon (J) [F'(0) + F''(0)J], & i = 1, \\ \left[a \hat{J}^2 / (2J) \right] F(J), & i = 2, \\ \left[a \hat{J}^2 / (2J) \right] \left[F'(0) + F''(0)J \right], & i = 3. \end{cases}$$

[Due to $\Upsilon - a\hat{J}^2/(2J) = \mathcal{O}(J^{-5})$ at large J, higher-order corrections are insignificant compared to those discussed below.] Hence, $U = U_1 F''(0) + U_2$, where

$$U_{1} = \frac{2}{a} \int_{0}^{\infty} \left\{ \Upsilon(J, a) J - \frac{a\hat{J}^{2}}{2} \right\} dJ,$$

$$U_{2} = \frac{2}{a} \int_{0}^{\infty} \left\{ \Upsilon(J, a) F'(0) - \frac{a\hat{J}^{2}}{2J} \left[F'(J) - F'(0) \right] \right\} dJ$$

It is convenient to rewrite U_1 in terms of the function

$$Q(j) = 1 - 2j\Psi(j) \tag{26}$$

[Fig. 3(b)], which we will also use below. Namely,

$$U_1 = -\varkappa \hat{J}^3 \sqrt{a}, \quad \varkappa = \int_0^\infty Q(z) \, dz. \tag{27}$$

The value of the numerical constant \varkappa can be readily estimated from Fig. 3(b) as $\varkappa \sim 0.5$; for a precise calculation, see the Appendix.

To simplify the expression for U_2 , notice that

$$\frac{\partial U_2}{\partial a} = 2F'(0) \int_0^\infty \frac{\partial}{\partial a} \left[\frac{\Upsilon(J,a)}{a} \right] dJ, \qquad (28)$$

which can be checked directly. On the other hand,

$$\frac{\partial}{\partial a} \left[\frac{\Upsilon(J,a)}{a} \right] = \frac{\hat{J}^2}{J} \frac{d(\Psi j)}{dj} \frac{\partial j(J,a)}{\partial a}$$
$$= -\frac{\hat{J}^2}{2a} \frac{d(\Psi j)}{dj} \frac{\partial j(J,a)}{\partial J}$$
$$= -\frac{\partial}{\partial J} \left[\frac{\Upsilon(J,a)J}{2a^2} \right], \tag{29}$$

and also

$$\int_{0}^{\infty} \frac{\partial}{\partial J} \left[\frac{\Upsilon(J, a)J}{2a^{2}} \right] dJ = \frac{\Upsilon(J, a)J}{2a^{2}} \Big|_{0}^{\infty} = \frac{\hat{J}^{2}}{4a}.$$
 (30)

This yields

$$\partial_a U_2 = -\hat{J}^2 F'(0)/(2a).$$
 (31)

Integrating the latter with respect to *a* gives

$$U_2 = -[\hat{J}^2 F'(0)/2] \ln a + U_c.$$
(32)

The integration constant U_c , independent of a, equals $U_2(a = 1)$, so it can be written as

$$U_c = \int_0^\infty \left\{ 2\Upsilon(J,1)F'(0) + \frac{\hat{J}^2}{J} [F'(J) - F'(0)] \right\} dJ.$$

Using that $\Upsilon(J, 1) = \hat{J}\Psi(J/\hat{J})$, one gets

$$U_c = \hat{J}^2 \int_0^\infty \left[F'(J) - F'(0) Q\left(\frac{J}{\hat{J}}\right) \right] \frac{dJ}{J}.$$
 (33)

[Notice that, although determined by essentially nonlinear dynamics in the narrow vicinity of the resonance, $Q(J/\hat{J})$ nevertheless can affect the integrand on the thermal scale.] Finally, Eq. (24) is summarized as

$$U = -\left[\hat{J}^2 F'(0)/2\right] \ln a - \varkappa \sqrt{a} \hat{J}^3 F''(0) + U_c.$$
(34)

B. Nonlinear frequency shift

Using Eqs. (33) and (34), we can rewrite Eq. (21) as

$$\epsilon(\omega,k) + \frac{\omega_p^2}{2k^2} C_1 \ln a + \frac{\omega \omega_p^2}{k^3} \varkappa C_2 \sqrt{a} = 0.$$
(35)

Here, we introduced

$$\epsilon = 1 - \frac{m^2 \omega_p^2}{k^6} \int_0^\infty \left[F'(J) - F'(0) Q\left(\frac{J}{\hat{J}}\right) \right] \frac{dJ}{J}$$
(36)

and also

$$C_1 = (m/k)^2 F'(0), \quad C_2 = (m/k)^3 F''(0).$$
 (37)

In particular, when F'(0) = 0, the nonlinear part of Eq. (35) is small, yielding that the nonlinear frequency shift ω_{NL} is also small; hence,

$$\omega_{\rm NL} = -\left(\frac{\partial\epsilon}{\partial\omega}\right)^{-1} \frac{\varkappa \omega_p^2}{k^2} \sqrt{\frac{e\tilde{\varphi}_0}{m}} C_2. \tag{38}$$

Yet, at nonzero F'(0), the nonlinear part of Eq. (35) diverges logarithmically at small *a*; i.e., wave interaction with resonant particles has a strong effect on ω .

C. Comparison with existing theories

Equations (35)–(38) generalize the existing NDR for quasimonochromatic eigenwaves in plasmas with smooth distributions of particles,^{16,18,21,25–30} namely, as follows. First of all, notice that ϵ , serving as a *generalized linear dielectric function*, is a functional of the action distribution. Unlike the commonly used distribution of "unperturbed" velocities $f_0(v)$, which depends on how the wave was excited,²¹ F(J) is defined unambiguously; thus, the above equations hold for any excitation scenario. (Notice that finding F(J) itself is hence found to be a separate problem, which is also where the effect of the quasistatic field¹⁸ comes in.) Second, even if put in terms of $f_0(v)$, Eqs. (35)–(38) cover a wider class of particle distributions. The latter is seen as follows.

For example, consider a homogeneous wave developed with fixed *u* slowly from zero amplitude, so each *J* is conserved,³¹ even through trapping and untrapping.^{18,32–34} Then $F(J) = F_0(J)$, with index 0, henceforth, denoting the initial state. Yet, since there was no wave in that state, Eq. (20) applies, so each ℓ th derivative of $F_0(J)$ reads as

$$F_0^{(\ell)}(J) = (k/m)^{\ell} \Big[f_0^{(\ell)}(v^+) + (-1)^{\ell} f_0^{(\ell)}(v^-) \Big].$$
(39)

Let us use bars to denote limits $f_0^{(\ell)}(v \to u \pm)$ and introduce

$$\bar{f}_0^{(\ell)}(v) = \begin{cases} \bar{f}_0^{(\ell)}(u-), & v < u, \\ \bar{f}_0^{(\ell)}(u+), & v > u, \end{cases}$$
(40)

which is a piecewise-constant function equal to the left and right limits for v < u and v > u correspondingly; in particular,

$$C_1 = \bar{f}_0'(u+) - \bar{f}_0'(u-), \tag{41}$$

$$C_2 = \bar{f}_0''(u+) + \bar{f}_0''(u-). \tag{42}$$

Hence, ϵ rewrites as follows:

$$\epsilon = 1 - \frac{\omega_p^2}{k^2} \int_{-\infty}^{\infty} \frac{f'_0(v) - Q\bar{f}'_0(v)}{v - u} \, dv, \tag{43}$$

where $Q \equiv Q(|v/u - 1|)$. (Remarkably, contributing to ϵ are both passing and trapped particles.)

First, compare Eq. (43) with the usual expression,³⁵

$$\epsilon_{\rm L} = 1 - \frac{\omega_p^2}{k^2} \mathsf{P} \int_{-\infty}^{\infty} \frac{f_0'(v)}{v - u} dv, \tag{44}$$

where P denoting the principal value.³⁶ For smooth $f_0(v)$, our ϵ can be put in the same form as ϵ_L , because $P \int_{-\infty}^{\infty} (v-u)^{-1} Q \bar{f}'_0(v) dv = 0$. However, Eq. (43) is valid also when $f'_0(v)$ is discontinuous across the resonance, a case in which ϵ_L is undefined. This is because the integrand in Eq. (43) is finite (piecewise-continuous), so the integral converges absolutely rather than existing only as a principal value (like ϵ_L does). Second, for smooth $f_0(v)$, when $C_1 = 0$ and $C_2 = 2f''_0(u)$, Eq. (38) for ω_{NL} matches that in Ref. 21, including the coefficient, since $\varkappa \approx 0.544$ (Appendix). Yet, unlike the existing theory, our results apply just as well for arbitrary C_1 and C_2 , in which case $f_0(v)$ may not be smooth while F(J) is.

IV. BEAM-LIKE DISTRIBUTIONS

Suppose now that, in addition to a smooth distribution, near the resonance, there is a phase-space clump or a hole, further termed uniformly as a beam with $F_b(J) \ge 0$ and some average spatial density $n_b(J) \ge 0$. Since particles contribute additively to the right-hand side in Eq. (11), we immediately obtain

$$\epsilon(\omega,k) + \frac{\omega_p^2}{2k^2} C_1 \ln a + \frac{\omega\omega_p^2}{k^3} \varkappa C_2 \sqrt{a} + \mathfrak{B} = 0, \qquad (45)$$

where \mathfrak{B} is the beam-driven term,

$$\mathfrak{B} = -\frac{2\omega_b^2}{an_b\omega^2} \int_0^\infty GF_b(J) \, dJ,\tag{46}$$

with $\omega_b^2 = 4\pi n_b e^2/m$. [Unlike F(J), the distribution $F_b(J)$ is normalized to the beam density here.] For example, consider a beam of deeply trapped particles, so that $F_b(J) = n_b \delta(J)$. Since G(0) = -1, one then has

$$\mathfrak{B} = 2\omega_b^2 / (a\omega^2). \tag{47}$$

Hence, when the beam nonlinearity dominates,

$$(\omega, k) + 2\omega_b^2 / \omega_E^2 = 0.$$
 (48)

For instance, within the model,³⁷

 ϵ

$$\epsilon(\omega,k) = 1 - \frac{\omega_p^2}{\omega^2 - 3k^2 v_T^2},\tag{49}$$

where v_T is the thermal speed, one thereby gets the following dispersion relation:

$$\omega^2 = \frac{\omega_p^2}{1 + 2\omega_b^2/\omega_E^2} + 3k^2 v_T^2,$$
(50)

in agreement with Ref. 15. Strictly speaking, the sinusoidalwave approximation required for this result to hold applies only at $\omega_b^2/\omega_E^2 \ll 1$,¹⁵ so one can further approximate Eq. (50) as

$$\omega^2 = \omega_p^2 (1 - 2\omega_b^2 / \omega_E^2) + 3k^2 v_T^2, \tag{51}$$

in agreement with Ref. 38. Since Eq. (49) implies $kv_T \ll \omega$, one has $\omega \approx \omega_p$, so the above result also rewrites as

$$\omega^2 = \omega_{\rm L}^2 - 2\omega_b^2/a,\tag{52}$$

where $\omega_{\rm L}^2 = \omega_p^2 + 3k^2 v_T^2$ corresponds to the linear limit. In particular, notice that $\omega_{\rm NL} = \mathcal{O}(a^{-1})$, with $\omega_{\rm NL} < 0$ for a clump and $\omega_{\rm NL} > 0$ for a hole. (Remember that $\omega_b^2 \propto n_b$ may have either sign.)

Now let n_b itself depend on the wave amplitude. For example, a Van Kampen mode would have ω_b^2/a independent of *a*, yielding that trapped particles result in a *linear* frequency shift, in agreement with the original linear theory.³⁹ Also, let us consider the case when F_b is constant across the trapping width, i.e., $F_b(J) = \overline{F}_b \Theta(J) \Theta(J_* - J)$, with Θ being the Heaviside step function. Hence, $n_b = \overline{F}_b J_*$ and

$$\mathfrak{B} = \frac{2\omega_b^2}{a\omega^2} \frac{\Psi(j_*)}{j_*} = \frac{8\bar{F}_b\hat{J}}{3\pi n\sqrt{a}} \frac{\omega_p^2}{\omega^2},\tag{53}$$

where we used $\Psi(j_*) = 4/(3\pi)$ (Appendix). Hence, the approximate formula is obtained, like above,

$$\omega^2 = \omega_{\rm L}^2 - \frac{8\omega_p^2 \tilde{J}F_b}{3\pi\sqrt{an}}.$$
(54)

Equation (54) also matches the result found previously, e.g., in Ref. 17.

V. DISCUSSION

The above analysis shows that the power index β in the scaling for the nonlinear frequency shift $\omega_{\rm NL} \propto a^{\beta}$ depends critically on the properties of F(J) at small J. The commonly anticipated $\beta = 1/2$ is realized when F(J) is smooth, an example being the case when the wave develops adiabatically from zero amplitude. However, perturbations to F on the trapping-width scale J_* result in additional nonlinearities which yield $\beta < 0$. In particular, this affects the sign of $\partial_a \omega_{\rm NL}$, which can be decisive for the wave stability (Paper III). Even more generally, describing such nonlinear kinetic effects in a fluid-like manner^{30,40,41} is possible, in fact, only in special cases (e.g., when a(t) is monotonic and u is fixed, as in Ref. 42), because $\omega_{\rm NL}$ is not a local function of a.

For example, consider a driven nonlinear mode with adiabatic chirp imposed.^{10,43–45} Due to the chirp, the wave phase velocity changes, carrying trapped particles along because of the autoresonance; then $F(J < J_*)$ remains intact. Yet $F(J > J_*)$ is changing, because passing particles do not conserve their actions as defined here.³¹ Hence, even for an initially smooth F(J), the chirp generally produces discontinuity of the action distribution at $J = J_*$, thereby changing the scaling for $\omega_{NL}(a)$. Thus, accounting for nonlocality in $\omega_{NL}(a)$ would be particularly important for trapped-particle waves whose phase velocity varies with time, modulated waves (Paper III) being an example.

Finally, let us also consider effects that can be driven by dissipation, in case when plasma is collisional. Since ω changes rapidly with small *a* in the presence of a phase-space clump or a hole, slow decay of *a* will cause frequency downshifting or upshifting, correspondingly. Yet, since β depends on how localized $F_b(J)$ is, another effect is anticipated, namely, as follows. Notice that a friction drag (say, proportional to the particle velocity) can cause condensation of the trapped distribution near the bottom of the wave potential trough.⁴⁶ Hence, peaking of F(J) can occur, and β can decrease gradually down to minus one. This represents a frequency sweeping mechanism additional to those considered in Refs. 47–50.

VI. CONCLUSIONS

In this paper, we expand on the results that were previously reported in our brief letter,¹¹ performing a systematic study of kinetic nonlinearities in NDR of an adiabatic sinusoidal Langmuir wave in collisionless plasma. Specifically, a general NDR is derived in a nondifferential form, allowing for an arbitrary distribution of trapped electrons. The linear dielectric function is generalized, and the nonlinear kinetic frequency shift $\omega_{\rm NL}$ is found analytically as a function of the wave amplitude a. Smooth distributions yield $\omega_{\rm NL} \propto \sqrt{a}$, as usual. However, beam-like distributions of trapped electrons result in different power laws, or even a logarithmic nonlinearity, which are derived as asymptotic limits of the same dispersion relation. Such beams are formed whenever the phase velocity changes, because the trapped distribution is in autoresonance and thus evolves differently from the passing distribution. Hence, even adiabatic $\omega_{\rm NL}(a)$ is generally nonlocal.

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APPENDIX: AUXILIARY FUNCTIONS

Here, we summarize explicit expressions for the special functions used in this paper.⁵¹ The analysis presented in the main text does not rely on these expressions *per se*, yet those are needed to calculate the numerical values of j_* , $\psi(j_*)$, and \varkappa .

Specifically, the normalized action j(r) reads as

$$j(r) = \frac{4}{\pi} \times \begin{cases} (r-1)\mathsf{K}(r) + \mathsf{E}(r), & r < 1, \\ r^{1/2}\mathsf{E}(r^{-1}), & r > 1 \end{cases}$$
(A1)

[Fig. 4(a)], where K and E are the complete elliptic integrals of the first and second kind, 5^{2}



FIG. 4. (a) Normalized action j(r), Eq. (A1). (b) Weight function g(r), Eq. (A6). The dashed lines denote the separatrix r = 1 and also the asymptote of g(r).

$$\mathsf{K}(\zeta) = \int_0^{\pi/2} (1 - \zeta \, \sin^2 \theta)^{-1/2} \, d\theta, \tag{A2}$$

$$\mathsf{E}(\zeta) = \int_{0}^{\pi/2} \left(1 - \zeta \, \sin^2 \theta\right)^{1/2} d\theta, \tag{A3}$$

with the following properties:

$$K(0) = E(0) = \pi/2, \qquad E(1) = 1.$$
 (A4)

In particular, one thereby gets $j(1) = 4/\pi$ (which can be found as well without introducing elliptic integrals), and

$$j'(r) = \frac{2}{\pi} \times \begin{cases} \mathsf{K}(r), & r < 1, \\ r^{-1/2} \mathsf{K}(r^{-1}), & r > 1, \end{cases}$$
(A5)

also yielding the oscillation period, with Eq. (18). From Eq. (15), one obtains then

$$g(r) = \begin{cases} 1 - 2\mathsf{E}(r)/\mathsf{K}(r), & r < 1, \\ 2r - 1 - 2r\mathsf{E}(r^{-1})/\mathsf{K}(r^{-1}), & r > 1 \end{cases}$$
(A6)

[Fig. 4(b)]. Further, the function $\Psi(j)$ is calculated as $\Psi(j) = \psi(r(j))$, with Eq. (22) yielding

$$\psi(r) = -\int_0^r g(\bar{r}) j'(\bar{r}) \, d\bar{r}.$$
 (A7)

Hence, Eqs. (A5) and (A6) result in



FIG. 5. (a) Function $\psi(r)$, Eq. (A8). (b) Function q(r), Eq. (A9). The dashed lines denote the separatrix r = 1 and also the asymptote of q(r).

$$\begin{split} \psi(r) &= \frac{4}{3\pi} \\ &\times \begin{cases} (1-r)\mathsf{K}(r) + (2r-1)\mathsf{E}(r), & r < 1, \\ r^{1/2}[2(1-r)\mathsf{K}(r^{-1}) + (2r-1)\mathsf{E}(r^{-1})], & r > 1, \end{cases} \end{split}$$

and, in particular, $\psi(1) = 4/(3\pi)$ [Fig. 5(a)].

Finally, Q(j) is calculated as Q(j) = q(r(j)), with

$$q(r) = 1 - 2j(r)\psi(r) \tag{A9}$$

[Fig. 5(b)]. Hence,

Λ

$$\varkappa = \int_0^\infty [1 - 2j(r)\psi(r)]j'(r)\,dr \approx 0.544.$$
 (A10)

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