Computational Study of Magnetic Islands in the W7-X and NCSX Stellarators

M. Drevlak, C. Nührenberg

IPP-Euratom Association, MPI für Plasmaphysik, 17491 Greifswald, Germany A.H. Boozer, S.R. Hudson, D. Monticello, A. Reiman Princeton Plasma Physics Laboratory, PO Box 451, Princeton NJ 08543, USA

The magnetic fields of stellarators and, in general, all not strictly axisymmetric toroidal fusion devices exhibit magnetic islands. They strongly influence the confinement properties of a fusion plasma and are exploited in divertor design. So, the existence and the structure of the magnetic islands are an important issue in configuration design. Codes that determine finite-plasma- β stellarator equilibria while fully accounting for their island structures exist (HINT [1], PIES [2]) and have been applied to the NCSX [3] and W7-X [4] stellarators. In the NCSX standard high- β scenario low-order rotational-transform values exist inside the plasma [5]. Figure 1 shows a poloidal cut of an NCSX case at $\langle \beta \rangle = 4\%$ with $0 < \theta < \pi$ and the normalized radius r < 1 of the PIES background coordinates. Several 3/m and 6/m is-

lands can be seen, e.g. 3/5 at $r \approx 0.75$ and 3/6 at $r \approx 0.6$ in the







3-periodic device.

Figure 2: PIES for W7-X at $\langle \beta \rangle \approx 5\%$.

For the W7-X stellarator the situation is different: In the standard high- β case no low-order rationals exist inside the plasma [6]. However, in the W7-X configuration space a case exists with rotational transform $\iota = 5/6$ inside the plasma at high plasma- β , as has been computationally demonstrated (see Fig. 2). The PIES code has been used to compute a free boundary equilibrium at $\langle \beta \rangle = 5\%$. In MFBE calculations preceding the PIES computation, a configuration of coil currents similar to one previously successful at $\beta = 4\%$ [6] was selected. This configuration employs the auxiliary coils, energised so as to maximise the plasma volume by shifting the magnetic axis inward and adjusting the *t* profile. The equilibrium found in the subsequent PIES analysis exhibits a low order resonance (5/6) inside the confinement region (see Fig. 2). The plasma volume is reduced to ~ $15m^3$ at this β . The rotational transform ranges between 5/6 and 5/5 in the vacuum field, it decreases with increasing plasma- β .

An alternative approach to the assessment of magnetic islands in finite- β stellarator equilibria has been developed with the method of perturbed equilibria [7]. Since a perturbed equilibrium represents a small deviation from an equilibrium ideal MHD stability theory and, hence, ideal MHD stability codes, e.g. the CAS3D code [8], can be used to determine



Figure 3: CAS3D results for a W7-X case at $\langle \beta \rangle = 5\%$.

a perturbed equilibrium. Discontinuities of the normal displacement at rational surfaces indicate surface currents which are used to model islands. The strength of such a surface current can be used to estimate the corresponding island width. The augmented CAS3D code was applied to equilibria neighbouring the cases of Figs. 1 and 2. A first result for W7-X is shown in Fig. 3 for a case with the rational rotational transform, t = 5/6, at approximately half of the total enclosed toroidal flux. The CAS3D analysis predicts an island width of ≈ 0.034 m which is comparable to the PIES code finding.

References

- T. Hayashi. In *Theory of Fusion Plasmas Chexbres 1988*, volume EUR 12149 EN, page 11, Bologna, 1989. Società Italiana di Fisica.
- [2] A. H. Reiman and H. Greenside. Comput. Phys. Commun., 43:157, 1986.
- [3] M. C. Zarnstorff, L. A. Berry, A. Brooks, et al. Plasma Phys. Control. Fusion, 43:A237, 2001.
- [4] G. Grieger, et al. In Plasma Physics and Controlled Nuclear Fusion Research 1990, volume 3 of Nucl. Fusion Suppl., page 525, Vienna, 1991. International Atomic Energy Agency.
- [5] S. R. Hudson, et al. Phys. Rev. Lett., 89:275003, 2002.
- [6] M. Drevlak, et al. Nucl. Fusion, 45:731, 2005.
- [7] A. H. Boozer. Phys. Plasmas, 6:831, 1999.
- [8] C. Nührenberg and A. H. Boozer. Phys. Plasmas, 10:2840, 2003.