

Multiregion Relaxed MHD toroidal states with flow

Tokamak



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Plan of talk

- Desiderata: Simple MHD model that can reduce to ideal MHD when appropriate but allows reconnection when not
- Dynamical MRxMHD: Force-free magnetic field⇔Euler fluid in each relaxation region
- Contrast 2 cases: Laminar flow (axisymmetry);
 Chaotic streamlines (non-axisymmetry)
- Preliminary numerical implementation

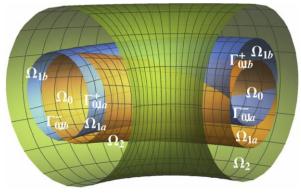


Basic ideas

Confinement is maintained by thin transport barriers (interfaces) $\Gamma_{i,j}$ dividing the plasma into toroidal sub-regions Ω_i

 \Box Force balance across *interfaces* $\Gamma_{i,j}$ allows stepped pressure





Interfaces may be separatrices of magnetic islands / Kelvin's cat's eyes in fluid Within the sub-regions Ω_i, use force-free magnetic fields, so no jXB force — field and fluid are decoupled

□ Scope of applicability

- "smart finite element" discretization for ideal-MHD 2D equilibrium calculations (incl. flow)
- new, quasi-relaxed, MHD allowing islands & chaos in 3D equilibrium and dynamics



MRxMHD theory

MRxMHD: M stands for **Multi-region** (aka waterbag) Rx stands for **Relaxed**; ...D stands for **Dynamics**

- \Box Full ideal-MHD constraints *apply only within the interfaces* $\Gamma_{i,i}$
- Within the sub-regions Ω_i use generalized Taylor relaxation theory – relax nearly all the ideal-MHD constraints except for conservation of macroscopic magnetic helicity, entropy, and microscopic mass
- □ Cross-helicity not constrained, so fluid and magnetic field couple *only at interfaces*, which are current/vortex sheets
- □ Generalize minimum energy equilibrium approach by instead extremizing the MHD action (Hamilton's Principle)
- Details in Dewar et al 2015



Euler-Lagrange Equations within Ω_i

• Mass conservation (microscopic constraint)

$$\frac{\partial \rho}{\partial t} = -\boldsymbol{\nabla} \boldsymbol{\cdot} (\rho \mathbf{v})$$

• $\delta p \Rightarrow$ Isothermal equation of state

$$p = \tau_i \rho$$
 (N.B. $\tau_i = C_{\mathrm{s}i}^2$)

δA ⇒ Beltrami equation
∇×B = μ_iB (N.B. ⇒ j×B = 0)
ξ ⇔ Momentum equation (*Euler fluid*)

$$p\left(\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v}\right) = -\nabla p$$

Steady axisymmetric toroidal flow

Piecewise-constant *vorticity* $\boldsymbol{\omega} \equiv \nabla \times \mathbf{v}$:

 $\boldsymbol{\omega}_i = 2\Omega_i \mathbf{e}_Z \iff \mathbf{v} = \Omega_i R \mathbf{e}_\phi \implies \boldsymbol{\omega}_i \times \mathbf{v} = -2\Omega_i^2 R \mathbf{e}_R = -\nabla v_\phi^2$

• Check Ideal Ohm's Law solvability condition for Φ :

$$\nabla \times (\mathbf{v} \times \mathbf{B}) = \Omega_i R \left(\nabla \cdot \mathbf{B} - \frac{1}{R} \frac{\partial B_{\phi}}{\partial \phi} \right) \mathbf{e}_{\phi} = 0 \quad \mathbf{v} \square \mathbf{v} \square$$

So rigid-body flow is compatible with ideal MHD for any axisymmetric **B**, *including* MRxMHD's force-free Beltrami field!

• Steady toroidal Euler flow momentum equation:

$$\boldsymbol{\omega} \times \mathbf{v} + \nabla \left(\frac{\nu_{\phi}^2}{2} + \ln \frac{\rho}{\rho_0} \right) = 0$$

• Gives Bernoulli equation $-\frac{v_{\phi}}{2} + \ln \frac{\rho}{\rho_0} = 0$: in agreement with, e.g., McClements & Hole 2010's *ideal MHD* result.



"Relaxed" non-axisymmetric flow

Static solution of Compressible Euler Fluid eqs in arb. Ω (Sato & Dewar arxiv):

- Dot momentum equation $\rho \mathbf{v} \cdot \nabla \mathbf{v} = -\tau \nabla \rho$ with \mathbf{v}/ρ and integrate to give *Bernoulli* equation $v^2/2 + \tau \ln(\rho/\rho_0) = \text{const}$ on each flow line.
- Only solution valid for *arbitrarily chaotic flow* within Ω is (with suitable choice of global constant ρ_0) $\rho = \rho_0 \exp(-v^2/2\tau)$
- Gives nonlinear Beltrami equation: $\nabla \times \mathbf{v} = \alpha_0 \exp(-v^2/2\tau)\mathbf{v}$



- Relaxed (Beltrami) fluid solution: vorticity parallel to flow
- Rigid toroidal flow solution: vorticity perpendicular to flow
- Toroidal kinetic energy terms in Bernoulli equations have opposite signs in the two solns.
- Unlike rigid toroidal flow, relaxed fluid does not in general satisfy ideal Ohm's Law solvability condition ∇ × (v × B) = 0.



- Are there steady 3-D MRxMHD solution that satisfy ∇ × (v × B) = 0?
- Can fluid relaxation be related to inverse cascade theory?
- Do adjacent non-axisymmetric layers with different mean velocities necessarily exchange angular momentum or is there a d'Alembert's paradox? — Do nonaxisymmetric steady flow equilibria *exist* in general, or only oscillatory solutions?



Preliminary SPECF implementation I

Starting equation:

 $\nabla \times \boldsymbol{u} = 0$ (Vorticity-free) $\rho = \rho_0 e^{-u^2/2\tau}$ (Bernoulli relationship) $\nabla \cdot \rho \boldsymbol{u} = 0$ (Continuity)

We get that

$$\nabla \cdot e^{-\frac{u^2}{2\tau}} (\nabla f + \boldsymbol{V}_0) = 0, \qquad *$$

where f is a single valued per function and

 $\boldsymbol{u} = \nabla f + \boldsymbol{V}_0 = \nabla f + \psi_{tV} \nabla \theta + \psi_{pV} \nabla \xi.$

This ensures that the toroidal and poloidal loop integral $\oint \mathbf{u} \cdot d\mathbf{l}$ equal to ψ_{tV} and ψ_{pV} , the toroidal and poloidal vorticity flux, respectively.

Using the Chebyshev-Fourier reps in SPEC, writing f into

$$f = \sum_{i,l} f_{e,i} T_{l,i}(s) \sin(m_i \theta - n_i \xi),$$

casting the equation * into matrix form

$$\underline{A(\boldsymbol{f}_{n-1})} \cdot \boldsymbol{f}_n = \underline{B(\boldsymbol{f}_{n-1})} \cdot \begin{pmatrix} \psi_{tV} \\ \psi_{pV} \end{pmatrix},$$

in which matrix <u>A</u> and <u>B</u> depend on <u>f</u> due to the Boltzmann exponential factor, and are calculated iteratively by taking the last solution of f until converged.

The boundary condition is $\boldsymbol{u} \cdot \boldsymbol{n} = u_s = 0.$

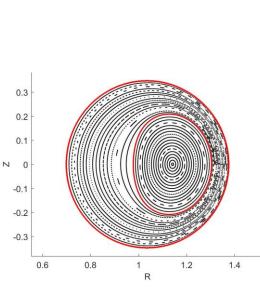
Force Balance:
$$[[p_0 e^{-\frac{u^2}{2\tau}} + \frac{1}{2}B^2]] = 0$$

Preliminary SPECF implementation 2

Case1: Axisymmetric test case (this is some random geometry...)

If β is low $\sim r^2/R^2$, the exponential factor to pressure does not influence force balance So we need high $\beta \sim r/R$ (in our case $\beta \sim 1/3$)

Constrain rotational transform (Helicity is not avaliable currently)

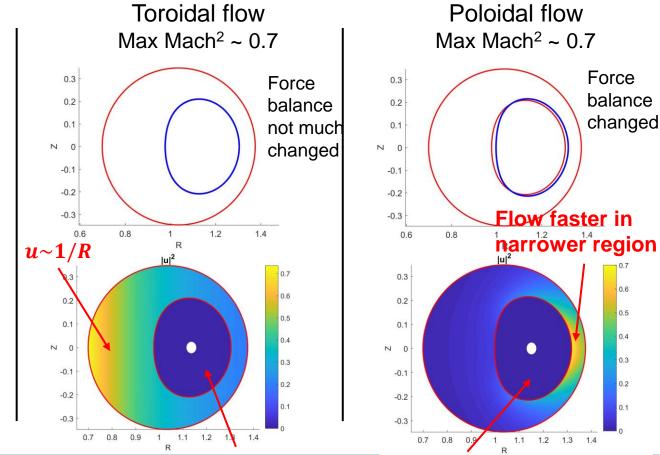


No flow

Australian

National University

I know I need to add many interfaces to match Michael's solution...



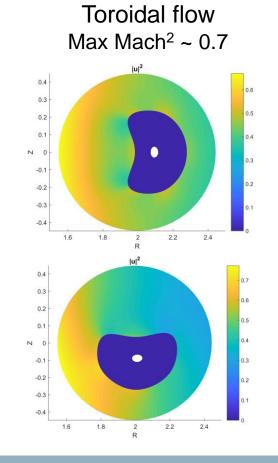
No flow at core because illly defined vorticity flux

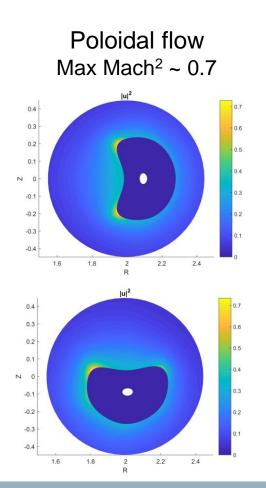


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Case 2: Reverse field pinches, single helical axis (SHAx) Very small beta, flow plays no role in force balance

No flow 0.5 (c)-0.5 1.5 2.5 2 (d)1.5 2 *R* [m] 2.5



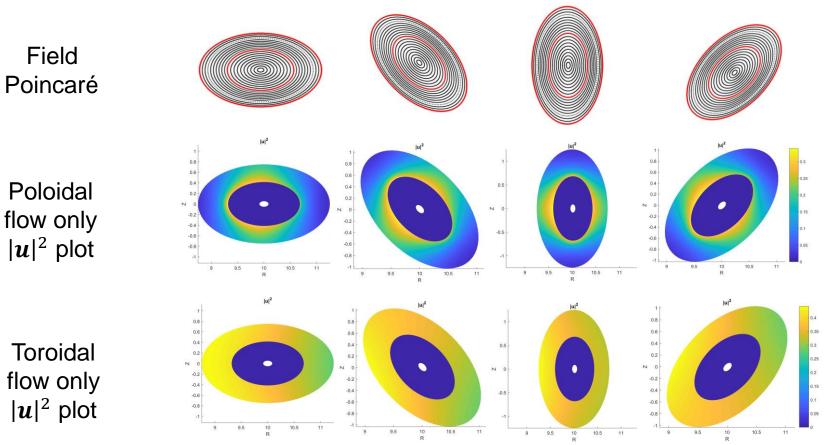


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Case 3: Stellerator test case (toroidal periodicity = 5) Note: zero β case No viscosity is used. There is d'Alembert's paradox in our case. One need to consider interface as a boundary layer (not finished).

But regardless of the lack of boundary layer, we get solutions.





- Have shown that MRxMHD is compatible with ideal MHD for *axisymmetric* toroidal flow equilibria
- Have found that the most general nonaxisymmetric "relaxed Euler flow" equilibria cannot reduce to the axisymmetric toroidal flow equilibria
- Have implemented a preliminary version of the SPEC code with flow (SPECF)
- Have enunciated some open questions that need to be addressed



Abstract

Multiregion Relaxed MHD toroidal states with flow <u>R.L. Dewar</u>,¹ Z.S. Qu,¹ N. Sato,² S.R. Hudson,³ M.J. Hole¹ ¹Mathematical Sciences Institute, The Australian National University, Canberra

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The action-based formulation¹ of Multiregion Relaxed MHD (MRxMHD) encompasses both steady-flow statics, and dynamics on a slower timescale than Taylor relaxation. We consider the case of a toroidal plasma laminated into multiple nested annular toroidal relaxation regions, separated by interfaces supporting current sheets. Unlike ideal MHD, Taylor relaxation allows reconnection at resonant surfaces to occur within these regions. However, the physical applicability of the model depends on the interfaces between them being ideal, i.e. *stable* against reconnection for times much longer than the relaxation timescale.

It has been postulated² that plasma flow may stabilize such current sheets even if they occur on surfaces that resonate with boundary perturbations in 3D geometries such as stellarators, or tokamaks with resonant magnetic perturbation (RMP) coils. This motivates the extension, now under development, of the 3D-MRxMHD-based *equilibrium* code SPEC³ to allow plasma flow with reasonably general flow profiles. However, it is not clear⁴ that stationary 3D states with other than rigid-rotation flow exist, motivating development of a 3D MRxMHD *initial value* code to model oscillatory states and nonlinear instabilities.

The formulation of Ref. 1 describes the plasma in each region as an ideal Euler fluid, which is too general for practical purposes as it allows all the turbulent complexity of such a fluid. This motivates developing a Taylor-like relaxation model⁵ for fluids, based on minimizing total energy with constant mass, entropy and fluid helicity (or, equivalently, minimizing fluid helicity at constant mass, entropy and energy). This leads to a compressible Beltrami equation, $\nabla \times \mathbf{v} = \alpha_0 \exp(-v^2/2\tau)\mathbf{v}$, where α_0 and τ are constant in each region, τ being the square of the isothermal sound speed in that region. The simplest case is $\alpha_0 = 0$, i.e. the flow has *zero vorticity*, but, because our relaxation regions are not simply connected, non-trivial rotation profiles can still be treated.

References:

- 1. R.L. Dewar, et al., J. Plasma Phys., 81, 515810604-1-22, (2015).
- 2. R.L. Dewar, S.R. Hudson et al., Phys. Plasmas, 24, 042507-1-18, (2017).
- 3. S.R. Hudson, R.L. Dewar et al., Phys. Plasmas 19, 112502-1–18, (2012).
- 4. G.R. Dennis, S.R. Hudson, R.L. Dewar and M.J. Hole, Phys. Plasmas **19**, 042501-1–9, (2014).
- 5. N. Sato and R.L. Dewar, *Relaxation of Compressible Euler Flow in a Toroidal Domain* <u>https://arxiv.org/pdf/1708.06193.pdf</u>.



Derived from MRxMHD Lagrangian

kinetic energy – MHD potential energy + Lagrange multiplier constraint terms:

• MHD Lagrangian density in region *i*

$$\mathcal{L}^{\text{MHD}} = \rho \frac{v^2}{2} - \frac{p}{\gamma - 1} - \frac{B^2}{2\mu_0}$$

• Constrained Lagrangian in region *i*

$$L_i = \int_{\Omega_i} \mathcal{L}^{\text{MHD}} dV + \tau_i (S_i - S_{i0}) + \mu_i (K_i - K_{i0})$$

• Helicity and entropy macroscopic invariants

$$K_i \equiv \int_{\Omega_i} \frac{\mathbf{A} \cdot \mathbf{B}}{2\mu_0} \, dV \qquad S_i \equiv \int_{\Omega_i} \frac{\rho}{\gamma - 1} \ln\left(\kappa \frac{p}{\rho^{\gamma}}\right) dV$$