#### Containing a star on earth: the promise of fusion energy



#### Walter Guttenfelder Princeton Plasma Physics Laboratory (PPPL)

Portland State University, Physics Dept. seminar Nov. 5, 2018





#### OUTLINE

- Nuclear fusion for energy (what & why)
- Plasmas
- Magnetic confinement of fusion plasma
- Achievements in fusion energy research
- Challenges and opportunities

#### Many quality-of-life metrics correlated with energy use





- Similar trends for
  - UN human development index
  - Income, average wage
  - Literacy, years in school
  - Reduced child mortality

Free material from www.gapminder.org

#### Many quality-of-life metrics correlated with energy use



- Similar trends for
  - UN human development index
  - Income, average wage
  - Literacy, years in school
  - Reduced child mortality

 Increased energy consumption in the industrial era has increased CO2 emissions

### Increased CO2 emissions $\rightarrow$ increased CO2 concentration $\rightarrow$ elevating global temperatures

Global warming relative to 1850-1900 (°C)



Source: IPCC Special Report on Global Warming of 1.5°C (Oct. 6, 2018)

 Climate concerns & growing global energy demand drives pursuit of a portfolio of alternative / renewable, non-carbon energy sources:
– solar, wind, nuclear fission, hydroelectric, geothermal ... and nuclear

fusion (this talk)

### Nuclear Fusion: Energy release occurs due to fusing two small nuclei



### Nuclear Fusion: Energy release occurs due to fusing two small nuclei



#### Opposite of <u>nuclear fission</u> that powers today's "nuclear" reactors

• Splitting large atoms also leads to energy release



### Energetics governed by binding energies / strong nuclear force



#### Energetics governed by binding energies / strong nuclear force

• Gain for fusion can be much larger than fission (both are far larger than chemical reactions)



#### Fusion with deuterium & tritium (D-T) is easiest

#### **Potential Candidates:**

- D + D  $\rightarrow$  <sup>3</sup>He + n + 3.27 MeV (50%)
- $D + D \rightarrow T + p + 4.03 \text{ MeV} (50\%)$
- $D + T \rightarrow {}^{4}He + n + 17.59 \text{ MeV}$
- $D + {}^{3}He \rightarrow {}^{4}He + p + 18.3MeV$

- Optimal temperature for D-T "thermonuclear" fusion ~ 150 M °C (15 keV)
- ~10 × hotter than the center of the sun!



Deuterium is abundant in seawater, but tritium half-life ~12.3 years (only ~50 kg on the planet)  $\rightarrow$  breed tritium from lithium

$$^{7}Li + n \rightarrow He + T + n - \text{energy}$$
  
 $^{6}Li + n \rightarrow He + T + \text{energy}$ 

- Reserves in South America
- Potential abundant source of Li also in seawater (via desalination / dialysis)

D + T → <sup>4</sup>He + n + 17.6 MeV + Li + n → <sup>4</sup>He + T + 2.3 MeV  $\approx$  D + Li → 2 <sup>4</sup>He + 20.0 MeV

#### Why study fusion energy research?

- No carbon emission
- Fuel is abundant (~thousands of years)



#### Why study fusion energy research?

- No carbon emission
- Fuel is abundant (~thousands of years)
- Inherently safe only grams (<minute) of fuel in the device</li>

no melt down/runaway concerns

- Very little (and short lived, low-level) radioactive material compared to nuclear fission
- Compared to non-carbon renewables (solar, wind) fusion is compact and continuous (not intermittent)
- Disadvantages: Hard to do!

#### Must overcome repulsive electrostatic force to fuse atomic nuclei

Force between two charged particles increases as they get closer





**Temperatures must be ~150 million degrees Celsius**  $\rightarrow$  no longer a gas, but a <u>plasma</u> (Core of the sun ~15 million C)



### What is plasma?

### Plasma: a gas of charged particles (the fourth state of matter)



• Plasma behaves qualitatively different than neutral gas due to collective (Coulomb) interactions & interactions with electric (E) & magnetic (B) fields  $\vec{F} = q \left( \vec{E} + \vec{v} \times \vec{B} \right)$ 

#### 99% of (known) matter in universe is plasma

 Sun, stars, interstellar and intergalactic medium account for most mass and are largely plasma

#### Ionized gas in the Milky Way







#### Numerous examples of plasmas on or near earth





neon signs



tv



satellite plasma thrusters

#### semiconductor processing



aurora



#### Plasmas exist over broad range of density and temperature → wide range of physics phenomenon!



# How do we create and contain a hot plasma on earth?

#### **Recipe to create a fusion plasma**

- 1. Establish an appropriate magnetic field
- 2. Inject appropriate gases (in a container at vacuum pressure)
- 3. Heat the gases

### Charged particles experience Lorentz force in a magnetic field → gyro-orbits

$$\vec{F} = q\left(\vec{E} + \vec{v} \times \vec{B}\right)$$

Magnetic force acts perpendicular to direction of particle
Particles follow circular gyro-orbits



#### Magnetic field confines particles away from boundaries

For a 5 Tesla magnetic field, 100 million C plasma

ion radius ~ 3 mm 1-2 meter electron radius ~ 0.05 mm device size

No magnetic field





#### Magnetic field confines particles away from boundaries



#### Solution: bend the field into a donut-shaped torus



### But toroidicity leads to vertical drifts from ∇B & curvature



 $\tau_{loss} \sim 5$  ms from vertical drifts (B~5 T, R~5 m, T~15 keV)

### Even worse, charge separation leads to faster E×B drifts out to the walls



 $\tau_{\text{loss}} \sim \mu s$  from E×B drifts (due to charge separation from vertical drifts)

### Solution: need a helical magnetic field for confined particle orbits



### Helical B field carries plasma from "bad curvature" region to "good curvature" region



Similar to how honey dipper prevents honey from dripping

#### Equilibrium establishes closed, nested magnetic "flux surfaces"

### Governed by magnetic hydrodynamic (MHD) force balance: $J \times B = \nabla P$



## How do we create the helical magnetic field?

#### The Tokamak (Russion acronym ~ "toroidal chamber with magnetic coils")

 Toroidal field from external coils + poloidal field from plasma current → helical field



#### At Princeton Plasma Physics Lab (PPPL): National Spherical Torus Experiment-Upgrade (NSTX-U)



#### At PPPL:

#### **National Spherical Torus Experiment-Upgrade (NSTX-U)**



#### At PPPL: National Spherical Torus Experiment-Upgrade (NSTX-U)




# Solenoid for inductive current drive

### **Toroidal field coils**



### **Shaping coils**





### Vacuum vessel

### Plasma current induced through solenoid via transformer action (Faraday's law of induction)





## Plasma current induced through solenoid via transformer action (Faraday's law of induction)



### Video of NSTX-U plasma



## Evidence of helical field seen in visible images of edge instabilities

MAST tokamak (UK)



### Plasma current induced through solenoid via transformer action (Faraday's law of induction)



### Stellarator concept uses complex 3D coils to generate helical magnetic field without plasma current

W7-X stellarator (Germany)



#### Trading engineering complexity for inherent steady-state operation

More complicated BUT more degrees of freedom for performance optimization



### We've created a magnetically confined plasma – how do we heat it?

### Mini particle accelerators (<u>Neutral Beam</u> <u>Injectors</u>) are used to heat the plasma



### Mini particle accelerators (<u>Neutral Beam</u> <u>Injectors</u>) are used to heat the plasma



### Microwave heating is also used (works similar to microwave ovens)



#### **RF** antenna

### Have achieved sufficient temperatures!

#### **Tokamak Fusion Test Reactor (PPPL, 1982-1997)**





### NBI & RF also used to drive plasma current as part of 100% non-inductive scenarios

Total plasma current = self-driven bootstrap current + NB current drive + RF current drive



# So what else do we need to make fusion <u>energy</u> a reality?

### **Power balance in a fusion reactor**



### **Power balance in a fusion reactor**

Typically requires a fusion gain Q >> 1 to account for thermal conversion, heating efficiency, cyro, plant, ...



### Fusion gain depends on the "triple product" $nT\tau_E$



### Fusion gain depends on the "triple product" $nT\tau_{\text{E}}$



### Fusion gain depends on the "triple product" $nT\tau_E$



# Confinement time is a measure of how well insulated the plasma is from the surrounding boundary



 $\tau_E = \frac{\text{stored energy}}{\text{rate of energy loss}}$ 

#### **For ignition (a self-sustaining, "burning plasma")** Q ~ $p \cdot \tau_F > 8$ atm·s (at ~150 million C)

p ~ 2-4 × atmospheric pressure  $\tau_{\rm E}$  ~ 2-4 seconds

#### Have come very close to plasma "break-even" (Q=1)



#### TFTR (PPPL, 1994)

10.7 MW fusion power 46 MW heating power Q=0.23

JET (UK, 1997) 16.1 MW fusion power 22 MW heating power Q=0.7

### Next step: ITER is being built to study "burning plasmas"

- Goal: 500 MW fusion power using 50 MW heating power
  - →large fusion gain Q = 10

<u>Seven partners</u> China, EU, India, Japan, Korea, Russia, US







#### Inside the tokamak pit (https://www.iter.org/album/construction)



### One toroidal field coil



### Not a bad place to visit!

#### Lavender of Senanque



#### **Ocre cliffs of Rousillon**



#### Les Beaux de Provence



#### Wine is good and not too expensive



# ITER will address a number of reactor relevant issues in an integrated fashion

- Demonstrate 500 MW fusion power for 400 sec, with large fusion power gain (Q=10)
- Study "burning plasma" regime ( $P_{\alpha} > P_{heat,ext}$ ) with selfheating via energetic particles (3.5 MeV  $\alpha$ 's)
- Test tritium breeding
- Demonstrate safety characteristics of fusion device (has already obtained nuclear licensing)
- ... (and more)

### ITER & next-step power plant projections are big (\$\$)

# What can we do to improve on this?

# Triple product (fusion gain) depends on stability, engineering and energy confinement

$$Q \sim nT\tau_E \sim \beta \cdot B^2 \cdot \tau_E$$

- Plasma beta β = nT/(B<sup>2</sup>/2µ<sub>0</sub>) = plasma / magnetic pressure
  → limited to β~5-10% by macroscopic (MHD) stability constraints (e.g. need to avoid "disruptions")
- Magnetic field strength B limited by superconductor technology (B<sub>crit</sub>, J<sub>crit</sub>) & mechanical stress limits
- $\tau_E$  limited by energy loss  $\rightarrow$  dominated by turbulence

## Turbulence in the plasma can (unfortunately) be very efficient at flushing out energy



## Turbulence in the plasma can (unfortunately) be very efficient at flushing out energy



# Why does turbulence develop in tokamaks?

## Analogy for turbulence in tokamaks - density gradient in the presence of gravity

- Higher density on top of lower density, with gravity acting downwards (Rayleigh-Taylor instability)
- Any small perturbation becomes unstable
- Convection mixes regions of different density

gravity density/pressure





### Inertial (centrifugal) force in toroidal field acts like an effective gravity



### Inertial (centrifugal) force in toroidal field acts like an effective gravity



Fast parallel dynamics + helical field lines provides stability  $\rightarrow$  gradient must surpass a threshold for instability
#### DIII-D Shot 121717

### GYRO Simulation Cray XIE, 256 MSPs

### Onset of turbulence reduces the achieved temperature that would have been present due only to diffusion



Temperature gradient (T<sub>hot</sub> - T<sub>cold</sub>)

## Onset of turbulence reduces the achieved temperature that would have been present due only to diffusion



Temperature gradient (T<sub>hot</sub> - T<sub>cold</sub>) Analogous to convective transport when heating a fluid from below ... boiling water (before the boiling)



Rayleigh, Benard instability (early 1900's)

## Numerous elements of US MFE research are actively addressing optimizing performance

$$Q \sim n T \tau_E \sim \beta \cdot B^2 \cdot \tau_E$$

- Refining operational scenarios to optimize beta, tau & steady-state (100% non-inductive) *simultaneously* DIII-D (General Atomics, San Diego)
- Pursuing very high beta (~30-40%) at low aspect ratio

   NSTX-U (PPPL) → low-A potentially useful as a compact volume neutron source for nuclear material qualification
- Pursing stronger magnetic fields: high temperature superconductors (HTS) enable larger B<sub>crit</sub>, J<sub>crit</sub>, T<sub>crit</sub>
  - MIT & Commonwealth Fusion Systems (CFS one of many privately funded fusion ventures)

#### Challenges (& career opportunities) to help enable a future with fusion energy

- Steady state operation with good confinement → plasma physics
- Handling intense heat fluxes at plasma-material boundary → plasma, chemistry & materials science
- Managing materials in a neutron environment & tritium breeding → <u>nuclear engineering</u>
- Better electromagnets → <u>superconductor R&D</u>
- Never-ending need for diagnostic development (<u>spectroscopy</u>, ...), data analysis (<u>big data analysis</u>, <u>machine learning</u>, ...), and just all around <u>good scientists &</u> <u>engineers!</u>

## Numerous tokamak, stellarator & other experiments are globally attacking these research topics

Lot's of fun places to visit / work (5)







#### **DOE-funded undergraduate internships**

- Summer Undergraduate Laboratory Internship (SULI): <u>https://science.energy.gov/wdts/suli/</u>
- Community College Internship (CCI): <u>https://science.energy.gov/wdts/cci/</u>
- Graduate Summer School at PPPL for grad students from non-fusion programs (<u>gss.pppl.gov</u>)

### Summary

- Nuclear fusion offers a promising solution for clean energy, especially for growing global energy demands
- The progress in magnetic fusion energy research has been immense – ITER is the next evolution on-thehorizon!
- Numerous US and international institutions are focused on solving remaining challenges to bring fusion energy to the market – always looking for new, young talent!





www.pppl.gov science.energy.gov/wdts/suli/ science.energy.gov/wdts/cci/



### **BACKUP SLIDES**

## Progress in achieved fusion performance outpaces (outpaced) Moore's Law



#### Comparing fusion-relevant performance among different magnetic energy confinement configurations



Z. Hartwig (MIT)

#### Princeton Plasma Physics Laboratory (PPPL) Plainsboro, New Jersey



# At PPPL, we try to understand many aspects of plasmas

For example...

#### Experiments to study astrophysical "reconnection" (solar flares)

**MRX** 

**Magnetic Reconnection Experiment** 

 Laboratory experiment to mimic interaction of solar flare impinging on earths magnetic field (relevant to telecommunications)

Onda d'urto Sole Vento solare Magnetopausa 🔩

## Experiments to study plasma thrusters for satellites and deep space exploration

The success of NASA's Dawn mission to orbit two asteroids depended on plasma thrusters



HTX Hall Thruster Experiment



### **Additional plasma research at PPPL**

- Astrophysics
- Plasma thrusters
- Basic plasma physics
- Nanotechnology
- Plasma-surface chemistry interaction
- Developing medical isotopes
- Plasma theory and simulation

#### Plasmas for nuclear fusion energy research

### Characteristics length & time scales that govern instabilities span orders of magnitude



Magnetic Fusion Energy Sciences